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Signal Analysis and Parametrization for EMP Data

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Presented are the theoretical background, program description, and computer listings that describe the techniques to reduce and analyze digital time-series data obtained in electromagnetic-pulse testing of military communications systems. The topics treated include data preprocessing, digital filtering, fast Fourier transform, fast Walsh transform, refined spectral densities, autocor-

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relation functions, waveform representation by parametrized functions, and nonlinear, least squares techniques. Several examples	
of the data are analyzed and discussed.	

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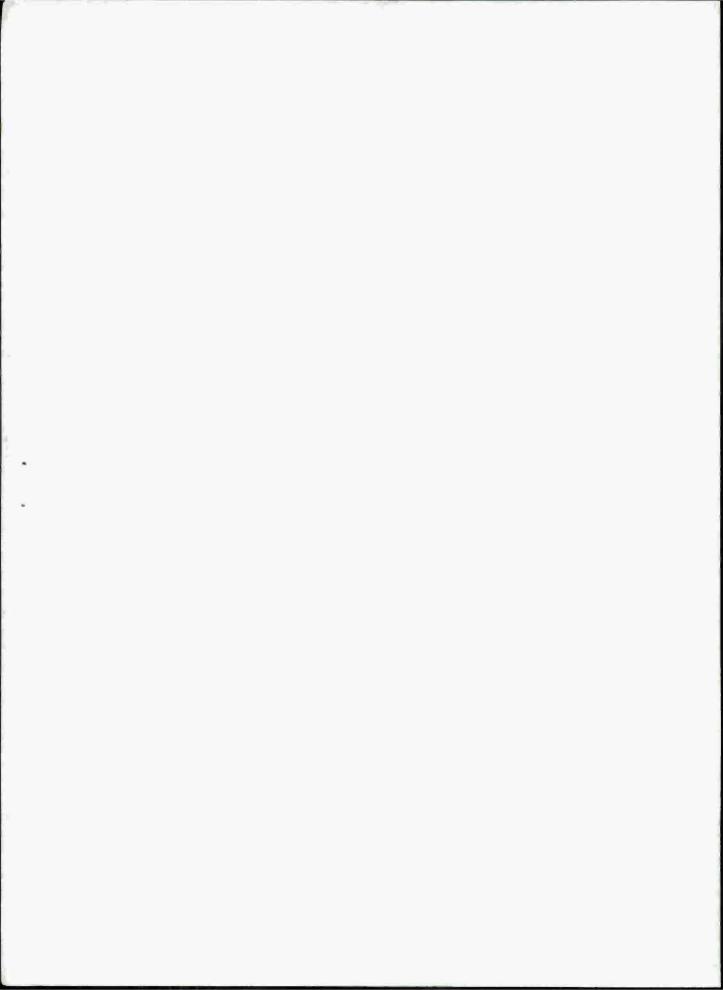
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#### 1. INTRODUCTION

This report presents the theoretical background and computer programs for the analysis techniques that were used in evaluating data obtained in the electromagnetic pulse (EMP) testing of military communication and weapons systems under the PREMPT program. 1

The data presented are initially obtained as a voltage-versus-time trace photographed on Polaroid film. This trace is then digitized and a time series of digital values is produced. The data are then processed in a digital computer. The various techniques employed in reducing and transforming the data are grouped under the generic title "signal analysis." Section 2 of this report gives a detailed description of all the algorithms; also, it contains complete instructions on how to use the signal-analysis program.

Another technique used in the data-reduction process is to represent the EMP waveform by a set of parametrized functions. This technique involves a least-squares fitting procedure which is discussed in section 3. Also, section 3 contains complete instructions on the use of a least-squares fitting program.

This report is not intended to be exhaustive on the subject of signal analysis but rather to present to the EMP community a basic software package that will: (1) accomplish most of the data reduction for EMP work and (2) be easily modified to include any additional techniques.

#### 2. SIGNAL ANALYSIS

This section presents the theoretical background and computer implementation of a number of techniques for reducing and transforming digital time series produced in EMP tests under the PREMPT program. A complete program listing annotated with comments is given in this section. Several versions of this program have been implemented on both the IBM 370-195 and CDC 6500 computers. All of the programming was done in FORTRAN. Some of the subroutines have been coded in assembly language for the NOVA minicomputer but are not reported here.<sup>2</sup>

<sup>&</sup>lt;sup>1</sup>The PREMPT program is a joint NMCSSC/DNA effort to determine the response of DCS to electromagnetic pulses generated by a high-altitude nuclear burst.

<sup>&</sup>lt;sup>2</sup>Further details are presented in "The Interactive Digitization and Editing System (IDES)" by Dr. Thomas A. Tumolillo, USA Harry Diamond Laboratories, Adelphi, MD 20783, (Aug 1973).

### 2.1 Data Preprocessing

Under the generic title "Data Preprocessing" is included the many minute details that are necessary to prepare the raw input data as obtained from a digitization of the waveform, so that it is suitable for transformation to the frequency domain.

The following sections discuss the method of referencing the trace to the scope graticule, scaling the data, time ordering, bit reversal, and a few of the simpler interpolation schemes. The software for two of the simpler interpolators, the linear Lagrange, and the linear least squares are presented. Higher order interpolators have been used occasionally in the PREMPT program, but are not included here because the simpler methods usually work. Simililarly, no techniques for time tying of digital records are given. Only the software are presented for the most commonly used grid and tablet referencing schemes, even though some of the more complicated procedures are discussed.

#### 2.1.1 Referencing the Trace to Graticule and Digitization Tablet

One important function of the data reduction is the determination of the rotation angle of the scope graticule with respect to the digitization tablet, the zero point (origin) of the graticule, and the scale factors for the X and Y axes in the graticule (or grid) coordinate system. There are many possible schemes that can be used to determine these factors; the most general method will be discussed in this section. The software exists for the general procedure as well as for the simpler specific case implemented in this signal analysis package.

For simplicity, but with no loss of generality, let the grid points be symmetrical about the origin of the grid coordinate system (fig. 1); call this origin ( $X_0,Y_0$ ). In the grid coordinate system the measure grid points are designated by the arrays XG(I) and YG(I). In the tablet coordinate system the measured grid points are designated by the arrays XT(I) and YT(I). The coordinate systems are shown in figure 1.

The transformation equation between the two systems is

$$\begin{bmatrix} X_{T} \\ Y_{T} \end{bmatrix} = \begin{cases} \cos \theta & -\sin \theta \\ \sin \theta & \cos \theta \end{cases} \begin{bmatrix} X_{G} \\ Y_{G} \end{bmatrix} + \begin{bmatrix} X_{O} \\ Y_{O} \end{bmatrix}$$
(1)

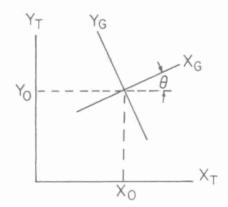


Figure 1. Diagram of coordinate systems.

Applying this transformation to the measured grid points, we have

$$\begin{bmatrix} \sum_{i=1}^{N} XT(i) \\ \sum_{i=1}^{N} YT(i) \end{bmatrix} = \begin{bmatrix} \cos\theta & -\sin\theta \\ \sin\theta & \cos\theta \end{bmatrix} \begin{bmatrix} \sum_{i=1}^{N} XG(i) \\ \sum_{i=1}^{N} YG(i) \end{bmatrix} + \begin{bmatrix} \sum_{i=1}^{N} X_{O} \\ \sum_{i=1}^{N} YG(i) \end{bmatrix} . \tag{2}$$

The points are symmetrical in the grid system; thus, we have

$$\sum_{I=1}^{N} xg(I) = \sum_{I=1}^{N} yg(I) = 0$$

Therefore,  $(X_0, Y_0)$  is given by

$$\begin{bmatrix} x_{O} \\ y_{O} \end{bmatrix} = \frac{1}{N} \begin{bmatrix} \sum_{I=1}^{N} XT(I) \\ \sum_{I=1}^{N} YT(I) \end{bmatrix}$$
(3)

It is convenient to shift the origin of the tablet coordinate system to the point  $(X_O, Y_C)$ ; thus,

$$\begin{bmatrix} \mathbf{X}_{\mathbf{T}}^{\prime} \\ \mathbf{Y}_{\mathbf{T}}^{\prime} \end{bmatrix} = \begin{bmatrix} \mathbf{X}_{\mathbf{T}} - \mathbf{Y}_{\mathbf{O}} \\ \mathbf{Y}_{\mathbf{T}} - \mathbf{Y}_{\mathbf{O}} \end{bmatrix} = \begin{pmatrix} \cos\theta & -\sin\theta \\ \sin\theta & \cos\theta \end{pmatrix} \begin{bmatrix} \mathbf{X}_{\mathbf{G}} \\ \mathbf{Y}_{\mathbf{G}} \end{bmatrix}$$

$$(4)$$

For a specific set of points in the grid system, the above transformation can be linearized. Take the line defined by  $Y_G = 0$ ; then we have

$$X_{T} = \cos\theta X_{G}, Y_{G} = 0 ..$$
 (5)

Let  $\alpha=\cos\theta$ , then  $X_T'=\alpha X_G$ . The best value of  $\theta$  can be estimated from the measured points along the grid X axis by using a least-squares technique. Let

$$\chi^{2} = \sum_{\mathbf{I}} [X\mathbf{T}(\mathbf{I}) - \alpha XG(\mathbf{I})]^{2}; \qquad (6)$$

minimizing  $\chi^2$  with respect to  $\alpha$ , we have

$$\frac{\partial(X^2)}{\partial \alpha} = 0 \Rightarrow \alpha = \cos\theta = \frac{\sum_{\mathbf{I}} XT(\mathbf{I}) * XG(\mathbf{I})}{\sum_{\mathbf{I}} XG(\mathbf{I}) * XG(\mathbf{I})}$$
(7)

The  $\sum_{i=1}^{\infty}$  means that we only sum over those points on the grid X axis.

We could now suitably define other straight lines on subsets of the measured grid points and get further estimates of  $\theta$ . It is more convenient to use a nonlinear least-squares technique, and extract the best value of  $\theta$ , by using all the points at once.

Let 
$$\vec{Z}(I) = \begin{bmatrix} XT(I) \\ YT(I) \end{bmatrix}$$
, (8)

$$\stackrel{\rightarrow}{\phi}(\theta, \mathbf{I}) = \begin{cases}
\cos \theta & -\sin \theta \\
\sin \theta & \cos \theta
\end{cases} \begin{bmatrix}
XG(\mathbf{I}) \\
YG(\mathbf{I})
\end{bmatrix} ,$$
(9)

and

$$\chi^{2} = \sum_{I=1}^{N} |\vec{Z}(I) - \vec{\phi}(\theta, I)|^{2}.$$
 (10)

We minimize  $\chi^2$  with respect to  $\theta$  by requiring  $\partial(\chi^2)/\partial\theta=0$ ; thus, we have

$$\frac{\partial(\chi^{2})}{\partial\theta} = \sum_{I=1}^{N} \left[ \frac{\partial\vec{\phi}^{\dagger}(\theta, I)}{\partial\theta} \cdot (\vec{Z}(I) - \vec{\phi}(\theta, I)) + (\vec{Z}(I) - \vec{\phi}^{\dagger}(\theta, I)) + (\vec{Z}(I) - \vec{\phi}^{\dagger}(\theta, I)) \right] = 0.$$
(11)

Here,  $\phi^{\dagger}$  is the adjoint of  $\Phi$ .

Assume now that we have defined an iterative process for evaluating  $\theta$  and  $\vec{\phi}$  , at the k iteration, we assume that  $\vec{\phi}$  is given by

$$\vec{\Phi}(\theta, \mathbf{I}) \sim = \vec{\Phi}(\theta^{\mathbf{k}}, \mathbf{I}) + \frac{\partial \vec{\Phi}(\theta^{\mathbf{k}}, \mathbf{I})}{\partial \theta} \cdot \Delta \theta^{\mathbf{k}}$$

$$\Delta \theta^{\mathbf{k}} = \theta^{\mathbf{k}+1} - \theta^{\mathbf{k}}$$
(12)

By substituting equation (12) into equation (11), a recursion relation is obtained for  $\theta$ . It can be shown after a slight algebraic manipulation that

$$\theta^{k+1} = \theta^k - C_1 \sin \theta^k + C_2 \cos \theta^k , \qquad (13)$$

where

$$C_{1} = \frac{\sum_{I=1}^{N} (XG(I) * XT(I) + YG(I) * YT(I))}{\sum_{I=1}^{N} (XG(I) * XG(I) + YG(I) * YG(I))},$$
(14)

$$C_{2} = \frac{\sum_{I=1}^{N} (XG(I)*YT(I) - YG(I)*XT(I))}{\sum_{I=1}^{N} (XG(I)*XG(I) + YG(I)*YG(I))},$$
(15)

The recursion relation generally converges quickly, as long as the initial estimate for  $\theta$  is close to the true value. We used for  $\theta^1$  the value determined in equation (7). The program considers that the iterative scheme has converged as long as  $\Delta\theta^k < 10^{-3}$  and  $\chi^2^{(k+1)} < \chi^2^{(k)}$ .

The scale factors for the X and Y axes can be easily calculated. For illustrative purposes, suppose we measured 28 grid points as shown in figure 2.

Initially, the arrays XG and YG are defined in a DATA statement so that the points [XG(I), YG(I)] correspond to a grid so that

$$XG(I) \in \{-1, -.8, -.6, ..., +.8, +1.\}$$

$$YG(I) \in \{-1, -.75, ..., +.75, +1.\}$$

so that for example [XG(27), YG(27)] = (+.4, -.5). In the program, one of the initial redefinitions of the arrays we make is

$$XG(I) \rightarrow XS*XG(I)$$
  
 $YG(I) \rightarrow YS*YG(I)$ 

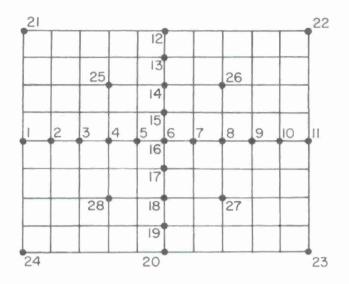


Figure 2. Measured grid points.

Scale factors XS and YS must be determined (fig. 3),

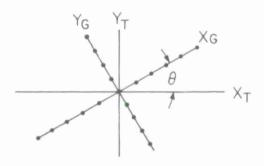


Figure 3. Rotation between G and T systems.

For the points on the  $X_{\widehat{G}}$  axis (I = 1,..., II), we have

$$[XS*XG(I)]^2 = XT(I)*XT(I) + YT(I)*YT(I)$$

The average value for XS is

$$XS = \underbrace{\sum_{I=1}^{11} (XT(I) * XT(I) + YT(I) * YT(I))}_{I=1} .$$
 (16)

Similarly, from the points on the  $Y_G$  axis (I = 12, 20), we have

$$YS = \frac{\sum_{I=12}^{20} (XT(I)*XT(I) + YT(I)*YT(I))}{\sum_{I=12}^{20} YG(I)*YG(I)}$$
(17)

A complete grid measurement is rarely carried out as a routine operation in the reduction of EMP data. Occasionally, it will be done to test the linearity of the system. The most common method is to measure two points on each axis and calculate all quantities from these numbers. The following equations were implemented in the signal analysis package. The routine calculates the grid rotation angle, the scale factor, and the origin point from two measured points on both the X and Y grid axes. Initially, the program reads in the coordinates of the two points measured on the X axis—X1 and X2—and on the Y axis—Y1 and Y2. Then the four grid points are read in the same order that the coordinate locations were read in. The four grid points are stored in the arrays XT(I), YT(I), I = 1,4. The origin of the grid (X0, Y0) is given by

$$XO = |X2*XT(1)-X1*XT(2)| / |X1-X2| ,$$

$$YO = |Y2*YT(3)-Y1*YT(4)| / |Y1-Y2| ,$$
(18)

The scale factors XS and YS are given by

$$XS = \sqrt{(XT(2)-XT(1))**2 + (YT(2)-YT(1))**2} / (X2-X1),$$

$$YS = \sqrt{(XT(4)-XT(3))**2 + (YT(4)-YT(3))**2} / (Y2-Y1),$$
(19)

The rotation angle is determined by

$$\tan(\theta) = 0.5* \left( \frac{\text{YT}(2) - \text{YT}(1)}{\text{XT}(2) - \text{XT}(1)} - \frac{\text{XT}(4) - \text{XT}(3)}{\text{YT}(4) - \text{YT}(3)} \right) ,$$

$$CT = \cos(\theta) = 1./\sqrt{1 + \tan(\theta) * *2} ,$$

$$ST = \sin(\theta) = \cos(\theta) * \tan(\theta) . \qquad (20)$$

Then the subroutine reads scale factors T (nanosecond/grid division) and V (volts/grid division), then recomputes the scale factors as

$$T+T/XS$$
,  $V+V/XS$ , (21)

which have units nanoseconds/tablet counts and volts/tablet counts. Then the (x,y) coordinates of the trace are transformed as,

$$X(I) \leftarrow (CT^*(X(I)-XO) + ST^*(Y(I)-YO)-XZ)^*T$$
,  
 $Y(I) \leftarrow (-ST^*(X(I)-XO) + CT^*(Y(I)-YO)-YZ)^*V$ , (22)

where (XZ, YZ) are the rotated coordinates of the "zero point of the trace," that is, the point on the trace at which the signal starts,

$$XZ \leftarrow CT^*(XZ-XO) + ST^*(YZ-YO)$$

$$YZ \leftarrow -ST^*(XZ-XO) + CT^*(YZ-YO)$$
(23)

#### 2.1.2 Time Ordering, Bit Reversal, and Interpolation Schemes

Time ordering of the array is a necessary procedure in order to remove errors introduced in the digitization process. Occasionally, there will be errors in the grid measurements that cause portions of the trace to fold back in the time sense after it is rotated. Similarly, inaccurate movement of the digitization operator's hand may also cause a few points to be folded back in the time sense. In some cases the digitization hardware will allow consecutive digital points to have the same time value. If these measurement ambiguities are not removed, they will cause considerable error in the high-frequency part of the transforms. This correction of the data is handled in subroutine CST OUT. This routine casts out those points in the array that are folded

back--that is, if XF is the input time array and XF(K) < XF(I) for K = I + 1, I + 2,... then SF(K) is deleted from the array. Similarly, if XF(K) = XF(I) for some set of K then the routine averages the amplitude YF(K) to create a single value at that value XF(I).

Bit reversal of the array refers to a specific reordering of the elements of a digital time series prior to its entering the FFT subroutine. It is done so that after transformation the frequency domain arrays are in ascending order of the frequency value. The term bit reversal arises from representing the index of an array I in base-two notation. For example, suppose we have the 65th element of a 1024 element array, then  $65_{10} = 0001000001_2$ . The reverse of the number is  $1000001000_2 = 520_{10}$ . To bit reverse, we swap the elements 65 and 520 of the original array.

In subroutine LNYQ, the Lagrangian methods for interpolation are used. A brief description of Lagrange interpolation is given here.

It is generally assumed that the function, f, interpolated here behaves like a polynomial; thus, in order to calculate f approximately at a point x, we find a polynomial approximation g for f good in the neighborhood of x. Lagrange showed that there is a unique polynomial of degree n having n + 1 values f at n + 1 distinct points x , i = 0,...n. That polynomial is  $g_n$ ,

$$g_{n}(x) = \sum_{i=0}^{n} f(x_{i}) \prod_{j=0}^{n} (x - x_{j})/(x_{i} - x_{j}) .$$

$$j=0$$

$$j\neq i$$

$$(24)$$

For the software presented in this package, n is restricted to the value 1. Thus,

$$g_{1}(x) = \frac{f(x_{1}) - f(fx_{0})}{x_{1} - x_{0}} x + \frac{x_{1}f(x_{0}) - x_{0}f(x_{1})}{x_{1} - x_{0}}$$

$$= C1 \cdot X + C2 \qquad (25)$$

In the program the function f is called YF(I), x, is replaced by the time array XF(I), and the interpolated values are put into the real part of the complex array YNYQA(K)

REAL [YNYQA(K)] =  $Cl \cdot X + C2$ ,

$$C1 = \frac{YF(I) - YF(I-1)}{XF(I) - XF(I-1)}, \qquad C2 = \frac{XF(I) * YF(I-1) - XF(I-1) * YF(1)}{XF(I) - XF(I-1)}. \tag{26}$$

In subroutine NYQST, a linear polynomial is fitted to a set of points in the arrays YF(I), XF(I), I=LB,...,LT. After determining the polynomial, the program evaluates it at the predetermined interpolation point.

The theory behind the least squares, polynomial fitting programs is straightforward. We need to minimize  $\chi^2$  with respect to the C(J) where

$$\chi^{2} = \sum_{I=LB}^{I,T} [YF(I) - \sum_{J=0}^{M} C(J) * XF(I) * * J]^{2}.$$
 (27)

Setting  $\frac{\partial(\chi^2)}{\partial \mathcal{C}(J)} = 0$  J = 0,..., M yields

$$\vec{B} = A \cdot \vec{C} \quad , \tag{28}$$

where the (K,L)<sup>th</sup> element of the matrix A is

$$A(K,L) = \sum_{I=LB}^{LT} XF(I) **(K + L - 2), K, L = 1,...,M,$$
 (29)

the  $K^{th}$  element of the vector  $\vec{B}$  is

$$\vec{B}(K) = \sum_{I=LB}^{LT} YF(I) * (XF(I) * * (K - 1)) \quad K = 1,...,M,$$
(30)

and the K  $^{\text{th}}$  element of the vector  $\vec{C}$  is just the K  $^{\text{th}}$  polynomial coefficient. Upon inversion we find the solution for  $\vec{C}$ 

$$\vec{C} = A^{-1} \cdot \vec{B} \tag{31}$$

Both of the interpolator subroutines take due account of the end points of the arrays and minimize the number of calculations when more than one interpolated point falls between the same time values.

### 2.2 Digital Filtering

This section briefly reviews the theory of digital filters. Frequently in EMP work, the signal is contaminated by high-frequency noise arising both from the nature of the measurements and the digitization process which reduces the continuous signal to a digital record. This high-frequency component can generally be eliminated by passing the digital record through a low-pass digital filter. Another useful application of low-pass filters is in reducing the number of digital values needed to accurately calculate the Fourier transform of a waveform at low frequencies. For example, if there is a signal with frequency content up to 250 MHz, the Nyquist criterion is satisfied by sampling the signal every 2.0 nsec. If the signal has a 2-usec duration, then 1000 numbers must be stored for the Fourier transform routine. However, if most of the significant frequency content is contained in a frequency band up to say 50 MHz, then after filtering, only 200 numbers must be stored to adequately represent the signal and obtain the Fourier transform without worrying about foldover effects. Bandpass and high-pass digital filters have application in EMP work when one is interested in studying only a certain region of the frequency spectrum so that correlations between equipment upset and damage and the induced signal can be determined.

#### 2.2.1 Theory of Digital Filtering and Z Transforms

In EMP work, there is generally (after interpolation or as a result of the digitization process) a sequence of numbers u(k), k=0,...N that must pass through a linear discrete system (generally a difference equation) in order to limit, in some manner, the frequency content of the signal. The output of the linear system is denoted here by y(k), k=0,...,N. By a linear discrete system is meant a system in which the output y(k) is expressed as a linear combination of inputs and past outputs; thus,

$$y(k) + \sum_{j=1}^{n} a(j)y(k-j) = \sum_{\ell=0}^{m} b(\ell)u(k-\ell)$$
, (32)

and y(k), u(k) = 0, k < 0.

The Z transform is used to simplify the analysis and synthesis of the digital filter represented by equation (32)—for example, generate the set of constants a(j) and  $b(\ell)$ .

The Z transform of a sequence of numbers f(k), k=0,...N f(k)=0, k<0 is defined by

$$Z[f(k)] = F(z) = \sum_{k=0}^{\infty} f(k)z^{-k}$$
, (33)

where z is an arbitrary complex number.

The Z transform of the input-output signals are related to one another by

$$Y(z) = H(z) U(z) , \qquad (34)$$

where U(z) and Y(z) denote the Z transforms of the input and output signals respectively, and H(z) is the system transfer function, and is given by

$$H(z) = \sum_{k=0}^{m} b(k) z^{k} / \left(1 + \sum_{j=1}^{n} a(j) z^{-j}\right)$$
 (35)

Proof: Multiply each side of equation (32) by  $z^{-k}$  and sum over all k.

$$\sum_{k=0}^{\infty} y(k)z^{-k} + \sum_{j=1}^{\infty} a(j) \sum_{k=0}^{\infty} y(k-j)z^{-k} = \sum_{\ell=0}^{m} b(\ell) \sum_{k=0}^{\infty} u(k-\ell)z^{-k} . (36)$$

Using the properties of the Z transform, this can be rewritten as

$$Y(z) + \sum_{j=1}^{n} a(j)z^{-j} Y(z) = \sum_{\ell=0}^{m} b(\ell)z^{-\ell} U(z);$$
 (37)

or, as

$$Y(z) \left[1 + \sum_{j=1}^{n} a(j)z^{-j}\right] = \left[\sum_{\ell=0}^{m} b(\ell)z^{-\ell}\right]U(z) , \qquad (38)$$

Thus, equation (35) follows and may be rewritten as

$$H(z) = b \prod_{i=1}^{r} (z-z_i) / \prod_{i=1}^{n} (z-p_i) .$$
(39)

Here, z and p are called the zeros and poles of the system transfer function H(z).

One of the most important properties of the transfer function is the fact that the location of the zeros and poles has an enormous effect on how the system transmits different types of inputs. Thus, a system can be synthesized that will pass some inputs and reject others by a proper selection of the zeros and poles.

Assume the application of a sinusoidal input  $u(k) = \sin(kwT)$  k=0, 1, 2,... to our systems. Then the resultant steady-state response, y(k), is given by

$$y(k) = \left| H(e^{iwT}) \right| \sin(kwT+\theta) . \tag{40}$$

Here, the sinusoidal response of the system is obtained by evaluating the system transfer function H(z) at  $z=e^{iwT}$ , where w is the radian frequency of the input sinusoid and T is the underlying sampling period.  $\theta$  is the phase angle of  $H\left(e^{iwT}\right)$ —that is,  $H\left(e^{iwT}\right) = H\left(e^{iwT}\right) = H\left($ 

$$U(z) = \frac{z s i w T}{\left(z - e^{i w T}\right) \left(z - e^{-i w T}\right)}$$

$$(41)$$

The Z transform of the system response is given by,

$$Y(z) = \frac{H(z)z\sin wT}{(z-e^{iwT})(z-e^{-iwT})} . \tag{42}$$

Since only stable systems are considered, all the poles of H(z) must be inside the unit circle. Therefore, none of the poles of H(z) is at  $e^{iw^T}$  or  $e^{-iw^T}$ . A partial fraction expansion of equation (42) yields

$$Y(z) = \frac{az}{(z-e^{iwT})} + \frac{bz}{(z-e^{-iwT})} + \text{terms due to poles of } H(z), \tag{43}$$

where

$$a = H(e^{iwT})/2i$$
 and  $b = -H(e^{-iwT})/2i$ .

Noting that  $H(e^{iwT}) = H(e^{-iwT}) *$  and writing  $H(e^{iwT}) = Me^{i\theta}$ , we find that

$$Y(z) = \frac{M}{2} \left[ \frac{ze^{i\theta}}{(z-e^{iwT})} - \frac{-ze^{-i\theta}}{(z-e^{-iwT})} \right] + \text{ terms due to poles of } H(z) . \tag{44}$$

Taking the inverse Z transform of equation (44) we obtain

$$y(k) = \frac{M}{2i} \left[ e^{i(kwT+\theta)} - e^{-i(kwT+\theta)} \right] + \text{transient response}$$

$$\text{generated by poles}$$
of  $H(z)$ :

or,

$$y(k) = M \sin(kwT+\theta) + y_{transient}(k)$$
 (46)

In the steady-state y transient (k) - 0 as k becomes large. Thus

$$y(k) = \left| H(e^{iwT}) \right| \sin(kwT + \theta) . \tag{47}$$

The quantity  $M = \left| H\left(e^{\mathrm{i} \, W \, T}\right) \right|$  is called the system gain factor. To filter out a given sinusoid, pick a system so that  $M \sim 0$ ; or, to amplify a given sinusoid, pick M > 1.

As has been observed in equation (39), the transfer function is a ratio of polynomials in the variable z. Therefore, the system gain factor that is equal to the magnitude of the transfer function evaluated at  $z=\exp(iwT)$  may always be expressed as a ratio of polynomials in the variables  $\cos(wT)$  and  $\sin(wT)$ . Thus, different filters can be synthesized by investigating ratios of trigonometric functions. For example, a low-pass filter with half-power point w, has the following squared gain factor,

$$|H(e^{iwT})|^2 = \frac{1}{1 + \frac{\tan^{2n}(wT/2)}{\tan^{2n}(w_1T/2)}}$$
 (48)

By a considerable amount of algebraic manipulation equation (48) can be written as

$$|H(z)|^{2} = \frac{\tan^{2n}(w_{1}T/2)(1+z)^{2n}}{\left[\tan^{2n}(w_{1}T/2) + (-1)^{n}\right]\left[(z-p_{1})(z-p_{2}) \cdot \cdot \cdot (z-p_{2n})\right]},$$
(49)

where  $z = \exp(iwT)$ , and the 2n poles  $p_i$  are given by

$$p_{i} = \frac{1 - \tan^{2}(w_{1}T/2) + \sqrt{-1} 2\tan(w_{1}T/2)\sin\theta_{i}}{1 - 2\tan(w_{1}T/2)\cos\theta_{i} + \tan^{2}(w_{1}T/2) r},$$
(50)

where,

 $\theta_{i} = (i-1) \pi/2$ , n odd

=  $(2i-1) \pi/2$ , n even .

It can be shown that of the 2n poles,  $p_i$ , exactly n lie inside the unit circle and n outside. Let  $p_1$ ,  $p_2$ ,..., $p_n$  denote the n poles inside the unit circle. The transfer function that has the desired squared-gain factor is given by

$$H(z) = \frac{b(1+z)^n}{\left(z-p_1\right)\left(z-p_2\right)\cdots\left(z-p_n\right)},$$
(51)

where b is chosen so that the steady-state, unit-step response has magnitude one, H(1)=1; thus,

$$b = \frac{(1-p_1)(1-p_2)...(1-p_n)}{2^n} . (52)$$

The remaining poles  $p_{n+1}$ ,  $p_{n+2}$ ,..., $p_{2n}$  associated with  $|H(z)|^2$ can be shown to arise from the process of determining the squared-gain factor.

From the foregoing, a well-defined procedure exists for synthesizing a low-pass filter. It can be summarized in the following steps:

- (a) Determine the half-power point w1,
- (b) Determine the value n--using equation (17)--by specifying the gain at frequency  $\mathbf{w}_{\text{p}}$ ,
- (c) Find the n roots p given by equation (49) which satisfy  $|p_{\rm j}|<1$ , and
- (d) Determine the difference equation which has the transfer function given by equations (51) and (52).

A squared-gain factor that corresponds to that of a high-pass filter is given by

$$\left| H\left(e^{i\mathbf{w}T}\right) \right| = \frac{1}{1 + \frac{\cot^{2n}(\mathbf{w}T/2)}{\cot^{2n}(\mathbf{w}_{2}T/2)}}$$
(53)

Here, the half-power point is denoted by  $w_2$ . Formulas analagous to equations (50) and (51) may be derived. Then, if the poles and zeros of a low-pass filter with half-power point  $(\pi/T-w_2)$  are rotated through  $\pi$  radians in the complex plane, the pole-zero pattern of a high-pass filter is obtained with half-power point  $w_2$  and the same gain-factor falloff outside its passband is obtained. Thus, if we have H(z) for the low-pass filter given by

$$H(z) = b \left(1+z\right)^{n} / \pi \left(z-p_{i}\right) , \qquad (54)$$

then the high-pass filter is given by H'(z)

H'(z) = b'(z-1)<sup>n</sup>/
$$\frac{n}{\pi}$$
(z+p<sub>i</sub>) . (55)

### 2.2.2 Examples of Filtering EMP Data

To illustrate the implementation of the algorithms described in section 1, a typical EMP waveform was selected from the vast amount of data collected at the Polk City AUTOVON EMP tests and processed. Figure 4 plots the digitized waveform after it has been digitized, time ordered, and interpolated at 1.63-nsec intervals using a Lagrange interpolator.

Figure 5 plots the power spectrum after passing the digital record through a fast Fourier transform routine. All the power is contained in two peaks at 13.2 and 24.0 MHz. Several filters were then synthesized and the digital record passed through filters before processing it through the fast Fourier transform routines. Table I lists the half-power points,  $\mathbf{w_1}$ , the gain at the higher frequency  $\mathbf{w_2}$  used to determine n, and the filter coefficients  $\mathbf{a(j)}$ ,  $\mathbf{j=1,...,n}$ ,  $\mathbf{b(j)}$ ,  $\mathbf{j=0...,n}$ , which were calculated for each synthesized filter. Figures 6 through 12 are plots of the power spectra obtained by using a fast Fourier transform routine on the filtered-time series. The general result is fairly evident from these plots—namely, that as the number of poles is increased and consequently the gain rolloff at the half-power point is increased, a sharper filtering is obtained.

Although it is not apparent from the plots of the power spectra presented, there is much greater definition of the peaks in the power spectrum relative to the noise background. The high-frequency noise was reduced by a factor of 100.

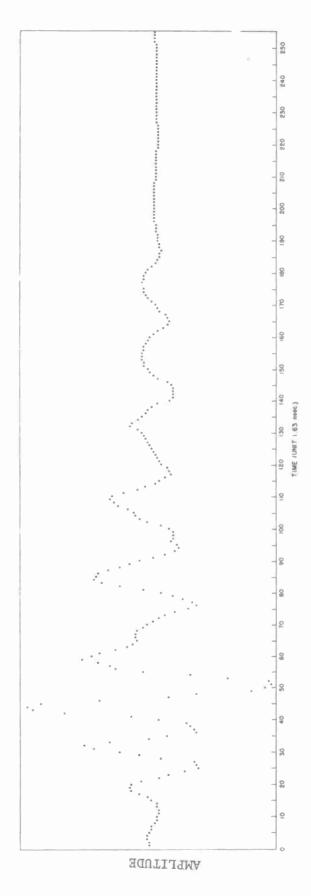


Figure 4. Plot of equispaced digital times series.

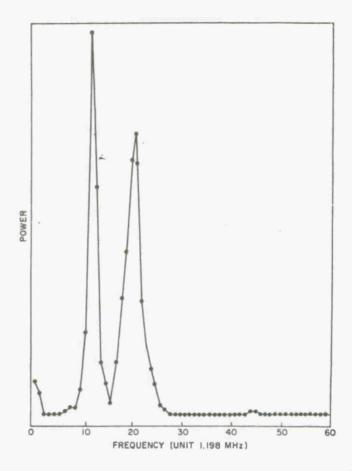


Figure 5. Plot of power spectrum obtained without filtering.

TABLE I. PARAMETERS CALCULATED FOR EACH SYNTHESIZED FILTER

Filter No.	Half-power point W <sub>1</sub> (MHz)	Gain W= W <sub>2</sub> > W <sub>1</sub>	w <sub>2</sub>	Sampling Interval T (µsec)	n	a(j),j=1,,n	b(j),j=1,n+1
1	20	0.4	25	1.63	1	-0.8135657E00	+0.9321713E00 +0.9321713E00
2	20 .	0.2	25	1.63	4	-0.3465070E 01 0.4533676E 01 -0.2652045E 01 0.5848047E 00	0.8540331E-04 0.3416131E-03 0.5124197E-03 0.8540331E-04
3	<b>20</b>	0.1	25	1.63	5	-0.4337327E 01 0.7563597E 01 -0.6625226E 01 0.2913780E 01 -0.5145599E 00	0.8235322E-05 0.4117661E-04 0.8235322E-04 0.8235322E-04 0.4117661E-04 0.8235322E-05
4	20	0.07	25	1.63	6	-0.5208673E 01 0.1135041E 02 -0.1324103E 02 0.8718896E 01 -0.3071778E 01 0.4523243E 00	0.7937438E-06 0.4762463E-05 0.1190616E-04 0.1587487E-04 0.1190616E-04 0.4762463E-05 0.7937438E-06
5	20	0.05	25	6.52	6	-0.2848268E 01 0.3845003E 01 -0.2959958E 01 0.1351626E 01 -0.3426920E 00 0.3745000E-01	0.1299385E-02 0.7796306E-02 0.1949077E-01 0.2598769E-01 0.1949077E-01 0.7796306E-02 0.1299385E-02
6	18	0.2	25	1.63	3	-0.2631725E 01 0.2328194E 01 0.6912119E 00	0.6571114E-03 0.1971334E-02 0.1971334E-02 0.6571114E-03
7	17.5	0.05	25	1.63	5	-0.4420094E 01 0.7844978E 01 -0.6986225E 01 0.3120767E 01 -0.5592871E 00	0.4385640E-05 0.2192819E-04 0.4385639E-04 0.4385639E-04 0.2192819E-04 0.4385640E-05

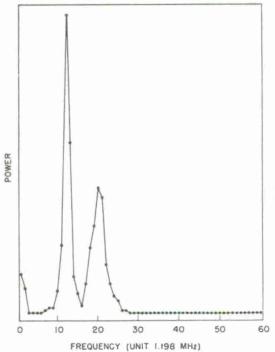
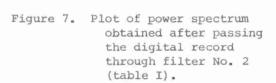
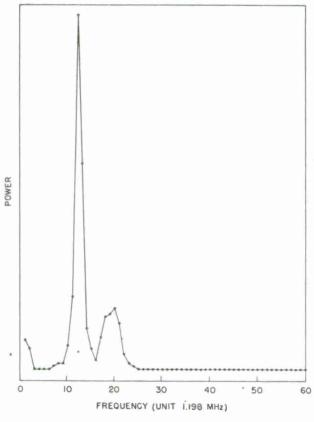


Figure 6. Plot of power spectrum obtained after passing the digital record through filter No. 1 (table I).





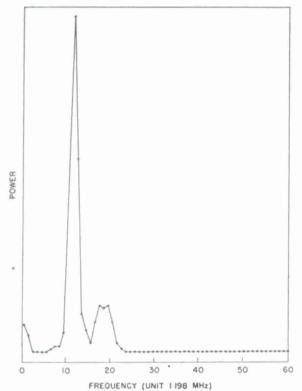
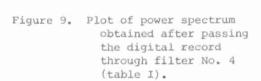
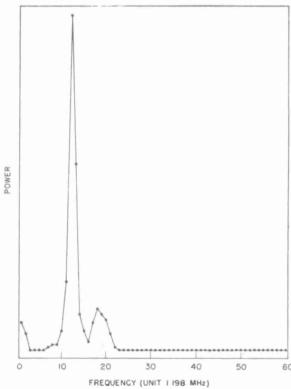


Figure 8. Plot of power spectrum obtained after passing the digital record through filter No. 3 (table I).





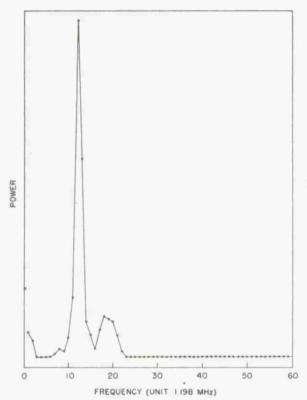
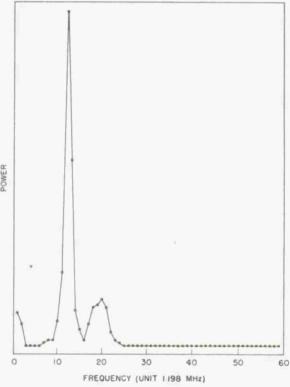


Figure 10. Plot of power spectrum obtains after passing the digital record through filter No. 5 (table I).

Figure 11. Plot of power spectrum obtained after passing the digital record through filter No. 6 (table I).



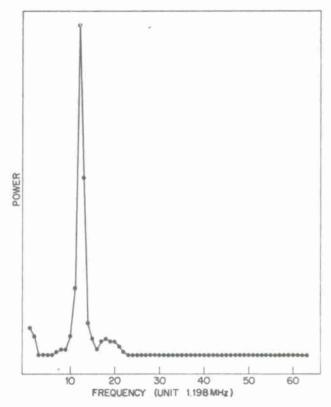


Figure 12. Plot of power spectrum obtained after passing the digital record through filter No. 7 (table I).

#### 2.3 Transform Techniques

Three different popular transform techniques are discussed below.

#### 2.3.1 Fast Fourier Transform

The fast Fourier transform is by far the method most preferred for generating the Fourier transform of a digital-time series. A short discussion of the method is given in this section.

 $F(\omega)$  of a function f(t) is defined as

$$F(\omega) = \int_{-\infty}^{\infty} f(t) e^{-i\omega t} dt . \qquad (56)$$

If f(t) is nonzero only over a finite time interval T, then it is a good approximation to write the Fourier transform as a discrete sum,

$$F(\omega) = \Delta t \sum_{k=0}^{N-1} f(k\Delta t) e^{-\omega k \Delta t}, \quad N\Delta t = T.$$
(57)

The inverse transform is

$$f(t) = \frac{1}{2\pi} \int_{-\infty}^{\infty} F(\omega) e^{i\omega t} d\omega.$$
 (58)

If  $F\left(\omega\right)$  has no (or a negligible) frequency content above  $\omega_{\mbox{max}}$  the inverse transform can be written as

$$f(t) = \frac{\Delta\omega}{2\pi} \sum_{r=0}^{N-1} F(r\Delta\omega) e^{itr\Delta\omega}, N\Delta\omega = 2\omega_{max} .$$
 (59)

For a band-limited signal, we know from the Nyquist criterion that the sampling interval  $\Delta t$  must be chosen so that  $\Delta t = 2\pi/2\omega_{\rm max}$ . Similarly, for a time-bounded signal, the frequency-sampling interval must be chosen so that  $\Delta \omega = 2\omega/T$ ,

$$\Delta t \Delta \omega \le \frac{T}{N} \frac{2\pi}{T} = \frac{2\pi}{N} \tag{60}$$

Thus,

$$F(r\Delta\omega) = \Delta t \sum_{k=0}^{N-1} f(k\Delta t)e^{-2\pi i r k/N}, r=0, 1,..., N-1.$$
 (61)

For notational convenience, we write  $F_r = F(r \Lambda \omega)$ ,  $f_k = f(k \Lambda t)$ ,  $W = e^{-2\pi i/N}$ , and drop the factor  $\Lambda t$ .

$$F_{r} = \sum_{k=0}^{N-1} f_{k} W^{rk} \qquad r = 0, ..., N-1.$$
 (62)

Divide the time series of points,  $f_k$ , into two functions  $g_k$  and  $h_k$ ,

$$g_k = f_{2k}$$
,  
 $h_k = f_{2k+1}$ ,  $k = 0, 1, 2...\frac{N}{2} - 1$  . (63)

The discrete Fourier transforms of  $\mathbf{g}_k$  and  $\mathbf{h}_k$  are  $\mathbf{G}_r$  and  $\mathbf{H}_r$  respectively.

$$G_{r} = \sum_{k=0}^{(N/2)-1} g_{k} e^{-2\pi i r k/(N/2)},$$

$$r = 0,1,...\frac{N}{2} - 1 \qquad (64)$$

$$H_{r} = \sum_{k=0}^{(N/2)-1} h_{k} e^{-2\pi i r k/(N/2)}.$$

Equations (61) and (62) may be rewritten as

$$F_{r} = \sum_{k=0}^{(N/2)-1} f_{2k} e^{-2\pi i r k/(N/2)} + e^{-2\pi i r/N} \sum_{k=0}^{(N/2)-1} f_{2k+1} e^{-2\pi i r k/(N/2)} . (65)$$

Using equations (63) and (64), we have

$$F_r = G_r + W H_r$$
  $0 \le r < N/2$   $F_{r+N/2} = G_r - W H_r$   $0 \le r < N/2$  . (66)

We may therefore compute the Fourier transform of a function sampled N times by evaluating two Fourier transforms of the function sampled N/2 times. The computations of  $G_r$  and  $H_r$  can be reduced to the evaluation of sequences of N/4 samples. If N =  $^2$ , n such reductions can be made by applying equations (63) and (66) first for N, then N/2, and finally for a two-point function. The Fourier transform of a one-point function

is just the sample itself. Beside the obvious savings in computer running time by using this successive reduction scheme, the transform can also be done in place—that is, in each stage of reduction the intermediate results are written over the original array. This in-place reduction requires a rearrangement of the original array called bit reversal (subroutine SRTFUR). A complete rearrangement and overwriting sequence for an eight-point sampled function is illustrated in figure 13.

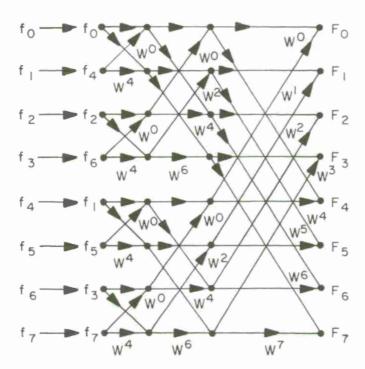


Figure 13. Rearrangement and overwriting sequence for an eight-point sampled function.

Each arrow in the diagram means that the term at the origin of the arrow must be added. A variable next to the arrow acts as a multiplier to the additive term. Thus, at the first overwriting, sequence  $\mathbf{f}_0$  is replaced by  $\mathbf{f}_0^{+W^0}*\mathbf{f}_4$  and  $\mathbf{f}_4$  is replaced by  $\mathbf{f}_0^{+W^0}*\mathbf{f}_4$ .

Subroutine FFT will compute the transform of any array with 2<sup>n</sup> elements as described above. The only restriction is that imposed by the finite memory size of the computer being used. Several examples of the power spectrum derived from the real and imaginary parts of the transform generated by FFT are shown in figures 5 through 12.

### 2.3.2 Fast Walsh Transform

The previous section transforms the digital-time series to the frequency domain by using a particular set of orthogonal functions—namely sines and cosines. Another set of orthogonal functions, which are used extensively in communications theory, are the Walsh functions that are used primarily to represent logic signals. Their most appealing feature is that the digital Walsh transform algorithm is about an order of magnitude faster than the Fourier transform algorithm.

The Walsh functions are wal( $k,\theta$ ), sal( $k,\theta$ ), and cal( $k,\theta$ ),

wal(
$$2k,\theta$$
) = cal( $k,\theta$ )  
wal( $2k-1,\theta$ ) = sal( $k,\theta$ )

They are defined on the time interval T,  $\theta$  is the normalized time  $\theta$ =t/T, and k is called the sequency. The sequency is equal to the average number of zero crossings of the function per unit time. The functions sal and cal are similar to the sine and cosine functions. The sequency of the Walsh functions plays a similar role as the frequency for the sinusoidal functions. One definition of the Walsh functions is through a difference equation,

$$wal(2k+p,\theta) = (-)^{\lfloor k/2 \rfloor + p} \left[ wal \left[ k, 2 \left( \theta, \frac{1}{4} \right) \right] + (-)^{k+p} wal \left[ k, 2 \left( \theta - \frac{1}{4} \right) \right] \right], \quad (67)$$

where

$$k = 0, 1, 2, \dots,$$

[k/2] is the largest integer less than or equal to k/2,

$$p = 0$$
 or 1, and

$$wal(0,\theta) = 1 \theta \le \frac{1}{2}$$

$$0 \theta > \frac{1}{2}$$
(68)

A few of the Walsh functions are shown in figure 14.

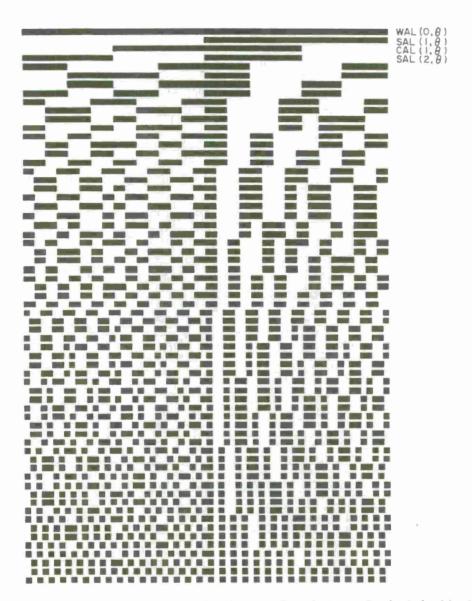


Figure 14. Examples of Walsh functions for interval  $-\frac{1}{2} < \theta < \frac{1}{2}$  (dark areas imply +1, light areas imply -1).

A signal f(t) may be expanded in a Walsh series

$$f(t) = a(0)wal(0,t) + \sum_{k=1}^{\infty} \left[ a_{c}(k)cal(k,t) + a_{s}(k)sal(k,t) \right],$$

$$a(0) = \int_{-T/2}^{T/2} f(t)dt,$$

$$a_{c}(k) = \int_{-T/2}^{T/2} f(t)cal(k,t) dt,$$

$$a_{s}(k) = \int_{-T/2}^{T/2} f(t)sal(k,t)dt.$$
(69)

Just as in a Fourier expansion, the sum of the squares of the expansion coefficients give the sequency energy spectrum. The Walsh power is

$$E(k) = a_C^2(k) + a_S^2(k)$$
 (70)

To evaluate the coefficients, a fast Walsh transform algorithm can be derived that is similar to the fast Fourier technique. The main difference is that the reduction cannot be done in place. The steps for an eight-point-sampled function is shown in figure 15. The arrows have the same meaning as in the overwriting sequence for the FFT. Thus, at the first overwrite,  $f_0$  is replaced by  $f_0+f_1$ .

For the EMP data of figure 4, the Walsh power was calculated and the results plotted in figure 16. The same figure plots the Walsh power after filtering the data through filter No. 7, table I.

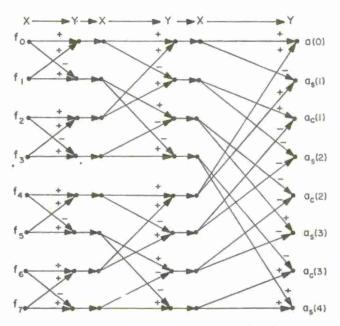


Figure 15. Overwriting sequence for Walsh transform.

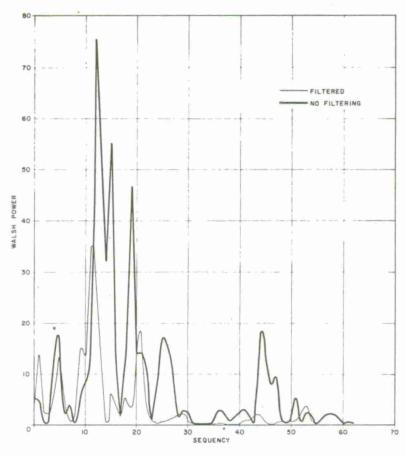


Figure 16. Walsh power for the time series of figure 4.

# 2.3.3 Refined Spectral Densities and Autocorrelation Function

In EMP work, there is a great need for implementing numerical algorithms for the processing and storage of digitized waveforms, which are the main output of all EMP tests. This subsection describes one algorithm to generate the power spectrum of a digital record from its autocorrelation function. The theory presented below is illustrated by several examples.

A procedure will now be defined for estimating the power spectrum of a uniformly spaced, discrete time series of finite length.

If  $C(\tau)$  is the autocovariance function for a time waveform X(t), then by definition

$$C(\tau) = \lim_{T \to \infty} \frac{1}{T} \int_{-T/2}^{T/2} X(t)X(t+\tau)dt . \qquad (71)$$

The power spectrum P(w) of the time waveform is then given by

$$P(w) = 2 \int_{0}^{\infty} \cos(w\tau) C(t)dt.$$
 (72)

For a uniformly spaced discrete-time series of finite length, denoted by  $X_0$ ,  $X_1$ ,...,  $X_n$  compute the mean lagged products,  $C_n$ , with lag interval  $\Delta \tau = h \Delta t$ , and  $\Delta t$  is the time interval between adjacent values of the time series.

$$C_r = \frac{1}{n-hr} \sum_{q=0}^{q=n-hr} X_q X_{qthr}, r=0, 1,..., m, m \le \frac{n}{h}$$
 (73)

Next, compute the "raw spectral density estimates" Vr.

$$V_{r} = \Delta \tau \cdot \left[ C_{o} + 2 \sum_{q=1}^{q=m-1} C_{q} \cos \frac{qr\pi}{m} + C_{m} \cos r\pi \right]$$
 (74)

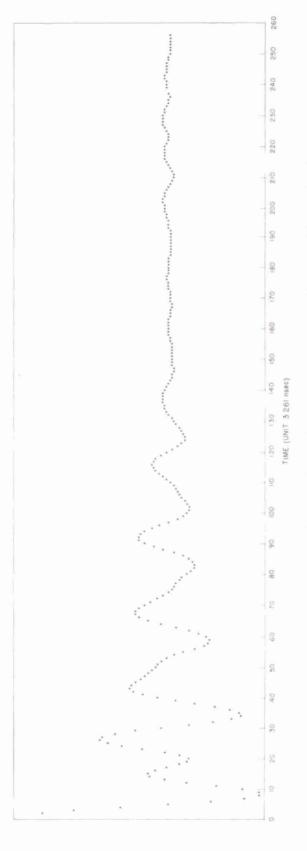
The frequency corresponding to r is  $r/2m\Delta\tau$ .

We next calculate the refined spectral density estimates according to

$$U_{r} = 0.23 V_{r-1} + 0.54 V_{r} + 0.23 V_{r+1} . (75)$$

To illustrate the implementation of the algorithms given by equations (73), (74), and (75), refer once again to figures 4 and 5. Figure 4 plots the digitized waveform after it has been digitized, time ordered, and interpolated at 1.63-nsec intervals with a Lagrange interpolator. Figure 5 plots the power spectrum after passing the digital record through a fast Fourier transform routine. Most of the power is contained in the two peaks at 13.2 and 24.0 MHz.

Figures 17 through 23 are plots of the mean lagged products and spectral densities for different values of the lag interval,  $\Delta t_{\rm c}$ . It is seen that all of the frequency content of the power spectrum is accurately calculated until the lag interval exceeds the Nyquist sampling rate,  $\Delta$  >  $t_{\rm N}$  = 1/2f  $_{\rm max}$ , where f  $_{\rm max}$  is the largest expected frequency content of the record. For these data,  $\Delta t_{\rm N}$  = 20 nsec. Figure 22 shows the power spectrum clearly broadened and thus fold-over effects on the lower frequency peak. Since the autocorrelation function is close to zero for times greater than 0.5 µsec, the power spectrum can be accurately calculated with a lag of 1.0 and a lesser number of mean lagged products (fig. 23).



The log Plot of mean lagged products,  $C_r$ ,  $r=1,\ldots,256$ . interval,  $\Delta \tau$ , is given by  $\Delta \tau = H \cdot \Delta \tau = 3.61$  nsec. Figure 17.

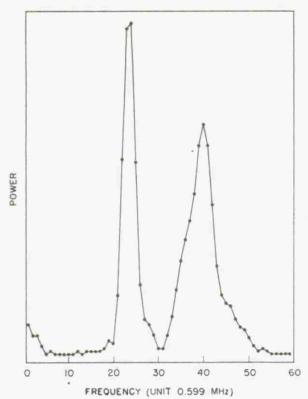
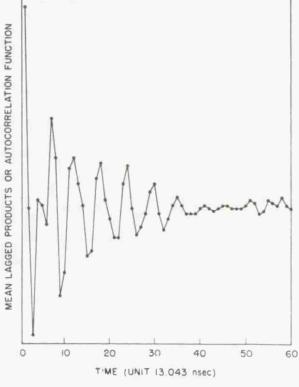


Figure 18. Refined spectral density estimates for the autocorrelation function of figure 17.





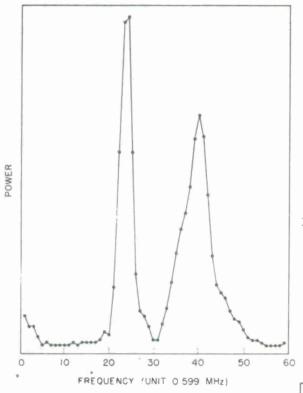
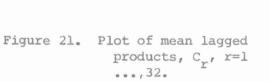
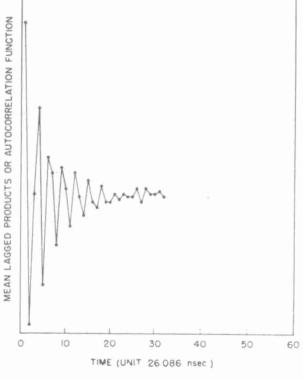


Figure 20. Refined spectral density estimates for the autocorrelation function of figure 19.





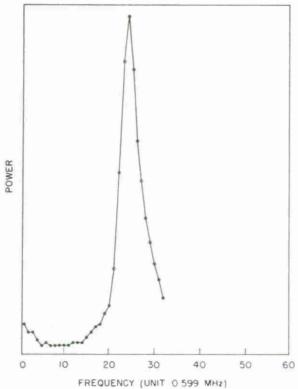
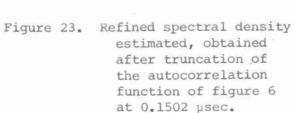
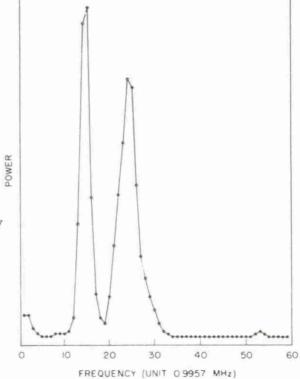


Figure 22. Refined spectral density estimates for the autocorrelation function of figure 21.





#### 3. PROGRAM DESCRIPTION AND INPUT FORMATS

A flow chart of the main program is presented in figure 24. Each subroutine performs a definite signal analysis function. As indicated, the user is privileged to employ the technique of his choice by setting the control parameters on the input cards. The subroutines and their functions are listed.

### 3.1 Subroutine Descriptions

MAIN: This routine directly or indirectly calls the rest of the subroutines and thus controls the passage of the program through all the signal analysis options.

READIN: This routine reads in all input parameters and data.

SCARTP: This routine calculates the rotation angle, the sine and cosine of the rotation angle, the origin (center) of the scope graticule, and the scale factors (tablet units/div) along both the time and voltage axes.

ROT: This routine rotates and scales the input time and amplitude arrays given in digitization-tablet units into units of time and volts.

CSTOUT: This routine checks the time ordering of the digital time series. It discards those points from the array whose time values,  $t_i$ ,  $i=1,\ldots,N$ , do not satisfy  $t_{i+1} \geq t_i$ . On those points that have the same time values, the program averages the amplitude values.

LNYQ: This routine interpolates at the Nyquist sampling intervals with a linear interpolator. If the time value, X, satisfies  $XF(I-1) \le X \le XF(I)$ , where XF is the time array, then the interpolated amplitude is Y = C1 \* X + C2 where

C1 = (YF(I) - YF(I-1)) / (XF(I) - XF(I-1))  $C2 = (YF(I-1) \times XF(I) - YF(I) \times XF(I-1)) / (XF(I) - XF(I-1))$ 

NYQST: This routine interpolates at the Nyquist sampling intervals by using a least-squares interpolator. If the time value, X, satisfies XF(I-1) < X < XF(I) where XF is the time array, then a least-squares fit of the function  $\Phi = a_1 X(I) + a_2$  is made to the set of points (XF(J), YF(J)), J=I-2, I-1, I, I+1. Thus, this routine is also a linear interpolator. Program modifications can be made to increase the number of points and/or the degree of the polynomial used in the fitting procedures.

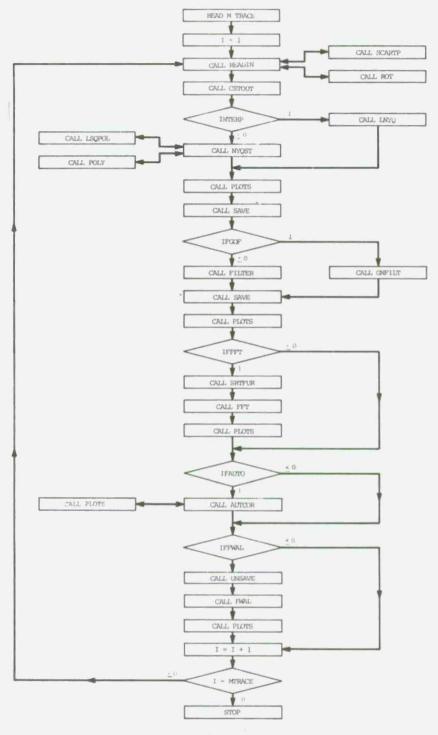


Figure 24. Flow chart for signal analysis program.

GNFILT: This routine will calculate the coefficients A(J) and B(J) for a Butterworth linear recursive low-pass digital filter. If Y(K) and U(K) are the output and input arrays to the filter, then

$$Y(K) = -\sum_{J=1}^{N} A(J) * Y(K-J) + \sum_{J=1}^{N+1} B(J) * U(K-J+1)$$
.

After calculating the coefficients, the routine filters the data.

- FILTER: This routine filters the data through a Butterworth linear low-pass digital filter when the coefficients are known.
- SRTFUR: This routine performs a bit reversal of the original Nyquist array. If I is the element of the array, then I =  $\sum_{n=0}^{m} a(n)*(2**n)$ , a(n) = 0 or 1. The element specified by I is exchanged with the element specified by J, where J =  $\sum_{n=0}^{m} a(m-n)*(2**n)$ .
  - FFT: This routine calculates the Fourier transform of a Nyquist sampled array that has its elements in bit-reversed order. The technique is generally known as the "fast Fourier transform."
  - FWAL: This routine calculates Walsh transforms of a Nyquist sampled array. The technique is generally known as the "fast Walsh transform."
- AUTCOR: This routine calculates the autocorrelation function for a Nyquist sampled array. From the autocorrelation function, the power spectrum is calculated by the method of refined spectral densities.

PLOTS: This routine plots an equispaced array.

SAVE: Utility subroutine for saving arrays.

UNSAVE: Utility subroutine for retrieving saved arrays.

# 3.2 Input Data Descriptions

Listed below are the input cards and their format.

CARD 1: READ MTRACE

FORMAT (I5)

MTRACE: The total number of the sets of data to be processed by the program.

CARD 2: READ FMT

FORMAT (20A4)

FMT: An array that contains the format of the input data arrays (X(I), Y(I)).

CARD 3: READ TITLE

FORMAT (20A4)

TITLE: An array that contains 80 characters of alpha-numeric information describing the data.

CARD 4: READ NROT, NSTAR, NPOW, ND, JD.

FORMAT (515)

NROT: If 0, then the input array must be rotated.

NSTAR: The number of points in the equispaced Nyquist array. It must be an interger power of two.

NPOW: NSTAR = 2\*\* NPOW.

ND: Used for plotting power spectra, every ND point will be plotted.

JD: Used for plotting time series, every JD<sup>th</sup> point will be plotted.

### CARD 5: READ INTERP, IFFFT, IFFWAL, IFAUTO, IFGOF

FORMAT (615)

INTERP = 0 Main calls NYQST

= 1 Main calls LNYO

IFFFT = 0 Main skips FFT

= 1 Main calls FFT

IFFWAL = 0 Main skips FWAL

= 1 Main calls FWAL

IFAUTO = 0 Main skips AUTCOR

= 1 Main calls AUTCOR

IFGOF = 0 Main calls FILTER

= 1 Main calls GNFILT.

### CARD 6: READ YSCAL, XSCAL

FORMAT (2(2X, E13.7))

YSCAL: A scale factor that multiplies the input amplitude array.

XSCAL: A scale factor that multiplies the input time values.

Note: The program assumes the time array is in nanoseconds after it is scaled.

#### CARD 7: READ JF

FORMAT (2X, I4)

JF: The number of points in the input time and amplitude arrays.

CARD 8: READ (XF(I), YF(I), I=1, JF)

FORMAT (FMT)

XF: An array which contains the initial time array.

YF: An array which contains the initial amplitude array.

CARD 9: READ XZ YZ FORMAT (2(1X, E11.4))

XZ,YZ: Are the tablet coordinates of the zero point of the trace.

CARD 10: READ X1, X2, Y1, Y2

FORMAT (4F510)

X1, X2: Coordinate labels for digitization points on the X axis of the grid.

Y1, Y2: Coordinate labels for digitization points on the Y axis of the grid.

CARD 11: READ (XT(I), YT(I), I=1,4)

FORMAT (2(1X, Ell.4))

XT,YT: Tablet coordinates for the four grid measurements defined by X1, X2, Y1, and Y2.

CARD 12: READ T, V

FORMAT (2(1X, E11.4))

T: A time-scale factor nanoseconds/division.

V: A voltage scale factor volts/division.

Note: Cards 9, 10, 11, and 12 only appear if NROT = 1.

CARD 13: READ F1, F2, GW2

FORMAT (3E13.7)

Fl: Half-power point for the low-pass digital filter given in megahertz.

F2: <F1 is another frequency at which we specify the gain, GW2, and thus determine the number of poles for the filter. F2 is in megahertz and 0 < 0.5.

Note: Card 13 only appears if IFGOF=1. If IFGOF=0 then

CARD 13: READ NP

FORMAT(IS)

CARD 14: READ (A(D), K=1, NP)

FORMAT (5(2X, E13.7))

READ (B(K), K=1, NP+1)

FORMAT (5(2X, E13.7))

If further analysis is desired repeat cards 2 through 12. The program will continue analyzing data until all input data are exhausted. If much similar data are to be processed, it would be convenient to define all the parameters internal to the program and then delete a great many input cards.

A computer printout of the signal analysis program is included as appendix A.

## 4. REPRESENTATION OF EMP WAVEFORMS BY PARAMETRIZED FUNCTIONS

As previously reported, it is very convenient to characterize an EMP-induced waveform by a finite set of parametrized functions. This allows an analyst to conveniently handle the thousands of waveforms generated in any given experiment. At present, most analysts deal with the data as a digital record consisting of  $\sim 10^3$  digital values per waveform and an equivalent number of digital values in the Fourier power spectrum. Although this is a valid approach, it is difficult and tedious to make comparisons between large sets of data and to discover trends in the data.

A large fraction of the data that is generated in EMP experiments consists of waveforms that can be described as a product of a growth factor, an exponential damping factor, and a sinusoidal function. In many of these records there is more than one dominant frequency. Most data are then fitted to a judicious mix of the functions  $\boldsymbol{\varphi}_n$ , where

$$\Phi_{n} = t^{n-1} \exp(-\xi_{n}t) \sin \omega_{n}t \qquad (76)$$

A linear combination of the  $\ \Phi_n$  generally suffices, thus a function  $\Phi$  is fit to the digital data, where

$$\Phi = \sum_{n=0}^{N} a_n \Phi_n ; \qquad (77)$$

or

$$\Phi = \sum_{n=0}^{N} a_n \Phi_n + \sum_{n'=0}^{N'} b_{n'} \Phi_{n'}$$
 (78)

There is no general rule that allows a blind selection of such functions; thus, at some point the data must be examined and a "guess" made for a good set of functions to characterize the data.

Following is a description of the method used to fit an N parameter nonlinear function to a set of data. Also, some examples of data are included, which were processed by using the computer codes especially developed for this problem. A listing of the computer program with detailed annotations is given in appendix B.

### 4.1 Method of Nonlinear Least Squares

For simplicity, consider the problem of fitting an N-parameter function with one independent variable  $X_i$ ,  $\Phi(X_i, P_i \dots P_N)$ , to the measured quantities  $Y(X_i)$  where the subscript i refers to the i data point. Thus, we find the parameters that minimize

$$S^{2} = \sum_{i=1}^{M} \left[ Y(X_{i}) - \Phi(X_{i}, P_{1}, \dots, P_{n}) \right]^{2} W_{i} , \qquad (79)$$

where  $\mbox{\tt M}$  is the number of data points, and  $\mbox{\tt W}_{\mbox{\tt i}}$  is the statistical weight of the i data point.

Let us assume now that an iterative procedure has been defined for determining the parameter P; that is, P; = the k iteration of the j parameter. Expand the function  $\Phi$  in a Taylor series about P, and truncate all but the linear terms. Then use the k iteration to determine the parameters for the (k+1) iteration. Let

$$\Phi = \Phi^{k} + \sum_{j=1}^{N} \left[ \frac{\partial \Phi}{\partial P_{j}} \right]^{k} \Delta P_{j}^{k} , \qquad (80)$$

where

$$\Delta P_{j}^{k} = P_{j}^{k+1} - P_{j}^{k} ,$$

and

$$\Phi^{k} = \Phi\left(X_{1}, P_{1}^{k}, \dots, P_{N}^{k}\right) .$$

Note that given  $P_j^k$ , we must now determine  $P_j^{k+1}$ . Substituting equation (80) into equation (79) yields in the linear approximation

$$\mathbf{S}^{2} \cong \mathbf{S}_{k}^{2} = \sum_{i=1}^{M} \left( \mathbf{Y} \left( \mathbf{X}_{i} \right) - \boldsymbol{\Phi}^{k} - \sum_{j=1}^{N} \left[ \frac{\partial \boldsymbol{\Phi}}{\partial \mathbf{P}_{j}} \right]^{k} \Delta \mathbf{P}_{j}^{k} \right)^{2} \mathbf{W}_{i} . \tag{81}$$

To minimize  $S_k^2$ , form

$$\frac{\partial S_{\mathbf{k}}^{2}}{\partial P_{\varrho}} = 0, \quad \ell=1,\dots,N, \tag{82}$$

which yields

$$\sum_{i=1}^{M} \left[ Y \left( X_{i} \right) - \Phi^{k} \right] \left[ \frac{\partial \Phi}{\partial P_{k}} \right]^{k} W_{i} = \sum_{i=1}^{M} \left[ \sum_{j=1}^{N} \left[ \frac{\partial \Phi}{\partial P_{j}} \right]^{k} \Delta P_{j}^{k} \right] \left[ \frac{\partial \Phi}{\partial P_{k}} \right]_{i}, \quad k=1,...,N. \quad (83)$$

Defining  $\psi_{i}^{k} = Y(x_{i}), - \Phi^{k}$ ,

$$z_{\ell i}^{k} = \left[\frac{\partial \Phi}{\partial P_{\ell}}\right]^{k}$$
,

$$C_{\ell}^{k} = \sum_{i=1}^{M} \psi_{i}^{k} Z_{\ell i}^{k} W_{i};$$
 (84)

then from equations (81) and (82),

$$\sum_{i=1}^{M} \left(z_{1i}^{k} z_{1i}^{k}\right) w_{i}, \dots, \sum \left(z_{1i}^{k} z_{Ni}^{k}\right) w_{i} \qquad \triangle P_{1}^{k} \qquad C_{1}^{k} \qquad C_{2}^{k} \qquad C$$

Or, more concisely,

$$\mathbf{A} \cdot \Delta \vec{\mathbf{P}} = \vec{\mathbf{C}} \quad . \tag{86}$$

Equation (83) can now be inverted to solve for the  $\Delta P_j^k$  and hence, the  $P_j^{k+1}$ . If now our iterative procedure is converging, the  $P_j^k$  should be closer to the values  $P_j^*$  which minimizes  $S^2$  and can then be used for the next iteration. This iterative procedure is continued until the use of  $P_j^k$  produces a chi-squared which differs from that using  $P_j^k$  by less than some preset value,  $V_j^k$  usually  $V_j^k = 10^{-6}$ ; that is,

$$|s_{k+1}^2 - s_k^2| < v$$
 (87)

To insure that the step  $\Delta P_{j}^{k}$  is converging, first note that in the linear approximation the quantity

$$D_{T}^{k} = \sum_{\ell=1}^{N} C_{\ell}^{k} \Delta P_{\ell}^{k} = \sum_{i=1}^{M} \sum_{j=1}^{N} \sum_{\ell=1}^{N} Z_{ji}^{k} \Delta P_{j}^{k} Z_{\ell i}^{k} W_{i} \Delta P_{\ell}^{k}, \tag{88}$$

$$= \sum_{i=1}^{M} w_{i} \left[ \sum_{j=1}^{N} z_{ji}^{k} \Delta P_{j}^{k} \right]^{2} \geq 0 ,$$

which implies the  $D_T^k$  must always be positive semidefinite. Thus, if at the  $k^{th}$  iteration  $D_T^k < 0$ , then change the sign of all the  $\Delta P_{\ell}^k$ . Another approach that shows this is to expand  $S^2$  in a Taylor series and to keep only the first deriviatives of  $\Phi$ ; that is,

$$S^{2} = S_{k}^{2} \begin{vmatrix} k & + \sum_{\ell} \frac{\partial S^{2}}{\partial P_{\ell}} \end{vmatrix}^{k} \Delta P_{\ell}^{k} + \frac{1}{2} \sum_{\ell,j} \frac{\partial S^{2}}{\partial P_{\ell}} \frac{\partial S^{2}}{\partial P_{j}} \end{vmatrix}^{k} \Delta P_{\ell}^{k} \Delta P_{j}^{k} + \dots,$$

$$= \sum_{i=1}^{M} w_{i} \left[ \left( \psi_{i}^{k} \right)^{2} - 2 \psi_{i}^{k} \sum_{k} z_{ki}^{k} \Delta P_{k}^{k} + \sum_{jk} z_{ji}^{k} z_{ki}^{k} \Delta P_{j}^{k} \Delta P_{k}^{k} \right] . \tag{89}$$

If  $\Delta P^k$  is considered as defining a vector, construct  $S^2$  to be a function of one variable  $\alpha$ , for example, by letting  $\Delta P^k_{\ell} \rightarrow \alpha \Delta P_{\ell}$ ; then we have

$$S^{2}(\alpha) = S_{k}^{2} - 2\alpha D_{T}^{k} + \alpha^{2} D_{T}^{k}$$
 (90)

Now the slope of  $S^2$   $(\alpha)$  evaluated at  $\alpha {=} 0$  is

$$\frac{ds^{2}(\alpha)}{d\alpha}\Big|_{\alpha=0} \approx -2D_{T}^{k}; \qquad (91)$$

and to reach a minimum this slope must be negative, that is,  $D_{\mathrm{T}}^{k} > 0$ , as in equation (86).

Although the condition  $D_T^k>0$  insures convergence, barring roundoff errors, the procedure can take divergent steps, thus oscillating and giving slow convergence. The following test has been used to overcome this problem and found to work satisfactorily on most data. It is first necessary to check  $D_T^k$  and, if it is negative, to change the sign of all the  $\Delta P_{\ell}^k$ . Then test to determine if the value  $S_{k+1}^2(\alpha=1) < S_k^2$ ; if it is, perform the test described by equation (85). If  $S_{k+1}^2[(\alpha)=1]>S_k^2$ ,  $S_{k+1}^2(\alpha)$  must be calculated for the following four values of  $\alpha$ . Thus,

$$\alpha = 1/2$$
,  $\alpha = D_T^k / (S_{k+1}^2 (\alpha=1) - S_K^2 + 2D_T^k)$ 

[determined by the parabola with slope -D  $_T^k$  at  $\alpha = 0$  and passing through the points  $S_k^2$  and  $S_{k+1}^2 \, (\alpha = 1)$  ],

$$\alpha = 1/D_{\rm T}^{\rm k}$$

(determined from the parabola given in equation (90)),

$$\alpha = \left[ s_{k+1}^2(\alpha=1) \ + \ 3s_k^2 \ - \ 4s_{k+1}^2(\alpha=1/2) \right] / 4 \left[ s_k^2 \ + \ s_{k+1}^2(\alpha=1) - 2s_{k+1}^2(\alpha=1/2) \right]$$

[determined by the parabola passing through the points  $S_k^2$ ,  $S_{k+1}^2$  ( $\alpha$ =1/2), and  $S_{k+1}^2$  ( $\alpha$ =1)] with the restriction that any value of  $\alpha$  < 10<sup>-2</sup> is ignored. From these four (or less) values of  $S_{k+1}(\alpha)$ , find the minimum, compare it with  $S^2$ , and perform the termination test equation (85) if  $S^2$  is improved. If, however,  $S^2$  is still not improved, then either bad data points remain or the initial starting parameters are too far from the ones that minimize  $S^2$ .

After achieving a minimum as determined by equation (87), an error analysis is performed to give an indication of how well the parameters can be determined from the data. The method used is that described by Cohen, Crowe, and Dumond. Their discussion deals only with linear least squares fitting and we have not made a detailed study for the nonlinear case. However, in the neighborhood of the minimum in  $S^2$  the linearization given by equation (81) should be a good approximation. With the above warning the error  $(\sigma_{\varrho})$  on the  $\ell^{th}$  parameter is given by,

$$\sigma_{\ell} = \sqrt{(A^{-1})_{\ell\ell} \chi^2} . \tag{92}$$

where

$$x^2 = S^2/(M-N)$$
, (93)

which is the chi-squared normalized with the number of degrees of freedom. To be conservative in our error estimate  $\chi^2$  is set equal to one if the fit gives a smaller value.

The other important quantities in the error analysis are the correlation coefficients defined by

$$\rho_{ij} = (A^{-1})_{ij} / \sqrt{(A^{-1})_{ii} (A^{-1})_{jj}}$$
(94)

The correlation coefficients are necessary for computing effects of error propagation when using the "best" parameters. Consider a function f which depends on L fitted parameters, then the error on f is given by

$$\sigma_{\mathbf{f}} = \sum_{\ell=1}^{L} \left( \frac{2f}{2P_{\ell}} \right)^{2} \sigma_{\ell}^{2} + 2 \sum_{\mathbf{i} \leq j} \rho_{\mathbf{i}j} \frac{2f}{2P_{\mathbf{i}}} \frac{2f}{2P_{\mathbf{j}}} \sigma_{\mathbf{i}}^{\alpha};$$

$$(95)$$

<sup>3</sup>E. R. Conen, K. M. Crowe, and J. W. M. Dumond, Fundamental Constants of Physics, Interscience Publishers, Inc., New York (1957), Ch. 7.

# 4.2 Parametrization of an EMP Waveform and its Autocorrelation Function

As simple examples of the use of the program the data of figure 4 has been analyzed in several different ways. The power spectrum of the time series data is shown in figure 5. It is seen that there are two predominant frequencies in the time waveform, 13.2 and 24 MHz. To isolate one of the frequencies (and obtain a less complicated time series), the data of figure 4 were digitally filtered, using a 6-pole low-pass digital filter with a half-power point at 17.5 MHz and a gain at 25 MHz of 0.05. The resulting time series is shown in figure 25, Its power spectrum is shown in figure 26. It is seen from figures 25 and 26 that the time series consists primarily of a sinusoid with frequency of 13.2 MHz. To the time series of figure 25 is fit the following function

$$\phi(t,\vec{P}) = P(1) t \exp(-P(2)t)\sin[P(3)t + P(4)]$$
 (96)

It took the program five iterations to converge to a solution. The initial estimates and the final fitted parameters are given in table II. A combined plot of the experimental data and the fitted function is shown in figure 27. Generally, the fit is good only in the central portion of the trace and it fails badly at the beginning and end of the trace. This clearly means that equation (96) is not the best representation of this trace.

From the filtered time series of figure 25, the autocorrelation (lagged products) function was formed as displayed in figure 28. The following function was fit to these data.

$$\phi(t, \vec{P}) = P(1) \exp[-P(2)t] \cos[P(3)t]$$
 (97)

It took the program two iterations to converge to a solution. The initial estimates and the final fitted parameters are given in table III. A combined plot of the data and the fitted functions is shown in figure 29. As can be seen the fit is quite good over the entire trace.

As our third example, the autocorrelation function for the data of figure 4 was calculated and is plotted in figure 30 (no filtering of any kind was done); it was fit by the following function,

$$\phi(t, \vec{P}) = P(1) \exp[-P(2)t] \cos[P(3)t] + P(4) \exp[-P(5)t] \cos[P(6)t]$$
 (98)

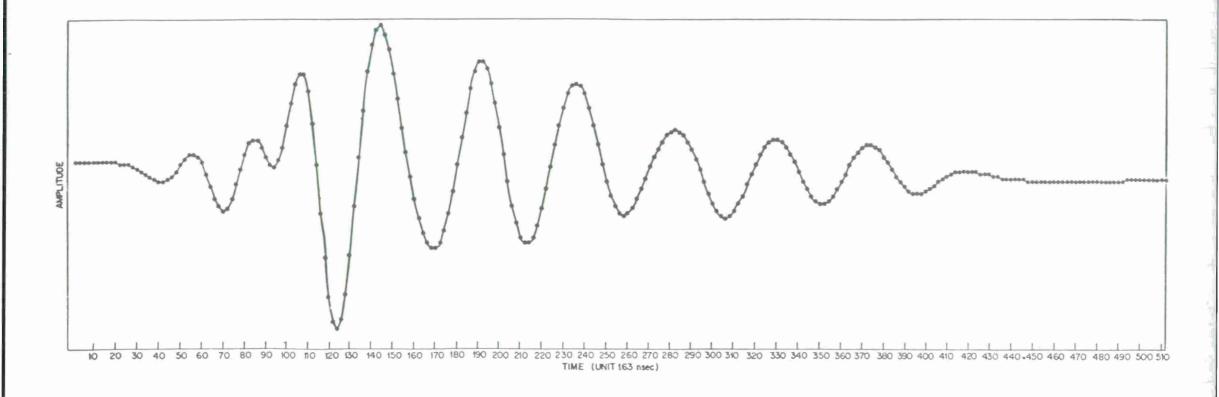


Figure 25. Time series of figure 4 after digital filtering.

The state of the s		

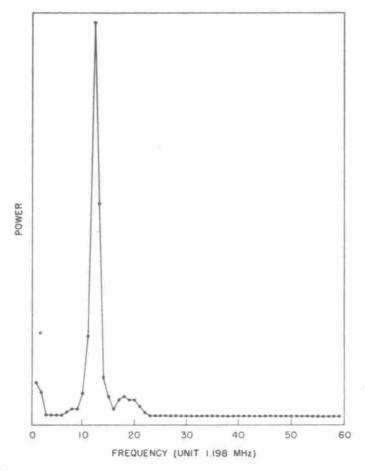


Figure 26. Power spectrum for the time series of figure 6.

It took the program three iterations to converge to a solution. The initial estimate and the final fitted parameters are given in table IV. A combined plot of the experimental data and the fitted function is shown in figure 31. The fit is quite excellent throughout the length of the trace.

Several attempts were made to fit the data of figure 4 with various forms of the functions listed in equation (76). Whenever n was greater than 2, the fitting procedure either diverged or produced a bad fit.

TABLE II. INITIAL ESTIMATES AND FINAL FITTED PARAMETERS FOR EQUATION (96) AND DATA OF FIGURE 25

	A THE TOTAL OF THE PARTY OF THE	
Parameter	Initial estimates	Final fitted values
P(1)	1.08	+0.742
P(2)	4.27	+4.958
P(3)	90.43	+84.433
P(4)	-19.59	-18.207

TABLE III. INITIAL ESTIMATES AND FINAL FITTED PARAMETERS FOR EQUATION (97)
AND DATA OF FIGURE 28

Parameter	Initial estimates	Final fitted values
P(1)	+0.001	+0.00102
P(2)	+4.27	+3.921
P(3)	+86.35	+84.196

TABLE IV. INITIAL ESTIMATES AND FINAL FITTED PARAMETERS FOR EQUATION (98) AND DATA OF FIGURE 30

Parameter	Initial estimates	Final fitted values
P(1)	0.0014	0.00108
P(2)	3.92	4.0289
P(3)	84.,20	84.7402
P(4)	0.001	0.00148
P(5)	3.92	9.8083
P(6)	147.58	146.0411

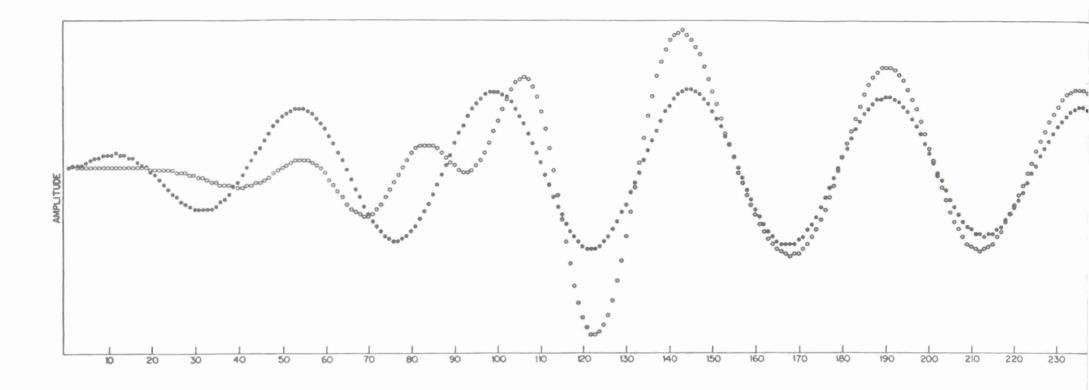
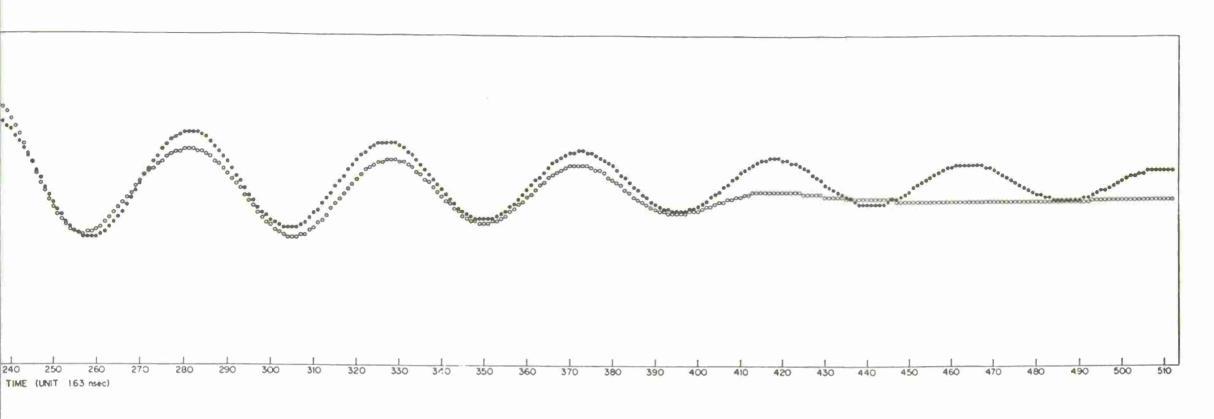


Figure 27. Plot of a least-squares fit of equation (96) to the time series of figure 6.



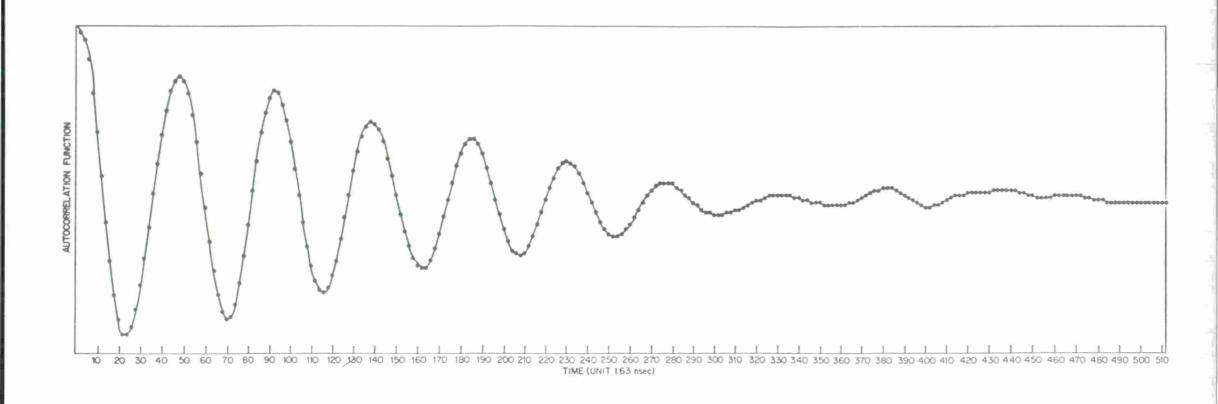


Figure 28. Autocorrelation function for the time series of figure 6.

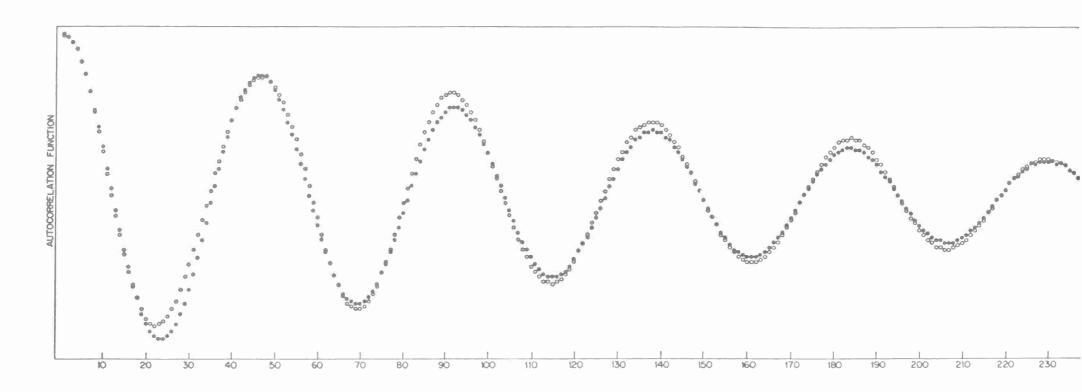
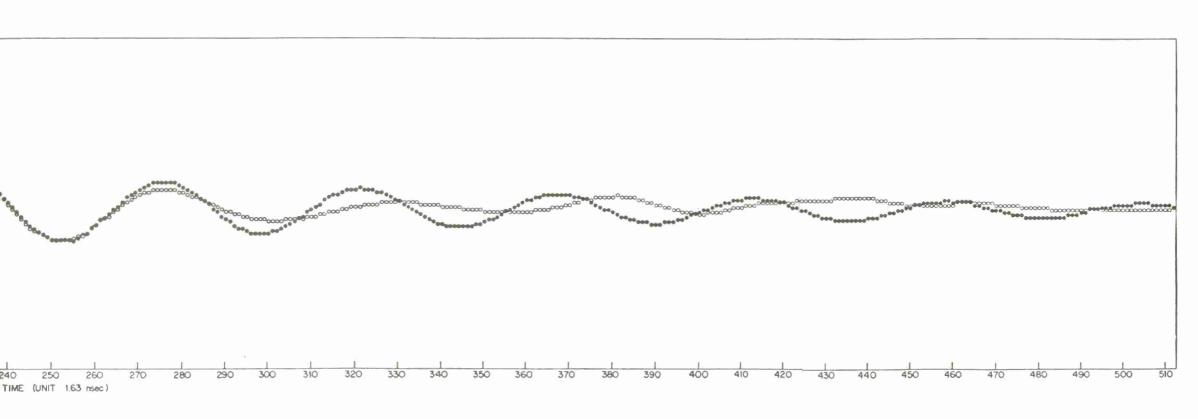


Figure 29. Plot of a least-squares fit of equation (97) to the autocorrelation function of figure 21.



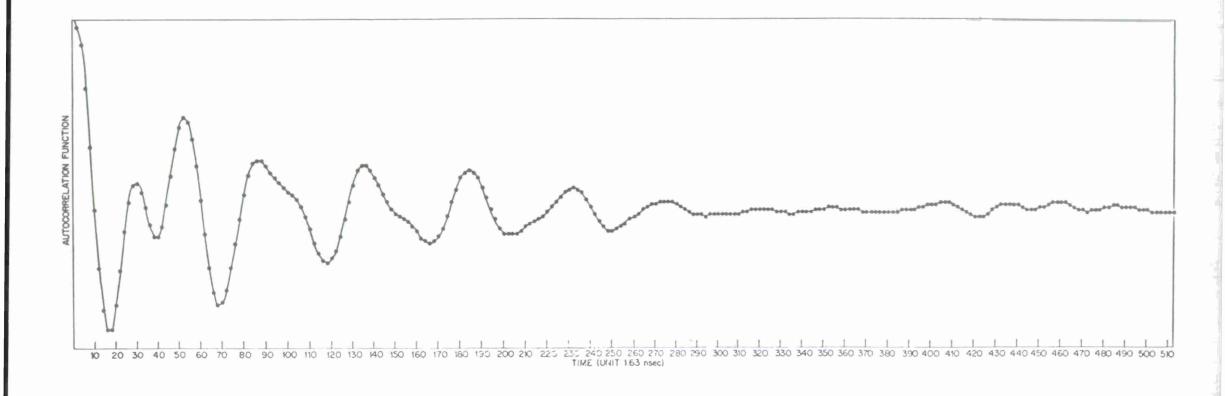


Figure 30. Autocorrelation function for the time series of figure 1.

a diam and a	
,	

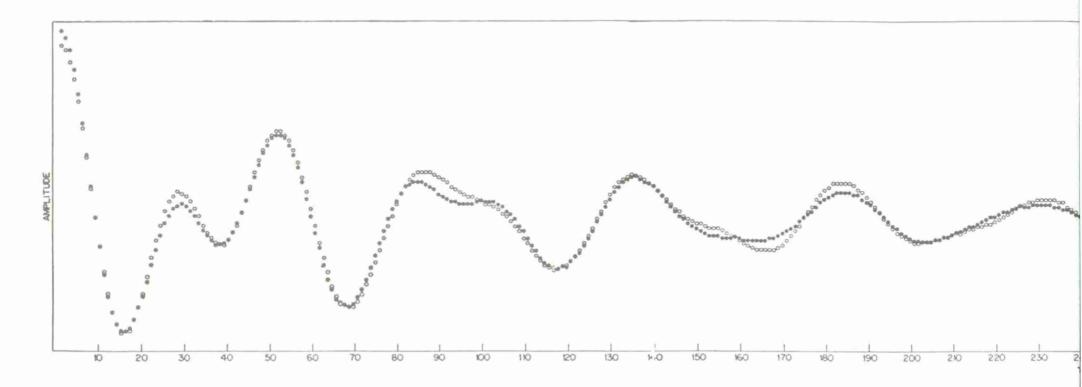
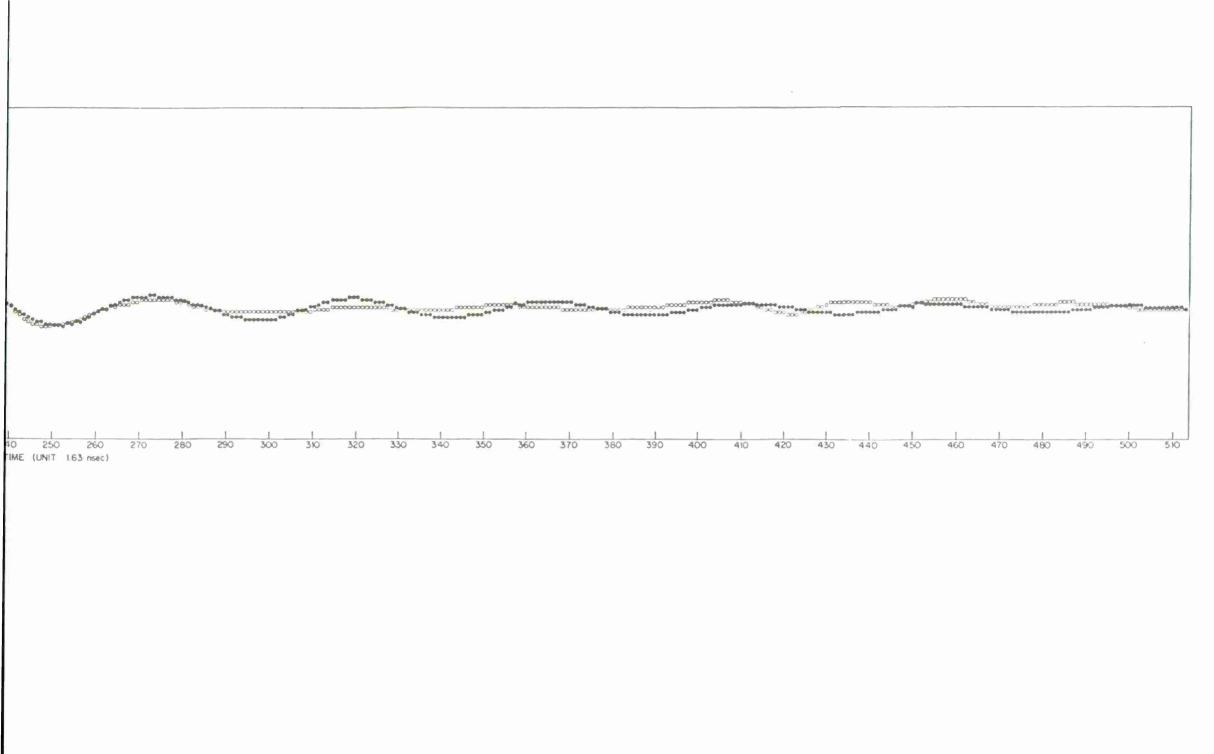


Figure 31. Plot of least-squares fit of equation (98) to the autocorrelation function of figure 11.



It seems clear that before fitting a function to the data (to obtain parametrization) a process such as digital filtering or formation of the autocorrelation function is necessary to isolate the functional dependence (at least the oscillatory part).

Following are parameters to the power spectrum of the record. Since the power spectrum,  $P(\omega)$ , is given by the cosine transform of the autocorrelation function, we have

$$P(\omega) = 2 \int_{0}^{\infty} \cos \omega t [P(1) \exp{-P(2)t} \cos{[P(3)t]} + P(4) \exp{-P(5)t} \cos{[P(6)t]} dt .$$
 (99)

Thus, after some algebraic manipulation, a power spectrum is given by

$$P(\omega) = \frac{P(1) P(2)}{[\omega - P(3)]^2 + P(2)^2} + \frac{P(4) P(5)}{[\omega - P(6)]^2 + P(5)^2}.$$
 (100)

It would have been just as easy to fit equation (100) to the power spectrum displayed in figure 5 and thus obtain a parametrization of the data. However, we have used a plot of equation (100) superimposed on the power spectrum of figure 5 as shown in figure 32.

Considering the storage of data, it is concluded that only six numbers must be stored to characterize the pulse instead of thousands. Thus, long-term storage costs are minimized and a high degree of analytical ease is realized in handling large blocks of data.

## 4.3 Program Listing and Description

The program is written in FORTRAN and consists of eleven subroutines and two user supplied functions; the program names with a description of each are listed below.

- MAIN This routine directly or indirectly calls the rest of the subroutines and thus controls the passage of the program through all the calculations used in the fit.
- READIN This routine reads in all the input parameters and data as well as setting up all the internal control parameters.
- LSQPHI This routine calculates the theoretical function, residuals, and chi square. The user must supply the function PHIFNC(I) for LSQPHI to use.

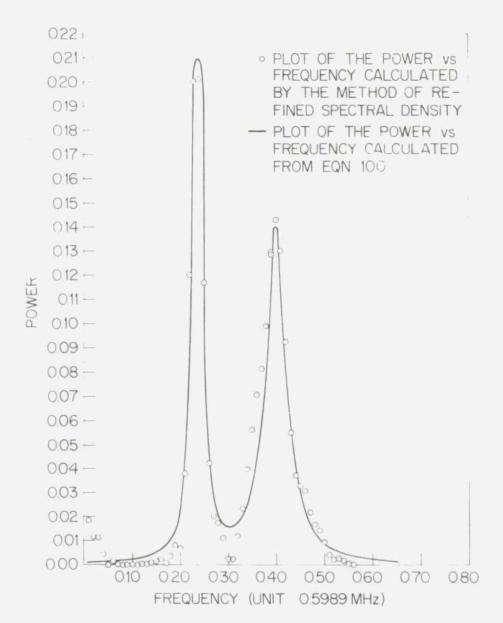


Figure 32. Plot of power spectrum obtained from equation (100) and that obtained from a numerical algorithm.

ELTS(I) This routine calculates all the derivatives of the theoretical function at the I<sup>th</sup> data point. The user must supply the function (DPHI(I)J) for ELTS to call.

GRIND This routine sets up the matrix of derivatives, the constants in the normal equations, then calls the matrix inverter.

TESTS This routine does the necessary testing of the changes in the parameters and performs the necessary changes.

ERRMAT This routine does all the calculations for a full error analysis on the fitted parameters.

CALS2 This routine is called by TESTS repeatedly for the purpose of swapping arrays and computing chi.

INVMAT Inverts a symmetric N x N matrix.

SCRIBE This routine writes out all the results of the program fit.

PLOTS Gives a plot of the experimental data and theoretical function on the same graph so that the goodness of the fit can be clearly seen

PHIFNC(I) PHIFNC =  $\phi(X(I), P(1), P(2), ..., P(N))$ 

$$\text{DPHI}\left(\text{I,J}\right) \quad \text{DPHI} \ = \ \frac{\text{d} \Phi}{\text{d} P\left(\text{J}\right)} \ \left(\text{X(I), P(1), P(2), \dots, P(N)}\right)$$

The flow chart of figure 33 shows the logical connection of the subroutines and functions.

Listed below are the input parameters, their meaning, and their format.

CARD 1 READ JSTOP, FMF FORMAT (15,5A4)

JSTOP Maximum number of iterations allowed for each fit. This is used to prevent excessive iterations in case there are errors in the experimental data or control parameters. Usually JSTOP  $\sim 100$ .

FMT The format with which the experimental data are to be read in, e.g., (6El3.7).

NOTE: Card 1 appears once independent of the number of sets of data to be fit.

CARD 2 READ TITLE FORMAT (20A4)

TITLE Any combination of alphanumeric data may appear in columns 1 through 80 to identify the fit.

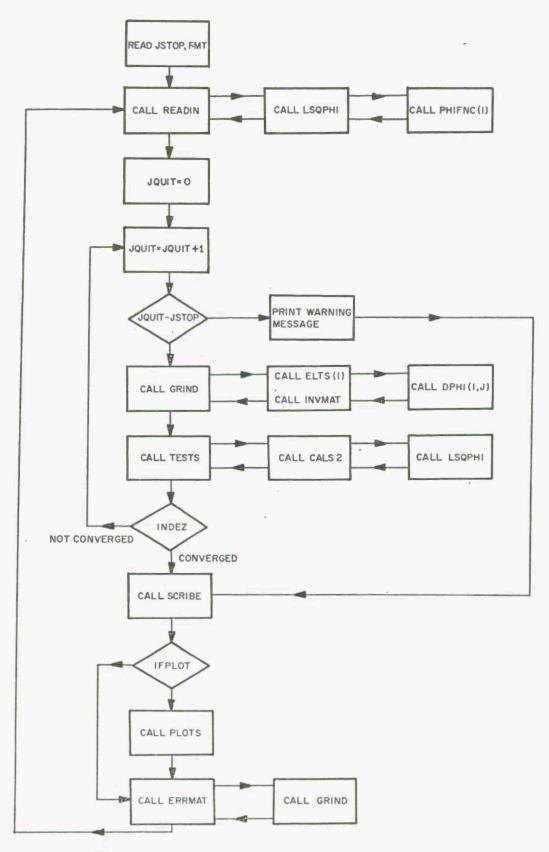


Figure 33. Flow chart of nonlinear least-squares fitting program.

CARD 3 READ NCHAN, NBADPT, NRUNS, NRP, NPSTEP, IFPLOT, IFCOR FORMAT (715)

NCHAN The number of experimental data points.

NBADPT The number of bad points in the data.

NRUNS The number of sets of experimental data to be added together.

NRP The number of independent parameters.

NPSTEP \(\geq 0\), used for plotting; that is, every NPSTEP th data point is plotted starting with the first (0 is internally set to 1).

IFPLOT = 0 no plot of data and fit
= l gives plot

IFCOR = -1 This prints and punches the correlation coefficients

= 0 Does not print or punch

= 1 Only prints

CARD 4 READ (P(I), I=1, NP)

P the array containing the fitted parameters.

CARDS 5 READ (Y(I), I=1, NCHAN) FORMAT (FMT)

Y the array containing the experimental data, one run following another for NRUNS worth of data.

CARDS 6 READ (X(I), I=1, NCHAN) FORMAT (FMT)

X the array containing values of the dependent variable for  $\phi$  and Y; for example,  $\phi(X(I), \overrightarrow{P})$ .

CARDS 7 READ (W(I), I=1, NCHAN) FORMAT (FMT)

W the array containing values of the statistical weights for the I  $\,$  data point.

## CARDS 8 READ (NBAD(I), I=J, NBADPT)

NBAD Index number of the bad data points in any order. These cards do not appear in the data deck if NBADPT = 0.

If multiple fits are desired, repeat cards 2 through 8. The program will continue doing fits until it has exhausted all of the input data.

## APPENDIX A. SIGNAL ANALYSIS PROGRAM

```
( JAN 73 )
```

## OS/360 FORTRAN H

```
COMPILER OPTIONS - NAME = MAIN.OPT=32.LINECNT=57.SIZE=3300K.
                  SOURCE . EBCDIC . NOLIST . DECK . LOAD . MAP . NOEDIT . ID . NOXREF
    C MAIN PROGRAM FOR THE ANALYSIS OF DIGITAL TIME SERIES
      PROGRAMMER DR. THOMAS A TUMOLILLO
                 USAMC HARRY DIAMOND LABORATORIES
                 WASHINGTON8D.C. 20438
      COMMON VARIABLES DEFINITION
                COMPLEX ARRAY USED TO STORE THE DIGITAL TIME SERIES AT VARIOUS
          AOYVY
                  STAGES OF THE ANALYSIS.
        NW
                  SYMBOLIC DESIGNATOR FOR THE OUTPUT LINE PRINTER
        VR
                  SYMBOLIC DESIGNATOR FOR THE INPUT CARD READER
         NT
                  SYMBOLIC DESIGNATOR FOR THE OUTPUT CARD PUNCH
    C.
         XF
                  ARRAY WHICH INITIALLY CONTAINS THE TIME VALUES
         ΥF
                  ARRAY WHICH INITIALLY CONTAINS THE AMPLITUDE VALUES
    C
         NSTAR
                 THE NUMBER OF INTERPOLATED POINTS 91T MUST BE A POWER OF TWO
    0
          NPOW
                 NSTAR=2 **NPGW
          ND .
                 USED FOR PLOTTING POWER SPECTRABEVERY ND-TH POINT WILL BE PLOTTED
    C
          JE
                 THE NUMBER OF POINTS IN THE INPUT DIGITAL TIME SERIES
                 THE COSINE OF THE ROTATION ANGLE
          CT
    C
          ST
                  THE SINE OF THE ROTATION ANGLE
          X 7
                  X-COORDINATE OF THE GRATICULE ORIGIN
          CY
                 Y-COORDINATE OF THE GRATICULE ORGIN
    C
          XS
                  SCALE FACTOR TABLET INTEGERS/SCALE DIVISION ALONG X AXIS
    C
          YS
                  SCALE FACTOR TABLET INTEGERS/SCALE DIVISION ALONG Y AXIS
          Ť
                 SCALE FACTOR NANOSECONDS/SCALE DIVISION
                 SCALE FACTOR VOLTS/SCALE DIVISION
                THE NYOUIST SAMPLING INTERVAL OR THE TIME INTERVAL BETWEEN
          DYYX
```

```
08
```

```
BETWEEN DIGITAL VALUES
      X 7
              THE X-COORD OF THE ZERO POINT OF THE TRACE
      YZ
              THE Y-COORD OF THE ZERO POINT OF THE TRACE
C
      GL
              USED FOR PLOTTING TIME SERIES. EVERY JD-TH POINT WILL BE PLOTTED
      INTERP
             INTERPO CALL NYOST .= 1 CALL LNYO
C
      IFFFT
              IFFFT=0 NO FFT.=1 CALL FFT
C
      IFFWAL
             IFFWAL=O NO WALSH TRANSFORM.=1 CALL FWAL
      IFAUTO IFAUTO=0 NO FT VIA AUTOCORRELATION.=1 CALL AUTCOR
C
      IFGOF
              IFGOF=O CALL FILTER,=1 CALL GNFILT
      LEPUNE
             IFPUNF=1 PUNCH OUT FREQUENCY AND POWER SPECTRA VALUES.=O DONT
      I F PUNW
             IFPUNW=1 PUNCH OUT SEQUENCY AND WALSH POWER VALUES.=O DON'T
C
      FPUNA
              IFPUNA=1 PUNCH DUT TIME AND AUTOCORRELATION VALUES.=O DONT
C
      IFPUND IFPUND=1 PUNCH OUT TIME AND AMPLITUDE VALUES OF INPUT DATA.=O DON
C
      IFPUNI IFPUNI=1 PUNCH OUT TIME AND INTERPOLATED VALUES,=0 DONT
C
      IFPUNR IFPUNR=1 PUNCH OUT FREQUENCY AND REFINED SPECTRAL DENSITIES.=ODON
      COMMON YNYOA (2048)
      COMMON NW.NR.XF(2048).YF(2048).NSTAR.NPOW.ND.JF.CT.ST.XO.YO.
     1 XS, YS, T, V, XNYQ
     2.NT.XZ.YZ.JD
      COMMON INTERP, IFFFT, IFFWAL, IFAUTO, IFGOF, IFPUNF, IFPUNW, IFPUNA,
     11FPUND.IFPUNI.IFPUNR
      COMPLEX YNYOA
      NW=6
      NR = 5
      NT = 7
      READ [NR.1] MTRACE
      FORMAT(15)
      DO 120 I=1.MTRACE
      CALL READIN
      CALL CSTOUT
      IF (INTERP) 13,13,23
10
      CALL NYOST
      GD TO 30
20
      CALL LNYQ
```

CONTINUE STOP END

.

CALL SAVE(O)
NPOWS=NPOW
NKEEP=NSTAR

CALL FILTER

IF(IFGOF) 40.40.50

30

40

16

```
COMPILER OPTIONS - NAME = MAIN.OPT=32.LINECNT=57.SIZE=3030K.
                    SOURCE . EBCDIC . NOLIST . DECK . LOAD . MAP . NOED IT . ID . NOXREF
            SUBROUTINE READIN
           COMMON YNYQA(2048)
           COMMON NW.NR.XF(2048).YF(2048).NSTAR.NPOW.ND.JF.CT.ST.XO.YO.
          1 XS, YS, T, V, XNYQ
           2.NT.XZ.YZ.JD
           COMMON INTERP, IFFFT, IFFWAL, IFAUTO, IFGOF, IFPUNF, IFPUNA, IFPUNA,
           1 I F PUND . I F PUNI . I F PUNR
           COMPLEX YNYOA
           DIMENSION FMT(20), TITLE(20)
           READ(NR. 1) FMT
           FORMAT(20A4)
           READ(NR. 1) TITLE
           WRITE(NW.2) TITLE
           FORMAT(1H1,20X,20A4)
           READ(NR.3) NROT, NSTAR, NPOW, ND, JD
     3
           FORMAT(515)
           WRITE (NW.13)
     13
           FORMAT (1X, NROT NSTAR NPOW ND
                                                 JDII
           WRITE(NW.14) NROT, NSTAR, NPOW, ND. JD
     14
           FORMAT([4.4[5./]
           READ(NR.3) INTERP. IFFFT, IFFWAL, IFAUTO, IFGOF
           WRITE(NW.11)
     11
           FORMAT(1X. 'INTERP [FFFT [FFWAL [FAUTO [FGOF']
           WRITE(NW.12)INTERP.IFFFT.IFFWAL.IFAUTO.IFGOF
     12
           FORMATIIX.216.217.16.617./)
           READ(NR.8) YSCAL, XSCAL
     8
           FORMAT(2(2X, E13, 7))
           WRITE (NW.15)
           FORMAT(IX, ' Y SCALE FACTOR X SCALE FACTOR ')
     15
```

WRITE (NW, 16) YSCAL, XSCAL FORMAT(2(2X, EL3.7), /)

WRITE(NW.5)

```
83
```

5

7

100

50

6

50

FORMATI' IMAGE OF INPUT DATA')

READ(NR, FMT) (XF(1), YF(1), 1=1. JF)

WRITE(NW, 4) (XF([], YF([], [=], JF)

FORMAT(/, ' SCALED AND ROTATED DATA')

WRITE(NW,4) (XF(I),YF(I),I=1,JF)

READ(NR.7) JF

FORMAT(2X, 141

DD 100 [=1, JF YF(I)=YSCAL\*YF(I) XF(I)=XSCAL\*XF(I)

CALL SCARTP CALL ROT WRITE(NW.6)

RETURN

XNYQ=XF(JF)/NSTAR

FORMAT(8(1X,E11.4))
IF(NROT-1)60,50,50

```
W XTONSAAN
```

```
COMPILER OPTIONS - NAME = MAIN, OPT=02.LINECNT=57.SIZE=0000K.
                  SOURCE, EBCDIC, NOLIST, DECK, LOAD, MAP, NOEDIT, ID, NOXREF
          SUBROUTINE GNEILT
          COMMON YNYOA (2048)
          COMMON NW.NR.XF(2048).YF(2048).NSTAR.NPOW.ND.JF.CT.ST.XO.YO.
         L XS, YS, T, V, XNYQ
         2.NT.XZ.YZ.JD
          COMMON INTERP. IFFET, IFFWAL, IFAUTO, IFGOF, IFPUNF, IFPUNW, IFPUNA,
         LIFPUND, IFPUNI, IFPUNR
          COMPLEX YNYQA, U, R, S, ROOT, BETA, UNIT
          EQUIVALENCE (ARRAY(1), YNYQA(1))
          DIMENSION ARRAY (4096)
          DIMENSION STH(40), CTH(40), INOM(40)
          DIMENSION U(40), R(20), S(20), ROOT(20), P(20),
         19(20),P2(20),A(20),B(20)
    C THIS ROUTINE WILL CALCULATE THE COEFFICIENTS A(J) AND B(J)
    C FOR A BUTTERWORTH LINEAR RECURSIVE FILTER. IF Y(K) AND U(K)
    C ARE THE OUTPUT AND INPUT ARRAYS FOR THE FILTER THEN.
        Y(K) = -SUM(1 TO N)(A(J) * Y(K-J)) + SUM(1 TO N+1)(B(J) * U(K-J+1))
      THE SAIN FACTOR FOR THE BUTTERWORTH FILTER IS GIVEN BY
        GAIN=/H(W)/**2=1/(1+(TAN(W*T/2)/TAN(W1*T/2))**(2*N))
      HERE T IS THE SAMPLING INTERVAL. N IS COMPUTED BY SPECIFYING
      THE HALF POWER POINT WI AND THE RELATIVE GAIN AT THE POINT
     W=W2>W1. AND THE TRANSFER FUNCTION H(Z) IS GIVEN BY THE TWO FORMULAS
        H(Z) = SUM(1 TC N+1)(B(J)*Z**(-J+1))/(1+SUM(1 TO N)(A(J)*Z**(-J))
```

H(Z) = BETA \* (1 + Z) \* \*N/(Z-P(1)) \* (Z-P(2) \* . . \*(Z-P(N)) .

```
753- 000
```

```
C HERE THE POLES OF THE FILTER. P(I). ARE GIVEN BY
  P(1) = (1 - TAN(ARG) * * 2 + (-1) * * 95 * 2 * TAN(ARG) * SIN(TH(1)) /
        (1-2*TAN(ARG)*COS(TH(I))+TAN(ARG)**2)
C WHERE ARG=WL*T/2,TH(I)=(I-1)*PI/N FOR N ODD, =(2*I-1)*PI/2*N
C FOR N EVEN. AND ONLY THOSE N P(I) ARE CHOSEN THAT SATISFY
C /P(1)<1.. THE NORMALIZING FACTOR BETA IS GIVEN BY
   BETA = (1-P(1))*(1-P(2))*...*(1-P(N))/2**N
C BY COMPARISON OF THE TWO EXPRESSIONS FOR THE TRANSFER FUNCTION
C THE FILTER COEFFICIENTS A(J) AND B(J) ARE DEDUCED
PI=3.1415927
C READ IN OR COMPUTE THE FREQUENCIES F1 AND F2. THE GAIN AT F2.GW28 AND
C THE SAMPLING INTERVAL
     TIME=0.001*XNYQ
     READ(NR.1) F1.F2.GW2
     FORMAT (3E13.7)
     WRITE(NW.2) F1.F2.GW2,TIME
     FORMAT(1H1, THE HALF POWER POINT IS ', E13.7, 'MHZ', /.
    1' THE GAIN AT ', E13.7, 'MHZ IS', E13.7,/.
     2' THE SAMPLING INTERVAL IS '.El3.7. MICROSECONDS'./)
C COMPUTE AND PRINT OUT THE NUMBER OF POLES FOR THIS FILTER
     NMAX=10
     ARG1=F1*TIME*PI
      TAG1 = TAN(ARGI)
```

```
ARG2=F2*TIME*PI
      RT = (TAN(ARG2)/TAGI)**2
      TVAL = (1.0-GW2)/GW2
      N = 1
      NJE=-1
10
      IF (RT**N.GE.TVAL) GO TO 20
      N = N + 1
      NOF=-NOE
      IF(N.GT.NMAX) GO TO 25
      GO TO 10
25
      WRITE(NW.3) NMAX
      FORMAT(1X, THE NUMBER OF POLES FOR THE FILTER EXCEEDED 1,/,
     1' THE ALLOWED LIMIT OF', [5]
      STOP
C
C COMPUTE ALL THE 2*N POLES WHICH ARE EITHER INSIDE OR OUTSIDE THE
C UNIT CIRCLE
C
20
      TWON=2*N
      I TWON= TWON
      FN=N
      WRITE(NW.4)
      WRITE(NW,5)
      DO 40 I=1, [TWON
      IF(NGE.EQ.1) GO TO 30
      TH=(I-1)*PI/EN
      GD TD 35
      TH=(2*[-1)*P[/TWON
30
35
      CTH(I)=COS(TH)
      STH(I)=SIN(TH)
      CONTINUE
40
      R1=1-TAG1**2
      R2=2*TAG1
```

D1=1+TAG1\*\*2

```
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```

NPI=N+I

```
APPENDIX A
```

```
K = 0
      DO 50 I=1. ITWON
      DENOM=D1-R2*CTH(I)
      P(I)=RI/DENOM
      Q[[]=R2*STH[[]/DENOM
      PQ(I)=P(I) ** 2+Q(I) ** 2
      WRITE(NW.6) 1,P(I),Q(I),PQ(I)
C
C SELECT ONLY THOSE POLES WHICH ARE INSIDE THE UNIT CIRCLE
C
      IF(PQ(I).GT.1.0) GO TO 50
      K = K + 1
      ROOT(K) = CMPLX(P(I), Q(I))
      CONTINUE
50
      FORMAT(1H/, 'INDEX', 5X, 'REAL PART', 5X, 'IMAG PART', 8X, 'SQUARE')
      FORMAT(IH/, 20x, 'FILTER POLE LOCATIONS')
      FJRMAT(1X, 15, 3(1X, E13, 7))
      IF (K.EQ.N) GD TO 60
      WRITE(NW.7) K, ITWON
      FORMAT(IH/, 'ONLY', 13, 'POLES OUT OF', 13, 'ARE INSIDE UNIT CIRCLE'
     1,/,1x,' PROGRAM EXECUTION HALTED')
      STOP
C COMPUTE BETA
C
60
      BETA=CMPLX(1.0.0.0)
      UNIT=BETA
      DD 70 I=1,N
7.0
      BETA=BETA*(UNIT-ROOT(I))/2.0
C COMPUTE THE BINOMIAL COEFFICIENTS
      INOM(1)=1
```

```
D7 75 I=2, NP1
      IVOM(I) = ((N-I+2) * INOM(I-I))/(I-I)
75
      WRITE(NW.14) N
14
      FORMAT( BINOMIAL COEFFICIENTS FOR (1+Z)**, 12,/,
     1' INDEX COEFFICIENT')
      WRITE(NW, 15) (I, INOM(I), I=1, NP1)
15
      FORMAT (4X, 12, 7X, 14)
C
C COMPUTE THE DIGITAL FILTER COEFFICIENTS
      IF(N-2180,90,100
      U(1) = -ROOT(1)
80
      GJ TO 125
90
      U(3)=CMPLX(1.0,0.0)
      U(2) = -RJOT(1) - ROOT(2)
      U(1)=RODT(1)*ROOT(2)
      GO TO 125
100
      U(1)=ROJT(1)*ROOT(2)
      U(2) = -ROOT(1) - ROOT(2)
      U(3) = CMPLX(1.0.0.0.0)
      L = 3
      R(2) = CMPLX(1.0,0.0)
      DO 120 K=3.N
      R(1) = -RGOT(K)
      00 110 I = 1, L
      S(1)=U(1)
110
      CALL POLMLTIZ, R, L, S, LR, U1
      L=LR
      CONTINUE
120
125
      WRITE(NW,8)
      FORMATILH/, INDEX', 5X, "REAL A(I)", 5X, "IMAG A(I)")
8
      DO 130 I=1.N
      IS=N+1-I
      RA=REAL (U(IS))
```

```
68
```

```
A(I) = RA
      AA = AIMAG(U(IS))
      WRITE(NW,9) I,RA,AA
130
      FORMAT(1X, 16, 2(1X, E13.7))
9
      NPI=N+1
      WRITE(NW,11)
      FORMAT(1H/, * INDEX', 5X, 'REAL B(1)', 5X, 'IMAG B(1)')
1.1
      NP2=N+2
      DO 140 I=1,NP1
      NR=NP2-I
      SHIT=[NOM(NR)
      S(I)=CMPLX(SHIT, 0.0) *BETA
      RB=REAL(S(I))
      B(1)=RB
      B3=AIMAG(S(I))
      WRITE(NW,9) I,RB,BB
140
C DIGITALLY FILTER THE INPUT TIME SERIES
C
      DO 170 K=1, NSTAR
      7 = 0 . 0
      DJ 150 J=1,NP1
      IVD=K-J+1
      IF([ND.LE.0] GO TO 155
      Z=Z+B(J) *AIMAG(YNYQA(IND))
150
      CONTINUE
      DJ 160 J=1,N
155
      IVD=K-J
      IF(IND.LE.0) GO TO 165
      Z=Z-A(J) *REAL(YNYQA(IND))
160
      CONTINUE
165
      X=AIMAG(YNYQA(K))
      YNYQA(K)=CMPLX(Z,X)
170
      CONTINUE
```

```
DD 180 K=1,NSTAR

X=REAL(YNYQA(K))

180 YNYQA(K)=CMPLX(X,0.0)

NS2=2*NSTAR

C
C PRINT OUT AND PLOT THE FILTERED TIME SERIES

C
WRITE(NW,12)
WRITE(NW,13) (ARRAY(I),I=1,NS2,2)

12 FORMAT(* FILTERED TIME SERIES*)

13 FORMAT(8(IX,EI3.7))
RETURN
END
```

DD 170 K=1.NSTAR

```
COMPILER OPTIONS - NAME = MAIN.OPT=32.LINECNT=57.SIZE=3333K.
                    SOURCE . EBCDIC . NOLIST . DECK . LOAD . MAP . NOED IT . ID . NOXREF
           SUBROUTINE FILTER
           THIS IS A RECURSIVE FILTER ROUTINE WHERE NP IS THE NUMBER OF POLES. A
           AND B ARE THE FILTER COEFFICIENTS.
           AIMAGIYNYDA) IS THE INPUT ARRAY.
     C
           REALLYNYOA) IS THE OUTPUT ARRAY
           REAL(YNYQA(I))=SUMB(J) *AIMAG(YNYQA(I-J+L) - SUMA(J)*REAL(YNYQA(I-J)).
           FOR NO FILTERING SET B(1)=1.B(K)=0.K.NE.1;A(K)=0 ALL K.
           COMMON YNYOA(2048)
           COMMON NW.NR.XE(2048).YE(2048).NSTAR.NPOW.ND.JE.CT.ST.XO.YO.
          I XS.YS.T.V.XNYQ
          2.NT.XZ.YZ.JD
           COMMON INTERP, IFFFT, IFFWAL, IFAUTO, IFGOF, IFPUNF, IFPUNW, IFPUNA.
          11FPUND, IFPUNI, IFPUNR
           COMPLEX YNYDA
           DIMENSION ARRAY (4096)
           DIMENSION A(10).B(10)
           EQUIVALENCE (ARRAY(1), YNYQA(1))
           READINR, 31 NP
           FORMAT([5]
     3
           READ(NR, 4) (A(K), K=1, NP)
           NPI=NP+1
           READ(NR, 4) (B(K), K=1, NP1)
           FORMAT(5(2X, E13, 7))
           WRITE(NW.101)
           FORMAT(1H1, FILTER COEFFICIENTS')
     1.01
           WRITE(NW.102) (K,A(K),K,B(K),K=1,NP)
           FORMAT(1X, 'A(',12,')=',E13.7,'B(',[2,']=',E13.7)
     102
           WRITE(NW.103) NPI.B(NPI)
           FORMAT(21X, '8(',12,')=',E13.7)
     103
     C DIGITALLY FILTER THE INPUT TIMESERIES
```

```
Z = 0 . 0
       DO 150 J=1.NP1
       IND=K-J+1
       IF(IND.LE.0) GO TO 155
       Z=Z+B(J) *A[MAG(YNYQA(IND))
150
      SUNTINUE
155
      DD 160 J=1.NP
       IND = K - J
      IF (IND.LE. 3) GO TO 165
      Z=Z-A(J) *REAL (YNYQA(IND))
160
      CONTINUE
165
      X=AIMAG(YNYQA(K))
      YNYQA(K) = CMPLX(Z, X)
170
      CONTINUE
      DJ 180 K=1.NSTAR
      X=REAL (YNYQA(K))
180
      YNYQA(K) = CMPLX(X.O.O)
C PRINT OUT THE FILTERED TIME SERIES
C
      NS2=2*NSTAR
       WRITE(NW, 12)
       WRITE(NW, 13) [ARRAY [1], I=1, NS2, 2)
       FORMAT(30X, FILTERED TIME SERIES',/)
12
13
      FORMAT (8(1X, E13, 7))
       RETURN
```

END

SOURCE . EBCDIC , NULIST , DECK , LOAD , MAP , NOED IT , ID , NOXREF

CDMMUN NW.NR.XF(2048).YF(2048).NSIAR.NPOW.ND.JF.CT.ST.XO.YO.

COMMON INTERP, IFFFT, IFFWAL, IFAUTO, IFGOF, IFPUNF, IFPUNW, IFPUNA,

COMPTLER OPTIONS - NAME = MAIN.OPT=32.LINECNT=57.SIZE=3000K.

SUBROUTINE INYO COMMON YNYCA (2048)

1 XS, YS, T, V, XNYQ 2. VT. XZ. YZ. JD

LIFPUND, IFPUNI, IFPUNR

20

30

40

90

100

110

115

```
COMPLEX YNYOA
      DIMENSION XOUT(4), YOUT(4)
C THIS ROUTINE INTERPOLATES AT THE NYQUIST INTERVALS
C USING A STRAIGHT LINE AS THE INTERPOLATING POLYNOMIAL
C IF XF(1-1)<X<XF(I) THEN THE INTERPOLATED VALUE
C IS Y WHERE Y=C1*X*C2 WITH
C = (YF(I) - YF(I-1))/(XF(I) - XF(I-1))
C = (Y = (I-1) * X = (I) - Y = (I) * X = (I-1)) / (X = (I) - X = (I-1))
      IFLAG=-1
      1 = 1
      I = 1
      IF(XF(I),GF,O)GO TO 30
      I = I + I
      IF(1-JF)20,150,150
      15 T = T
      YNYOA(L) = CMPLX(0.0.YF(I))
      X = XF(JST) + L * XNYQ
     IF(X-XF(1))100,130,140
     IFLAG=IFLAG+1
      DENOM= XE(I) - XE(I-1)
      C1 = (YE(I) - YE(I-I)) / DENOM
      C2=(YF([-1) * XF([) - YF([] * XF([-1]) / DENOM
     Y=C1 = X + C 2
     1 = 1 + 1
      1F(L-NSTAR)120,120,150
```

```
120
     YNYQA(L) = CMPLX(0.0, Y)
      GD TO 40
130
      Y = Y F (I)
      GO TO 116
140
     IFLAG=-1
      I = I + I
      IFII-JF190,90,145
145
     I = JF
      GO TO LLO
150
     WRITE(NW, 3)
      WRITE(NW.1)
      00 160 K=1, NSTAR, 4
      DO 155 J=1.4
      JP = J + K - 1
      XUIT(J) = XF(JST) + (JP-1) * XNYQ
155
      YOUT (J) = A[MAG(YNYQA(JP))
      WRITE(NW, 2) [(XOUT(L), YOUT(L)), L=1,4)
160
      SUNTINUE
      FORMAT(8(2X, E13.7))
      FORMAT (4(11X, 'TIME', 6X, 'INTERP. Y'1)
      FORMAT(IHI.20X, 'LISTING OF THE LINEARLY INTERPOLATED TIME SERIES')
3
      RETURN
      END
```

```
9
```

```
COMPTIER OPTIONS - NAME - MAIN.OPT.SC-TROSHIER OPTIONS - NAME - N
                                                                SOURCE . EBCDIC . NOLIST . DECK . LOAD . MAP . NOED IT . ID . NOKREE
                                    SUBROUTINE NYOST
                C NYOST TAKES NON-EQUI SPACED ARRAY YF.XF AND PRODUCES AN EQUISPACED COMPLEX
                C ARRAY YNYQA. INITIALLY RE(YNYQA) = O AND [M(YNYQA)=INTERPOLATED VALUES
                C EQUISPACED ARRAY.
                C YF(I). I=I.JF GIVEN AS V/DIV
                C XF(I) . I=1.JF GIVEN AS NS/DIV
                C. XNYO IS THE NYOUIST TIME INTERVAL IN NS/DIV
                C NSTAR = 2**NPOW IS GIVEN
                C INITIALIZE VALUES
                                    DIMENSION KOUT(4), YOUT(4)
                                    COMMON YNYGA (2048)
                                   COMMON NW.NR.XF(2048),YF(2048),NSTAR,NPOW,ND,JF,C1,ST,X0,Y0,
                                  1 XS.YS.T.V.XNYQ
                                 2.NT.XZ.YZ.JD
                                   COMMON INTERP, IFFFT, IFFWAL, IFAUTO, IFGOF, IFPUNF, IFPUNA, IFPUNA,
                                  LIFPUND, IFPUNI, IFPUNR
                                    COMPLEX YNYOA
                                    DIMENSION A(4.4), B(4)
                                    LTSAV=0
                                    1 B SA V= 3
                                    1 SA V=0
                                    1 = 1
                                    ISI=1
                C
                (
                                     SEARCH FOR FIRST XF > 0
                C
                 20
                                   IF(XF(ISI).GE.OIGO TO 30
                                    IST=IST+1
                                    IF(IST - JF) 20,150,150
                30
                                    JSTRT=IST
```

```
YNYQA(L)=CMPLX(0.0,YF(JSTRT))
       I = JSTRT
      XVAL = XF (JSTRT) + L * XNYQ
70
75
      IF (XF(I) GT. XVALIGO TO 80
      [ = [ + ]
      IF(I - JF) 75,115,75
83
      LT = I + I
      18=1-2
      IFILB .LE. JSTRTIGO TO 110
      I = I - 1
85
      1=1+1
C
      CHECK IF BOUNDS ARE SAME AS LAST TIME
      IF(LT.EQ.LISAV .AND. LB.EQ.LBSAV .AND. I.EQ.ISAV)GO TO 100
      NEW BOUNDS
      LTSAV=LT
      L3 SAV=LB
      ISAV=I
      IF(L - NSTAR) 95,95,120
      FIT THE FOUR POINTS TO A CUBIC
95
      CALL LSQPOL(A, B, 2, LB, LT)
      DENDM = A(1,1) *A(2,2) - A(1,2) **2
      11 = A(2,2) * B(1) - A(1,2) * B(2)
      T2 = A(1,2) * B(1) - A(1,1) * B(2)
      B(1)=T1/DENOM
      B(2) = -T2/DENOM
      YNYQA(L)=CMPLX(0.0,POLY(B,2,XVAL))
100
      IF (L-NSTAR) 70,120,120
```

```
97
```

```
APPENDIX
```

```
NEAR BEGINNING OF TRACE
C
C
110
      LB=JSTRT
      LT=L3+3
      I=18
      GO TO 85
      NEAR END OF TRACE
115
      L8=JF-3
      LT=JF
      I = I - 1
      GO TO 85
150
      WRITE(NW, 1)
      FORMAT(IX, ' ALL TIMES ARE NEGATIVE')
      STOP
      WRITE(NW,3)
120
      WRITE (NW, 4)
      )) 200 K=1,NSTAR,4
      DJ 190 J=1.4
      IND=K+J-1
      XOUT(J) = XF(JSTRT) + ([ND-L] * XNYQ
190
      YOUT(J)=A[MAG(YNYQA([ND))
      WRITE(NW, 2)(((XOUT(L), YOUT(L)), L=1,4))
      CONTINUE
200
2
      FORMAT(8(2X,E13.7))
      FORMAT (1H1, 'LISTING OF LEAST SQUARES INTERPOLATED TIME SERIES')
3
      FORMATI4(LLX, 'TIME', 6X, 'INTERP. /'))
      RETURN
      END
```

```
CUMPILER OPTIONS - NAME = MAIN.OPT=02.LINECNT=57.SIZE=0000K,
                   SOURCE, EBCDIC, NOLIST, DECK, LOAD, MAP, NOEDIT, ID, NOXREF
           SUBROUTINE AUTCOR(K.NSAVE.NPSAV)
           COMMON YNYOA (2048)
           COMMON NW.NR.XF(2048).YF(2048).NSTAR.NPOW.ND.JF.CT.ST.XO.YO.
          1 XS.YS.T.V.XNYQ
          2.NT.XZ.YZ.JD
           COMMON INTERP, IFFFT, IFFWAL, IFAUTO, IFGOF, IFPUNF, IFPUNW, IFPUNA,
          IIFPUND.IFPUNI.IFPUNR
           COMPLEX YNYOA
           DIMENSION FREQ(2048).POWER(2048).ZF(2048).OUT(8)
           DIMENSION ARRAY(4096)
           EQUIVALENCE (ARRAY(1), YNYQA(1))
           FQUIVALENCE (FREQ(1), ZF(1)), (POWER(1), XF(1))
           WRITE(NW.1)
           FORMAT('I LISTING OF THE AUTOCORRELATION FUNCTION')
           PI = 3.1415927
     C
     C COMPUTE AND PRINT OUT THE VALUES OF NSAVE.M. AND IH
    C
           KMI = K - 1
           IH=1
           M=NSAVE+NSAVE*KM1/10
           NSTAR = M
           WRITE(NW, 1001) NSAVE, IH, M
    C
    C COMPUTE DTAU=IH*XNYQ
           DTAU=IH*XNYQ
    1001 FORMATI' THERE ARE', 13, VALUES IN THE DIGITAL TIME SERIES', /,
          I' THE LAG INTERVAL IS DTAU=IH*DT WHERE IH=',13./.'THERE ARE',13.'
          2VALUES IN THE DIGITAL AUTOCORRELATION FUNCTION 1./)
    C
    C COMPUTE THE AUTOCORRELATION FUNCTION AND STORE IT IN THE ARRAY XF
```

```
APPENDIX A
```

```
C
      DO 30 IR=1.M
      SUM=0.0
      [NMHR=NSAVE-IH*([R-1]
      DD 20 19=1, INMHR
      [MESS=10+[H*(IR-1)
20
      SUM=SUM+ YF (IQ) * YF (IMESS)
      SUM=SUM/[NMHR
      XF(IR)=SUM
30
      CONTINUE
C
C PRINT OUT AND PLOT THE AUTOCORRELATION FUNCTION
C
      WRITE(NW.5)
      FORMAT(IX, 4(10X, 'TIME', 7X, 'AUTOCOR'))
5
      DD 35 I=1.M
     YYYQA[I] = CMPLX[XF[I],0.0]
35
      XNYSV=XNYQ
      XNYQ=DTAU
      D3 37 I=1,M,4
      DO 36 J=1.4
      DUT(2*J-1)=XNYQ*(1-1+J-1)
36
      JUT(2*J) = XF(I+J-1)
      WRITE(NW,4) (CUT(L),L=1.8)
      FORMAT(4(7X,E13.7))
69
4
     FJRMAT(1X,8(1X,E13.71)
     CONTINUE
37
      CALL PLOTS(4)
      XNYQ=XNYSV
C
C COMPUTE THE RAW SPECTRAL DENSITY ESTIMATES AND STORE THEM
C IN THE ARRAY ZF.
C
```

7 \*

DO 50 IR=1.M

```
IPO=IR-1
      SUM = XF(1) + XF(M) * (-1.) * * IPO
      MM1 = M - 1
      DO 40 IQ=2, MM1
40
      SUM = SUM + 2.0 \times XF([Q) \times COS(P[*([Q-1] \times ([R-1])/M))
      Z=(IR)=)TAU*SUM
50
      CONTINUE
C
C COMPUTE THE REFINED SPECTRAL DENSITIES AND STORE THEM IN THE
C ARRAY XF
C
      XF(1)=0.54*ZF(1)+0.23*ZF(2)
      XF(M)=0.23*ZF(M-1)+0.54*ZF(M)
      D3 63 I=2,MM1
      XF([]=0.23*ZF([-1]+0.54*ZF([]+0.23*ZF([+1]
60
      DJ 70 [=1.M
      ZF(I) = (I-1)/(2*M*DTAU)
73
C
C PRINT OUT AND PLOT THE REFINED SPECTRAL DENSITIES
      WRITE(NW.7)
7
      FORMATI'IREFINED SPECTRAL DENSITY ESTIMATES FOR THE AUTOCORRELATIO
     IN FUNCTION 1
      WRITE(NW.3)
      FORMAT(1H/,4(5X, 'FREQUENCY',9X, 'POWER'),/)
3
      D3 80 I=1,M,4
      WRITE(NW,4) (FREQ(I+J-L), POWER(I+J-L), J=1,4)
80
      NSTAR = 2 * M
      XNYSV=XNYQ
      UAT(=QYVX
      D7 93 I=1.M
90
      YNYQA(I) = CMPLX(POWER(I), 0.0)
      CALL PLOTS(2)
      XYYQ=XNYSV
      NS TAR = NSAVE
      RETURN
      END
```

```
COMPILER OPTIONS - NAME = MAIN, OPT=02, LINECNT=57, SIZE=0000K.
                    SOURCE, EBCDIC, NOLIST, DECK, LOAD, MAP, NOEDIT, ID, NOXREF
            SUBROUTINE FET
     C
     C THIS ROUTINE CALCULATES THE FOURIER TRANSFORM OF A NYOUIST SAMPLED ARRAY
     C THAT HAS ITS ELEMENTS IN BIT REVERSED ORDER. THE TECHNIQUE IS GENERALLY
     C KNOWN AS THE FAST FOURTER TRANSFORM.
     C
           COMMON YNYGA (2048)
           COMMON NW.NR.XF(2048).YF(2048).NSTAR.NPOW.ND.JF.CT.ST.XO.YO.
           1 XS, YS, T, V, XNYQ
           2, VI, XZ, YZ, JD
           COMMON INTERP, IFFFT, IFFWAL, IFAUTO, IFGOF, IFPUNF, IFPUNW, IFPUNA,
          11FPUND.IFPUNI.IFPUNR
           COMPLEX YNYQA, W, WE, FACTOR, SWAP, ZIP, S
           TP1=6, 2831854
           DD 40 I=1. NPOW
           T [ = 2 * * 1
           ARG=TPI/TI
           S=CMPLX(COS(ARG), SIN(ARG))
           [ \forall 1 = 1 - 1]
           11=2**IM1
           KSTUP=II+I
           K = 1
           W=CMPLX(1.0.0.0)
     20
           SWAP=YNYQA(K)
           KPII=K+II
           FACTOR=YNYQA(KPIL)*W.
           YNYQA(K) = SWAP+FACTOR
           YNYOA(KP[1])=SWAP-FACTOR
           K = K + 1
           W=W+S
           IF (K-KSTOP)20,30,20
     30
           K=KP[]+1
```

```
KSTOP=K+II
     W=CMPLX(1.0,0.0)
     IF (KSTOP.LE.NSTAR) GO TO 20
40
     CONTINUE
     DELF=1./(NSTAR *XNYQ)
     WRITE(NW, 1)
      XSQ=XNYQ**2
     DO 50 I=1.NSTAR.ND
      ZIP=YNYQA(I)
      RZIP=RFAL(ZIP)
      AIZIP=AIMAG(ZIP)
     FREQ=(1-1) *DELF
      PAR=(RZIP**2+AIZIP**2)*XSQ
      PHS=ATAN(ALZIP/RZIP)
      WRITE(NW, 2) FREQ, RZIP, AIZIP, PWR, PHS
     YNYQA(I)=CMPLX(PWR,PHS)
50
     CONTINUE
                                                                      POW
                                     REAL F(W)
                                                   IMAG F(W)
     FORMAT(1H1, FREQUENCY
1
                 PHASE 1)
     1ER
     FORMAT(5(1X, E13.7))
      RETURN
     END
```

K = K + 1

```
COMPILER OPTIONS - NAME = MAIN.OPT=02.LINECNT=57.SIZE=0000K.
                    SOURCE, EBCDIC, NOLIST, DECK, LOAD, MAP, NOEDIT, ID, NOXREF
            SUBROUTINE FWAL
     C
     C THIS ROUTINE CALCULATES WALSH TRANSFORMS OF A NYOUIST
     C SAMPLED ARRAY. THE TECHNIQUE IS GENERALLY KNOWN AS THE FAST WALSH TRANSFORM.
            COMMON YNYDA(2048)
           COMMON NW.NR.XF(2048), YF(2048), NSTAR.NPOW.ND.JF, CT.ST.X0.Y0.
           1 XS.YS.T.V.XNYQ
          2.NI.XZ.YZ.JD
            COMMON INTERP. IFFET. IFFWAL, IFAUTO. IFGOF, IFPUNE, IFPUNW, IFPUNA.
          11FPUND.IFPUNI.IFPUNR
           COMPLEX YNYOA
           D3 10 I=1.NSTAR
           XF([]=REAL(YNYQA([]))
     10
         YF([]=0.0
           DO 40 I=1. NPOW
           [M] = [-1]
           I1 = 2 * * IM1
           KSTOP=I1+1
           L = 1
           K = 1
           J = 0
           W = 1 \cdot 0
     20
         S1 = \langle F(K) \rangle
           KP [1=K+[1
           S2=XF(KPII)
           SW2= S2 *W
           INDEX=J/2
            SIG=(-1) ** [NDEX
           YF(L)=(S1+SW2)+SIG
           YF(L+1)=(-S1+SW2)*SIG
```

```
J = J + 1
      L=L+2
      W = -W
      IF(K-KSTOP)20,30,20
30
      K=KPI1+1
      KSTOP=K+II
      J = 0
      W = 1
      IF (KSTOP.LE.NSTAR) GO TO 20
      DO 35 [2X=1, NSTAR
      XF([ZX] = YF([ZX])
35
      YF ( [ Z X ] = 0 . 0
      CONTINUE
40
      DELF=1./(NSTAR *XNYQ)
      WRITE(NW, 1)
      WRITE(NW.2)
      NSMI=NSTAR-I
      G=XF(1)**2
      L = 0
      YNYOA(L+1)=CMPLX(G.J.O)
      WRITE(NW, 3) L, XF(L), G
      DO 50 I=2, NSM1, 2
      L=L+1
      AS=XF(I)
      AC=XF([+1]
      G=AS**2+AC**2
      YNYQA(L+1)=CMPLX(G,0.0)
50
      WRITE(NW, 4) L, AS, AC, G
      L=L+1
      G=XF(NSTAR) **2
      YNYQA(L+1)=CMPLX(G,0.0)
      WRITE(NW.5) L, XF(NSTAR), G
      FORMAT (LHI, CUTPUT OF THE WALSH TRANSFORM!)
2
      FORMAT(' I',9X,'AS(I)',9X,'AC(I)',10X,'G(I)')
```

3 FORMAT(1X,15,14X,2(1X,E13.7))
4 FORMAT(1X,15,3(1X,E13.7))
5 FORMAT(1X,15,1X,E13.7,15X,E13.7)
RETURN
END

INTERCHANGE ROWS TO PUT PIVOT ELEMENT ON DIAGONAL

OS/360 FORTRAN H

' [ JAN 73 ]

C

```
TELEMETY 1
```

```
IFIIROW.EQ. [COLUM] GO TO 60
              00 50 J=1.N
                     SWAP=A(IROW.J)
                     A(IROW, J) = A(ICOLUM, J)
                     A(ICOLUM. J) = SWAP
50
                     CONTINUE
              SWAP=B(IROW)
              B(IROW)=B(ICOLUM)
              B ( I COL UM ) = SWAP
              PIVOT=A(ICOLUM, ICOLUM)
              INDEX([,2)=ICOLUM
60
              INDEX(I.1) = IROW
              IF(PIVOT.EQ.O.O) GO TO 130
C
      DIVIDE PIVOT ROW BY PIVOT ELEMENT
C
              A(ICOLUM, ICOLUM) = 1.0
              DO 70 J=1.N
                     A(ICOLUM, J) = A(ICOLUM, J) / PIVOT
70
                     CONTINUE
              3(ICOLUM)=B(ICOLUM)/PIVOT
C
      REDUCE NON-PIVOT ROWS
C
C
              00 90 J=1.N
                     IFIJ.EQ.ICOLUMI GO TO 90
                     T = A ( J . I C OLUM)
                     A(J, ICOLUM) = 0.0
                     DO 80 K=1.N
                            A(J.K) = A(J.K) - A(ICOLUM,K) *T
                            CONTINUE
83
                     B(J) = B(J) - B(ICOLUM) *T
                     CONTINUE
90
100
              CONTINUE
```

```
\mathbb{C}
C
      INTERCHANGE CCLUMNS
      DJ 120 [=1,N
              J = N + L - I
              IF(INDEX(J,1).EQ.INDEX(J,2)) GO TO 120
              JROW=INDEX(J,1)
              JCOLUM=INDEX(J,2)
              00 110 K=1,N
                      SWAP = A(K, JROW)
                      A(K, JROW) = A(K, JCOLUM)
                      A(K, JCOLUM) = SWAP
110
                     CONTINUE
120
              CONTINUE
130
      RETURY
      END
```

```
109
```

```
COMPILER OPTIONS - NAME = MAIN. OPT=02, LINECHT=57, SIZE=0000K,
                    SOURCE, EBCDIC, NOLIST, DECK, LOAD, MAP, NOEDIT, ID, NOXREF
            SUBROUTINE LSCPOL(A, B, N, LB, LT)
     C
     C THIS ROUTINE CALCULATES THE MATRIX A AND VECTOR B WHICH ARE USED TO SOLVE FOR
     C THE VECTOR X IN THE EQUATION A.X=B. THE VECTOR X HAS ELEMENTS WHICH CORRESPON
     C TO THE POLYNOMIAL COEFFICIENTS IN SUMCOEFFILI*Z**I. THE COEFFICIENTS ARE
     C DETERMINED IN A LEAST SQUARES SOLUTION OF THE POLYNOMIAL TO SOME DATA.
           COMMON YNYDA(2048)
           COMMON NW.NR.XF(2048).YF(2048).NSTAR.NPOW.ND.JF.CT.ST.XO.YO.
          I XS.YS.T.V.XNYO
          CL.SY.SX.TV.S
           COMPLEX YNYOA
           DIMENSION A(4,41,8(4),C(10)
           L2=2*V-1
           NPTS=LT-LB+1
           DD 10 I=1.N
           B(I)=0.0
           DO 10 J=1.N
     1.0
           A[[,]]=0.0
           DO 5 L=1.L0
     5
           C(L)=0.0
           C(1)=NPTS
           DO 15 I=LB.LI
           00 13 K=1.N
           KMI = K - 1
           IF(XF(I).EQ.O.O.AND.KMI.EQ.O) GO TO 113
           8(K)=8(K)+YF(1)*(XF(1)**KM1)
           GO TO 13
     113
           B(K) = B(K) + YF(I)
     1.3
          CONTINUE
           D3 14 L=2.L0
           \lfloor |M| \rfloor = \lfloor -1 \rfloor
```

NIHAN=NCHAN+1

PHI(NCHAN) = REAL (YNYQA(I))

```
COMPILER OPTIONS - NAME = MAIN.OPT=02.LINECNT=57.SIZE=0000K.
                    SOURCE, EBCDIC, NOLIST, DECK, LOAD, MAP, NOEDIT, ID, NOXREF
            SUBROUTINE PLCTS(N)
     C*** PLOT ONLY THE LOG OF THE POWER FOR THE MOMENT
           COMMON YNYOAL20481
           COMMON NW, NR, XF(2048), YF(2048), NSTAR, NPOW, ND, JF, CT, ST, XO, YO,
          1 XS.YS.T.V.XNYQ
          2. VT, XZ, YZ, JD
           COMMON INTERPOIFEFT. IFFWAL . IFAUTO . IFGOF . IFPUNF . IFPUNA . IFPUNA .
          LIFPUND . IFPUNI . IFPUNR
           COMPLEX YNYOA
           DIMENSION ARR(126), PHI(1024), Y(1024)
           DATA BLANK, STAR, ZERO, PLUS, BAD/1H , 1H*, 1HO, 1H+, 1HB/
           NCHAN=0
           GD TO (100,200,300,400), N
     100
           00 110 1=1.NSTAR.JD
           NCHAN=NCHAN+1
           PHI(NCHAN) = AIMAG(YNYQA(I))
     110
           Y(NCHAN) = PH((NCHAN)
           GO TO 500
     2 2 2
           NI=NSTAR/2
           DO 210 I=1.N1.ND
           NC HAN= NC HAN+ L
           PHI(NCHAN) = REAL (YNYOA(I))
     210
           Y(NCHAN) = PHI(NCHAN)
           GO TO 500
     300
           D3 313 I=1.NSTAR.ND.
           NC HAN= NC HAN+ L
           PHI(NCHAN) = AIMAG(YNYQA(I))
     310
           Y(NCHAN) = PHI(NCHAN)
           GO TO 500
           DO 410 I=1, NSTAR, JD
     400
```

```
ARR(I)=BLANK
          10
                       CONTINUE
          ( 本本本
                FIND MAXIMUM AND MINIMUM OF EITHER THE DATA OR THE FIT.
                DD 30 I=1, NCHAN, NPSTEP
                       IF (Y(I).GT.AMAX)
                                            AMAX=Y(I)
                       IF (Y(I).LT.AMIN)
                                           AMIN=Y(I)
          20
                       IF (PHI(I) . GT . AMAX)
                                              AMAX=PHI(I)
112
                       IF (PHI(I).LT.AMIN)
                                              AMIN=PHI(I)
          30
                       CONTINUE
          C***
                FIND THE BIN SIZE.
                D3IN=(AMAX-AMIN)/124.0
                FIND WHERE TO PUT THE CURVES.
                DD 8C I=1, NCHAN, NPSTEP
                       II=IFIX((PHI(I)-AMIN)/DBIN)+2
                       JJ=IFIX((Y(I)-AMIN)/DBIN)+2
          40
                       IF (II. NE. JJ) GO TO 50
                       ARR(III)=PLUS
                       GO TO 635
          50
                       ARRIJJI=ZERO
```

ARR(II) = STAR

WRITE(NW.2) K. (ARR(K). K=1.126)

GO TO (61,62,62,611,N

K = (I - I) \* JD

K=([-1] \*NO

GD TO 70

INITIALIZE FOR FINDING MAXIMUM AND MINIMUM.

INITIALIZE THE PLOTTING ARRAY TO BLANKS.

THIS ROUTINE DOES THE PLOTTING OF THE DATA AND THEORETICAL FIT.

Y(NCHAN) = PHI (NCHAN)

NPSTEP=1

WRITE(NW.11

AMAX=PHI(1) AMIN=AMAX

D7 10 I=1,126

410

500

(\*\*\*

( \*\*\*

60

61

62

70

```
ARR(II)=BLANK
            ARRIJJI = BLANK
80
           CONTINUE
     DELF=1./(NSTAR *XNYQ)
     GJ TJ (81,82,83,84),N
     WRITE(NW,4)
81
     FORMAT( PLOT OF AMPLITUDE VS TIME AFTER INTERPOLATION )
     WRITE(NW,5) XNYQ
5
     FORMAT(' TO OBTAIN TIME VALUES MULTPLY XCOORD BY', E13.7)
     RETURN
84
     WRITE(NW, 6)
     FORMAT(' PLOT OF THE FILTERED TIME SERIES')
6
     WRITE(NW,5) XNYQ
     RETURN
     WRITE(NW,7)
82
     FORMAT(' PLOT OF THE POWER VS. FREQUENCY')
7
     WRITE(NW.3) DELF
     RETURN
     WRITE(NW,8)
83
     FORMAT(' PLOT OF THE PHASE VS. FREQUENCY')
     WRITE(NW.3) DELF
     RETURN
     FORMAT(' TO OBTAIN FREQUENCY VALUES MULTIPLY X VALUE BY', E13.7)
3
     FORMAT (1H1)
1
     FORMAT (1X,14,2X,125A1)
```

END

```
COMPILER OPTIONS - NAME MAIN. OPT=02.LINECNT=57.SIZE=0000K.
                    SOURCE, EBCDIC, NOLIST, DECK, LOAD, MAP, NOEDIT, ID, NOXREF
            SUBROUTINE CSTOUT
     C
           CHECK TIME ORDERING OF THE POINTS AND CAST OUT THOSE POINTS NOT IN THE
     C
           PROPER TIME CROER
           IT ALSO AVERAGES THOSE POINTS WHICH HAVE THE SAME X VALUES
           COMMON YNYGA (2048)
           COMMON NW.NR.XF(2048).YF(2048),NSTAR.NPOW,ND,JF,CT,ST,XO,YO,
           1 XS, YS, T, V, XNYQ
          2, NT, XZ, YZ, JD
           COMMON INTERP, IFFFT, IFFWAL, IFAUTO, IFGOF, IFPUNE, IFPUNW, IFPUNA,
          LIFPUND, IFPUNI, IFPUNR
           COMPLEX YNYOA
           LSUM=1
           J = 1
           JBAD=3
           K = 2
           YSUM=YF(1)
           XSAV=XF(1)
           IF (XF(K)-XSAV)90,100,110
     80
     90
           K = K + 1
           JBAD = JBAD + 1
           IF(K-JF)83,83,113
     100
           LSUM=LSUM+1
           YSUM=YSUM+YF(K)
           K = K + 1
           IF(K-JF)80,80,30
     110
          YF(J)=YSUM/LSUM
           XF(J)=XSAV
```

IF(K-JF)120,130,30

XSAV=XF(K) YSUM=YF(K) LSUV=1 K=K+1

END

```
COMPILER OPTIONS - NAME MAIN, OPT=02, LINECNT=57, SIZE=0000K,
                    SOURCE, EBCDIC, NOLIST, DECK, LOAD, MAP, NOEDIT, ID, NOXREF
           SUBROUTINE SAVE (N)
           COMMON YNYOA (2048)
           COMMON NW.NR.XF(2048).YF(2048).NSTAR.NPOW.ND.JF.CT.ST.XD.YD.
          1 XS, YS, T, V, XNYQ
          CL.SY.SX.TV.S
           COMMON INTERP, IFFFT, IFFWAL, IFAUTO, IFGOF, IFPUNF, IFPUNW, IFPUNA,
          IIFPUND.IFPUNI.IFPUNR
           COMPLEX YNYQA
           IF(N) 10,10,20
           DO 15 1=1, NSTAR
     10
     15
           YF([]=A[MAG (YNYQA([])]
           RETURN
     20
           DO 25 I=1.NSTAR
     25
           YF(I)=REAL (YNYQA(I))
           RETURN
```

END

APPENDIX

Di

```
COMPILER OPTIONS - NAME = MAIN. OPT = 02. LINECNT = 57. SIZE = 0000K.
                    SOURCE, EBCDIC, NOLIST, DECK, LOAD, MAP, NOEDIT, ID, NOXREF
           SUBROUTINE UNSAVE(N)
           COMMON YNYQA (2048)
           COMMON NW.NR.XF(2048).YF(2048).NSTAR.NPOW.ND.JF.CT.ST.XO.YO.
          1 XS.YS.T.V.XNYQ
          2.NT.XZ.YZ.JD
           COMMON INTERP-IFFFT-IFFWAL-(FAUTO-IFGOF-IFPUNF-IFPUNW-IFPUNA-
           11FPUND.IFPUNI.IFPUNR
           COMPLEX YNYOA
           IF (N) 10.10.20
     10
           DO 15 I=1.NSTAR
           YNYQA(I)=CMPLX(0.0,YF(I))
     15
           RETURN
           D7 25 I=1, NSTAR
     20
           YNYQA([]=CMPLX(YF([],0.0)
     25
           RE TURN
```

```
COMPILER OPTIONS - NAME = MAIN, OPT=02, LINECNT=57, SIZE=0000K, SOURCE, EBCDIC, NOLIST, DECK, LOAD, MAP, NOEDIT, ID, NOXREF SUBROUTINE ROI
```

C ROTATE ZERO POINT OF THE TRACE
COMMON YNYQA(2048)
COMMON NW,NR,XF(2048),YF(2048),NSTAR,NPOW,ND,JF,CT,ST,XO,YO,
1 XS,YS,T,V,XNYQ
2,YT,XZ,YZ,JD

COMMON INTERP, IFFFT, IFFWAL, IFAUTO, IFGOF, IFPUNF, IFPUNW, IFPUNA, IIFPUND, IFPUNI, IFPUNR

COMPLEX YNYQA XZS=XZ-XO

YZS=YZ-YO XZ=CT\*XZS+ST\*YZS YZ=-ST\*XZS+CT\*YZS

C ROTATE AND SCALE INPUT ARRAY

D3 13 I=1,JF SWX=XF(I) - X3 SWY= YF(I) - Y0

YF(I) = (-ST\*SWX+CT\*SWY - YZ)\*V

XF(I)=(CT\*SWX+ST\*SWY-XZ)\*T RETURN

END

THE TA = A TAN ( TAN T )

```
COMPILER OPTIONS - NAME = MAIN. OPT = 02. LINECNT = 57. SIZE = 0000K.
                     SOURCE, EBCDIC, NOLIST, DECK, LOAD, MAP, NOEDIT, ID, NOXREF
            SUBROUTINE SCARTP
            COMMON YNYOA (2048)
            COMMON NW.NR.XF(2048).YF(2048).NSTAR.NPOW.ND.JF.CT.ST.XO.YO.
           1 XS.YS.T.V.XNYO
           2.NT.XZ.YZ.JD
            COMMON INTERP. IFFFT. IFFWAL. IFAUTO. IFGOF. IFPUNF. IFPUNW. IFPUNA.
           1 IFPUND, IFPUNI, IFPUNR
            COMPLEX YNYOA
            DIMENSION XT(4).YT(4)
            READINR. 21XZ.YZ
            READ(NR, 1) X1, X2, Y1, Y2
     1
            FORMAT (415)
           READ(NR, 2) (XT(I),YT(I),I=1,4)
     2
           FORMAT(2(1X,E11.4))
            READ(NR, 2) T.V
     3
           FORMAT(IX, 'THETA=', EL3.7/IX, 'COS(THETA)=', EL3.7/
           X 1X, 'SIN(THETA) = ', El3.7/1X, 'COORDINATE ORIGIN S=', El3.7/
           X 1X. COORDINATE ORIGIN Y = '.El3.7.1X. SCALE FACTOR X = '.El3.7/
           X IX. SCALE FACTOR Y = 1.E13.7/1
     4
           FORMATILX, 'XAXIS PTS X, Y COORDINATES'./.
           X 3(1X,E13.7),/,3(1X,E13.7))
           FORMATILX, 'YAXIS PTS X,Y COORDINATES',/,
           X 3(1X,E13,7),/,3(1X,E13,7))
            XO = ABS (X2 * XT(1) - X1 * XT(2))/ABS(Y1 - X2)
            YO = ABS (Y2 + YT(3) - Y1 + YT(4))/ABS(Y1 - Y2)
            XS=SXT((XT(2)-XT(1))+*2+(YT(2)-YT(1))+*2)/AS (X2-X1)
            YS = SQRT((XT(4) - XT(3)) + *2 + (YT(4) - YT(3)) * *2) / ABS (Y2 - Y1.)
            IF ( YT ( 2 ) . EQ . YT ( 1 ) ) GO TO 20
            TANT = (YT(2) - YT(1)) / (XT(2) - XT(1)) - (XT(4) - XT(3)) / (YT(4) - YT(3))) / 2.0
            CT=1.0/SQRT(1.0+TANT**2)
            ST=TANT+CT
```

```
GO TO 30

THETA=0.0

ST=0.0

CT=1.0

T=T/XS

V=V/YS

WRITE(NW,3) THETA,CT,ST,X0,Y0,XS,YS

WRITE(NW,4) X1,XT(1),YT(2),X2,XT(2),YT(2)

WRITE(NW,5) Y1,XT(3),YT(3),Y2,XT(4),YT(4)

RETURN
END
```

RE TURN END

```
COMPILER OPTIONS - NAME = MAIN.OPT=02.LINECNT=57.SIZE=0000K.
                    SOURCE, EBCDIC, NOLIST, DECK, LOAD, MAP, NOEDIT, ID, NOXREF
           SUBROUTINE SRIFUR
     C
     C
           THIS ROUTINE PERFORMS A BIT REVERSAL OF THE ORIGINAL NYQUIST ARRAY.
     C
           IF I IS THE INDEX OF AN ELEMENT OF THE ARRAY THEN I=SUM(A(N)*2**N)
           WHERE THE LIMITS OF THE SUM ARE O AND M AND A(N) IS O OR 1.
           THE ELEMENT SPECIFIED BY I IS EXCHANGED WITH THE ELEMENT SPECIFIED
     C
     C
           BY J. WHERE J=SUM(A(M-N) *2**N.
           EXAMPLE: SUPPOSE [=011010]. THEN J=1010110.
           NSTAR=2**NPDW
           COMMON YNYOA(2048)
           COMMON NW.NR.XF(2048).YF(2048).NSTAR.NPOW.ND.JF.CT.ST.X0.YO.
          1 XS. YS. T. V. XNYQ
          2.NT.XZ.YZ.JD
           COMMON INTERP. IFFFT. IFFWAL. IFAUTO. IFGOF. IFPUNF, IFPUNW. IFPUNA,
          1 IFPUND, IFPUNI, IFPUNR
           COMPLEX YNYQA, SWAP
           DIMENSION INT(15)
           DO 10 I=I.NPOW
     10
          INT([])=2**(NPOW-[])
           D7 30 1=1.NSTAR
           J = I - I
           ISUM=0
           DO 20 K= 1 . NP OW
           ISUM=ISUM + MOD(J,2)*INT(K)
     20
           J= 1/2
           ISUM=ISUM+L
           IF (ISUM-1)30.30.25
     25
           SWAP=YNYQA(I)
           YNYOA(I)=YNYOA(ISUM)
           YNYQA(ISUM) = SWAP
     30
           CONTINUE
```

```
COMPILER OPTIONS - NAME= MAIN, OPT=02, LINECNT=57, SIZE=0000K,
SOURCE, EBCDIC, NOLIST, DECK, LOAD, MAP, NOEDIT, ID, NOXREF
FUNCTION POLY(B, N, XVAL)

C
C EVALUATES A POLYNOMIAL FUNCTION
C
DIMENSION B(4)
POLY=0.0
XP=1.0
DO 10 I=1, N
POLY=POLY + B(I) * XP

10 XP=XP*XVAL
RETURN
END
```

	-	78 7 7	350	
ar i				
i-				
,				
				- 1
				7 (

## APPENDIX B. SIGNAL PROGRAM

```
( JAN 73 )
                                    OS/360 FORTRAN H
COMPILER OPTIONS - NAME MAIN. OPT = 02. LINECHT = 57. SIZE = 0000K.
                   SOURCE, EBCDIC, NOLIST, DECK, LOAD, MAP, NOEDIT, ID, NOXREF
     C
     C MAIN PROGRAM FOR A GENERAL NON-LINEAR LEAST SQUARES FIT OF A
     C PARAMETRIZED FUNCTION TO A SET OF DATA Y
           NON LINEAR LEAST SQUARES FITTING PROGRAM
          *********COMMON INFORMATION*******************
                   MATRIX OF EQN. 8 AND INVERSE AFTER INVERSION
                  VECTOR IN EQN. 8
           DPR
                CHANGES IN THE RESTRICTED PARAMETERS
           FMT
                  FORMAT FOR THE EXPERIMENTAL DATA
                  RESTRICTED PARAMETERS
           PZ
           PSAVR ARRAY FOR TEMPORARY SAVING OF PR
           Ζ .
                   ARRAY OF DERIVATIVES
           PHI
                  VALUE OF THEORETICAL FUNCTION
           RESID RESIDUALS
               ARRAY OF INDEPENDENT VARIABLES SUCH AS TIME
                  THE EXPERIMENTAL DATA
                  THE STATISTICAL WEIGHTS
           IFCOR DETERMINES DISPOSITION OF CORRELATION COEFFICIENTS
           IFPLOT DETERMINES IF THE PLOTTING PROGRAM IS CALLED
           INDEZ
                   INTERNAL PARAMETER COMMUNICATING BETWEEN MAIN AND TESTS
                   THE INFORMATION THAT THE PROGRAM HAS OR HAS NOT
                   CONVERGED
```

```
C
      N3
              THE NUMBER OF THE BEGINNING PARAMETER IN A GIVEN
              REGION OF PARAMETER SPACE
      NBADPT
              THE NUMBER OF BAD DATA POINTS
      NCHAN
              THE NUMBER OF CHANNELS
      NE
              THE NUMBER OF THE LAST PARAMETER IN A GIVEN REGION OF
              PARAMETER SPACE
      NP
              NUMBER OF PARAMETERS TO BE FIT IN A GIVEN REGION OF
              PARAMETER SPACE
      NPSTEP
              USED FOR PLOTTING
      VP
              THE NUMBER OF PARAMETERS TO BE FIT IN A GINEN REGION
              OF PARAMETER SPACE
      NRP
              THE NUMBER OF RESTRICTED PARAMETERS
      AA
              ALPHA IN EQN. 12
      IFD
              CHI-SQUARED
      CHIB
              CHI-SQUARED BEFORE EACH ITERATION
      CHIA
              CHI-SQUARED AFTER EACH ITERATION
      DEGE
              THE NUMBER OF DEGREES OF FREEDOM
C
C
     I/O INFORMATION
C * * *
     NREAD IS THE SYMBOLIC DESIGNATION FOR THE CARD READER
[本本本
     NARITE IS THE SYMBOLIC DESIGNATION FOR THE LINE PRINTER
     NPUNCH IS THE SYMBOLIC DESIGNATION FOR THE CARD PUNCH
C***
C ***
     THE I/O CHANNELS ARE DEFINED IN THE NEXT THREE STATEMENTS
      COMMON A(50,50),C(50),DPR(50),FMT(5),PR(50),PSAVR(50),Z(50),P(50)
      COMMON PHI(1024), RESID(1024), X(1024), Y(1024), W(1024)
      COMMON IFCOR, IFPLOT, INDEZ. NB. NBADPT. NCHAN. NE. NP. NPSTEP. NRP.
             NREAD NWRITE NPUNCH
      COMMON AA, CHI, CHIB, CHIA, DEGF
      NREAD=5
      NWRITE=6
      NPUNCH=7
```

```
APPENDIX I
```

```
C GET THE MAXIMUM # OF ITERATIONS AND THE INPUT DATA FORMAT
      READ(NREAD, 1) JSTOP, FMT
C GET THE INPUT DATA FOR THE N-TH FIT
10
    CALL READIN
C INITIALIZE ITERATION COUNTER
      D=TIUGL
      WRITE(NWRITE, 2)
C INCREMENT ITERATION COUNTER
C
20
    JQUIT=JQUIT+1
C TEST # OF ITERATIONS
      IF ( JOUIT . LE . JSTOP 1 GO TO 30
      WRITE(NWRITE,3)
  GO TO 40
C SET UP MATRIX OF DERIVATES AND SOLVE FOR DELTA PARAMETERS
30
   CALL GRIND(O)
C TEST AND MAKE CHANGES IN THE PARAMETERS
      CALL TESTS
C INDEZ=O IF THE FIT HAS CONVERGED
```

```
IF (INDEZ) 20,40,40
C WRITE DUT THE PARAMETERS AND THE THEORETICAL FIT
40
      CALL SCRIBE
C TEST IF A PLOT IS DESIRED
      IF(IFPLOT)60,70,60
C DO THE PLOTTING
60
     CALL PLOTS
C DD THE ERROR ANALYSIS
70
      CALL ERRMAT
C GO GET THE DATA FOR THE NEXT FIT
C
      IF(1)80,80,10
      FORMAT (15,544)
      FORMATILH- , 'CHANGES IN PARAMETERS')
                     THE NUMBER OF ITERATIONS EXCEEDED JSTOP. THE'./;
             ' PROGRAM WILL AUTOMATICALLY GO TO THE PLOTTING AND ',/,
             · ERROR ANALYSIS ROUTINES IN WHICH CAAE THE ANSWERS DO 1,/,
             ' NOT REPRESENT THE TRUE CHI-SQUARE SOLUTUON')
      STOP
80
      END
```

C

```
DEFENDED D
```

```
( JAN 73 )
                                       OS/360 FORTRAN H
COMPILER OPTIONS - NAME MAIN, OPT=02, LINECHT=57, SIZE=0000K,
                    SOURCE, EBCDIC, NOLIST, DECK, LOAD, MAP, NOEDII, ID, NOXREF
            SUBROUTINE LSCPHI
     C
     C THIS ROUTINE CALCULATES THE THEORETICAL FUNCTION, RESIDUALS,
     C AND CHI-SQUARE. THE USER MUST SUPPLYTHE FUNCTION PHIENC
     C
           COMMON A(50,50),C(50),DPR(50),FMT(5),PR(50),PSAVR(50),Z(50),P(50)
           COMMON PHI (1024), RESID(1024), X(1024), Y(1024), W(1024)
           COMMON IFCOR, IFPLOT, INDEZ, NB, NBADPT, NCHAN, NE, NP, NPSTEP, NRP,
                 NREAD . NWRITE . NPUNCH
           COMMON AA, CHI, CHIB, CHIA, DEGF
     C CALCULATE CHI, PHI, AND RESID
           CHI=0.0
           DD 10 I=1.NCHAN
           PHI(I)=PHIFNC(I)
           RESID(I) = (Y(I) - PHI(I)) * W(I)
           CHI=CHI+RESID(I)**2
     10
           CONTINUE
     C DIVIDE CHI BY THE # OF DEGREES OF FREEDOM
           CHI=CHI/DEGF
           RETURN
           END
```

```
COMPILER OPTIONS - NAME MAIN, OPT=02, LINECNT=57, SIZE=0000K,
                    SOURCE, EBCDIC, NOLIST, DECK, LDAD, MAP, NOEDIT, ID, NOXREF
           SJBROUTINE ELTS(I)
     C
     C THIS ROUTINE CALCULATES THE DERIVATIVES OF PHI(I).
     C THE USER MUST SUPPLY THE FUNCTION DPHILL.JI
           COMMON A(50,50),C(50),DPR(50),FMT(5),PR(50),PSAVR(50),Z(50),P(50)
           COMMON PHI(1024), RESID(1024), X(1024), Y(1024), W(1024)
           COMMON IFCOR, IFPLOT, INDEZ, NB, NBADPT, NCHAN, NE, NP, NPSTEP, NRP,
                   VREAD, NWRITE, NPUNCH
           COMMON AA, CHI, CHIB, CHIA, DEGF
           DO 10 J=NB, NE
           Z(J) = DPHI(I,J)
     10
           CONTINUE
           RETURN
           END
```

NB = L NE = NRP

```
COMPILER OPTIONS - NAME MAIN. OPT=02. LINECHT=57. SIZE=0000K.
                    SOURCE, EBCDIC, NOLIST, DECK, LOAD, MAP, NOEDIT, ID, NOXREF
           SUBROUTINE READIN
     C
     C*** THIS ROUTINE READS IN ALL THE INPUT PARAMETERS, DATA, AND SETS UP
     C *** INTERNAL CONTROL PARAMETERS.
     C
           COMMON A(50.50),C(50),DPR(50),FMT(5),PR(50),PSAVR(50),Z(50),P(50)
           COMMON PHI(1024), RESID(1024), X(1024), Y(1024), W(1024)
           COMMON IFCOR, IFPLOT, INDEZ, NB, NBADPT, NCHAN, NE, NP, NPSTEP, NRP,
                   NREAD, NWRITE, NPUNCH
           COMMON AA, CHI, CHIB, CHIA, DEGF
           DIMENSION NBAD(1024), YRUN(1024), TITLE(20)
           EQUIVALENCE (X(1), YRUN(1), NBAD(1)), (W(1), TITLE(1))
           WRITE(NWRITE, 1001)
           READ(NREAD, 1) TITLE
           WRITE(NWRITE, 2) TITLE
     C
     Caax
           READ AND WRITE THE CONTROL PARAMETERS.
     C
           READ(NREAD.3) NCHAN.NBADPT.NRUNS.NRP.NPSTEP.IFPLOT.IFCOR
           WRITE(NWRITE, 1002)
           WRITE(NWRITE, 4) NCHAN, NBADPT, NRUNS, NRP, NPSTEP, IFPLOT, IFCOR
     C
     C***
           CALCULATE THE INTERNAL CONTROL PARAMETERS.
     C
           IF (NPSTEP.EQ.C) NPSTEP=1
           NPARAM=NRP
           NP = NRP
```

C

[ 本本本

190

WRITE(NWRITE, 1006)

```
133
```

```
APPENDIX B
```

```
C*** READ AND WRITE CHANNEL # OF BAD POINTS.
      READ(NREAD, 9) (NBAD(I), I=1, NBADPT)
      WRITE(NWRITE.9)(NBAD(I), I=1, NBADPT)
C*** DISCARD BAD DATA POINTS.
      DO 200 I=1.NBADPT
             J=NBAD(1)
             C . C = (L)Y
             W(J) = 0.0
200
             CONTINUE
C*** WRITE DATA.
210
    WRITE(NWRITE, 1007)
      WRITE(NWRITE, 12)(Y(I), I=1, NCHAN)
      (8CC1, STIRWN) STIRW
      WRITE(NWRITE, 12) (X(I), I=1, NCHAN)
      WRITEINWRITE, 10091
      WRITEINWRITE, 121 (WIII, I=1, NCHAN)
      WRITE (NWRITE, LOOL)
C
C***
      CALCULATE THE # OF DEGREES OF FREEDOM FOR THE FIT.
C
      DEGF=NCHAN-NBADPT-NRP
C
C * * *
      CALCULATE THE INITIAL THEORETICAL FUNCTION AND CHI-SQUARED.
      CALL LSQPHI
      CHIB=CHI
C
C*** SET CHI AFTER AN ITERATION TO 100.0 IN CASE PROGRAM CAN MAKE NO
C*** SUCCESSFUL ITERATIONS ..
C
      CHIA=100.0
1001 FORMAT (1H1,/,1H-)
     FORMAT (2044)
```

```
FORMAT (20X, 20A4)
3
      FORMAT(715)
                                              NRP NPSTEP IFPLOT
1002 FORMAT (1H- NCHAN NBADPT
                                     NRUNS
     1'IFCOR')
      FORMAT(1x, 718)
      FORMAT (8F10.4)
1003 FORMAT (1H-'PARAMETER NUMBER', 5X, 'PARAMETER')
      FORMAT (14X, 13, 4X, F10.2)
      FORMAT (4(15,F10.5))
      FORMAT (1415)
      FORMAT ( '-CHANNEL NUMBER OF THE BAD DATA POINTS')
1006
      FORMAT (1H1, /, 1H-, 30x, 'EXPERIMENTAL DATA', //)
1007
      FORMAT(LH1, 30x, 'X VALUES FO THE EXPERIMENTAL DATA')
1008
      FORMAT(1H1,30x, W VALUES FO THE EXPERIMENTAL DATA')
1009
      FJRMAT(1)(1X,E12.6))
12
      RETURN
      END
```

SOURCE. EBCDIC, NOLIST, DECK, LOAD, MAP, NOEDIT, ID, NOXREF

COMMON A(50.50).C(50).DPR(50).FMT(5).PR(50).PSAVR(50).Z(50).P(50)

C\*\*\* THIS ROUTINE DOES THE TESTING OF THE CHANGES IN THE PARAMETERS

COMMON IFCOR . IFPLOT . INDEZ . NB . NBADPT . NCHAN . NE . NP . NPSTEP . NRP .

COMMON PHI (1024) . RESID(1024) . X (1024) . Y (1024) . W (1024)

COMPILER OPTIONS - NAME MAIN. OPT=02. LINECHT=57. SIZE=0000K.

SUBROUTINE TESTS

```
NREAD . NWRITE . NPUNCH
      COMMON AA.CHI.CHIB.CHIA.DEGF
      DIMENSION WED(4).ZED(4)
C*** AND PERFORMS THE NECESSARY CHANGES.
C*** INITIALIZE FOR SUMMING.
      DT = 0.0
C*** SUM AND SAVE THE PARAMETERS.
      DO 10 I=NB.NE
            DI = DI + DPR(I) * C(I)
            PSAVR(I)=PR(I)
10
            CONTINUE
C*** TEST FOR DT < ZERO.
      IF(DT)20.40.40
C*** CHANGE THE SIGN OF THE DELTA P'S IF DT< ZERO.
20
      DD 30 I=NB.NE
            DPR(I) = -DPR(I)
30
            CONTINUE
      DT = -DT
C*** USE THE FULL LENGTH OF THE CHANGES VECTOR TO FIND A NEW CHI.
40
      AA = 1.0
C*** FIND THE NEW CHI.
      CALL CALS2
C*** TEST FOR NEW CHI SMALLER THAN OLD. IF NOT GO TO STATEMENT 90.
      IF(CHIB.LE.CHI) GO TO 90
C*** TEST FOR THE CHANGE IN CHI LESS THAN TERMINATION VALUE.
```

```
50
      IF((CHIB-CHI).GT.O.1E-08) GO TO 80
C*** REACHES HERE IF TERMINATED.
      INDEZ=0
     SET CHI AFTER ITERATION EQUAL TO CHI.
C***
60
      CHIA=CHI
C***
      WRITE OUT CHANGES IN PARAMETERS, LENGTH OF VECTOR, CHI BEFORE,
     AND CHI AFTER.
      WRITE(NWRITE, 1)(I, DPR(I), I=NB, NE)
      WRITE(NWRITE.2)AA.CHIB.CHIA
C*** SET CHI BEFORE TO CHI AFTER.
70
      CHIB=CHIA
      RETURN
     REACHES HERE IF NOT CONVERGED YET AND GOES BACK TO WRITE OUT.
      INDEZ=-1
80
      GO TO 60
     REACHES HERE IF FULL LENGTH OF VECTOR WILL NOT LOWER CHI AND
C*** TRIES THE PARABOLAS DESCRIBED IN WRITE UP.
90
      S=CHI
      AA=0.5
      CALL CALS2
      ZED(1)=CHI
      WE'D(1) = AA
      N=1
      K = 0
      AA=DT/(S-CHIB+2.0*DT)
133
      K = K + 1
     IF THE FACTOR MULTIPLYING THE CHANGES VECTOR IS TOO SMALL SKIP IT
     FOR THE CHI AFTER CALCULATION.
      IF (AA.LT.O.1E-01)
                         GO TO 110
      CALL CALS2
      N=N+1
```

```
ZED(N) =CHI
      WED(N) = AA
110
      IF(K-2)120,130,150
120
      AA=1.0/DT
      60 TO 100
130
      ZED1=CHIB+S-2.0*ZED(1)
      IF (ZED1)140, 150, 140
140
      AA=(S+3.0*CHIB-4.0*ZED(1))/(4.0*ZED1)
      GD TO 100
      AA = 0.5
150
      CHI=ZED(1)
      DO 160 I=1.N
             IFICHI.GT.ZED(II) GO TO 160
             CHI=ZED(I)
             AA=WED(I)
160
             CONTINUE
C*** TEST THE SMALLEST VALUE GIVEN BY THE PARABOLAS TO SEE IF IT IS
      SMALLER THAN CHI BEFORE, IF NOT RESET PARAMETERS AND GO TO NEXT
C***
C*** SUBSPACE.
      IF (CHIB. LE. CHI)
                        GO TO 170
      CALL CALS2
      GO TO 50
C*** RESET PARAMETERS.
170
      DO 180 I=NB, NE
             PR(I)=PSAVR(I)
180
             CONTINUE
      AA = 0 . 0
      CALL CALS2
      INDEZ=0
```

WRITE(NWRITE, 1001)

WRITE(NWRITE, 1)(I, DPR(I), I=NB, NE)

190

```
FURMAT (6(4H DP(,I2,2H)=,E12.5,2K))

FURMAT (4H AA=,F14.7,5X,5HCHIB=,E14.7,3X,5HCHIA=,E14.7,7)

1001 FURMAT ('-THE FOLLOWING DPR(I) PRODUCED A DIVERGENT STEP WHICH ',

COULD NOT BE FIXED')

RETURN
END
```

CONTINUE

```
COMPILER OPTIONS - NAME MAIN, OPT=02, LINECHT=57, SIZE=0000K.
                    SOURCE . EBCDIC . NOLIST . DECK . LOAD . MAP . NOED IT . ID . NOXREF
           SUBROUTINE GRINDLINV)
     C
     C*** THIS ROUTINE SETS UP THE MATRIX OF DERIVATIVES. THE CONSTANTS IN
     C*** THE NORMAL EQUATIONS, THEN CALLS THE MATRIX INVERTER, AND FINALLY
     C*** SHIFTS THE CONSTANTS AND THE SOLUTION TO MATCH THE PARAMETERS.
     C
           COMMON A(50,50),C(50),DPR(50),FMT(5),PR(50),PSAVR(50),Z(50),P(50)
           CJMMJN PHI(1024), RESID(1024), X(1024), Y(1024), W(1024)
           COMMON IFCOR, IFPLOT, INDEZ, NB, NBADPT, NCHAN, NE, NP, NPSTEP, NRP,
                  NREAD . NWRITE . NPUNCH
           COMMON AA, CHI, CHIB, CHIA, DEGF
           DIMENSION B(50)
           EQUIVALENCE (DPR(1).8(1))
     C*** INITIALIZE FOR SUMMING.
           DO 20 I=1.NP
                 DO 10 J=I.NP
                        A(I,J) = 0.0
     10
                        CONTINUE
                 B(1)=0.0
                 CONTINUE
     20
           SET UP MATRIX AND CONSTANTS.
           D3. 70 E=1.NCHAN
     C***
                 TEST FOR BAD DATA POINTS.
                 IF(Y(I))40.70.40
     40
                 CALL ELTS(I)
                 DO 60 J=1.NP'
                        DC 50 K= J. NP
                              A(J,K)=A(J,K)+Z(J)*Z(K)
     50
                              CONTINUE
                        B(J)=B(J)+Z(J)*RESID(I)
     60
                        CONTINUE
```

```
C*** GET LOWER HALF OF MATRIX (THE MATRIX IS SYMETERIC).
      DJ 93 J=1, NP
            00 80 K= J, NP
                  A(K,J) = A(J,K)
80
                  CONTINUE
C
                  SAVE THE CONSTANTS.
            (L)8=(L)3
90
            CONTINUE
     CALL THE MATRIX INVERTER.
      CALL INVMAT (NP, INV, ZZZ)
120
    RETURN
      END
```

```
COMPILER OPTIONS - NAME = MAIN.OPT=02.LINECNT=57.SIZE=0000K.
                   SOURCE . EBCDIC . NOLIST . DECK . LOAD . MAP . NOEDIT . ID . NOXREF
           SUBROUTINE INVMAT (N.INV.PIVOT)
          COMMON A(50,50),C(50),DPR(50),FMT(5),PR(50),PSAVR(50),Z(50),P(50)
           COMMON PHI (1024) • RESID(1024) • X(1024) • Y(1024) • W(1024)
           COMMON IFCOR . IF PLOT . INDEZ . NB . NBADPT . NCHAN . NE . NP . NPSTEP . NRP .
                  NREAD . NWRITE . NPUNCH
           COMMON AA. CHI. CHIB. CHIA. DEGF
           DIMENSION B(50) . INDC(50) . INDR(50) . IPIVOT(50)
          EQUIVALENCE (JROW. IROW), (JCOLUM. ICOLUM), (AMAX. SWAP.T)
           EQUIVALENCE (DPR(1).B(1))
     C*** MATRIX INVERSION WITH ACCOMPANYING SOLUTION OF LINEAR EQUATIONS.
     C******** NOTE. INV=1 IF THE INVERSE IS DESIRED.
     C *** INITIALIZATION
           00 10 I=1.N
                 IPIVGT(I)=0
     10
                 CONTINUE
           DO 100 I=1.N
     [ ***
                  SEARCH FOR PIVOT ELEMENT.
                  AMA X=0.0
                  DO 40 J=1.N
                        IF(IPIVOT(J).EQ.1) GO TO 40
                        DO 30 K=1.N
                              IF(IPIVOT(K)-1)20,30,130
     20
                              IF (ABS (AMAX).GT.ABS (A(J.K))) GO TO 30
                              IROW=J
                              I COLUM=K
                              AMAX=A(J.K)
     30
                              CONTINUE
     40
                        CONTINUE
                  IPIVOT(ICOLUM) = IPIVOT(ICOLUM) + 1
    C***
                  INTERCHANGE ROWS TO PUT PIVOT ELEMENT ON DIAGONAL.
                  IF (IROW.EQ.ICOLUM) GO TO 60
```

CONTINUE

INDR(I)=IROW INDC(I)=ICOLUM

IF(INV.NE.1) GO TO 100

DO 50 J=1.N

SWAP=B(IROW)

SWAP=A(IROW, J)

CONTINUE

PIVOT=A(ICOLUM, ICOLUM)

B(IROW)=B(ICOLUM) B ( ICOLUM) = SWAP

A(ICOLUM, J) = SWAP

A(IROW, J) = A(ICOLUM, J)

B(J) = B(J) - B(ICOLUM) \*T

SET INDEX IF INVERSE IS DESIRED.

50

63

90

C \* \* \*

```
100
              CONTINUE
C * * *
     INTERCHANGE COLUMNS IF INVERSE IS DESIRED.
      IF(INV.NE.1) GO TO 130
      DO 120 I=1,N
              J = N + L - I
              IF(INDR(J).EQ.INDC(J)) GO TO 120
              JROW=[NDR(J)
              JCOLUM=[NDC(J)
              DO 110 K=1,N
                     SWAP=A(K, JROW)
                     A(K, JROW) = A(K, JCOLUM)
                     A(K, JCOLUM) = SWAP
                     CONTINUE
110
              CONTINUE
120
130
      RETURN
      END
```

```
COMPILER OPTIONS - NAME - MAIN-OPT=02.LINECNT=57.SIZE=0000K.
                    SOURCE, EBCDIC, NOLIST, DECK, LOAD, MAP, NOEDIT, ID, NOXREF
           SUBROUTINE CALS2
     C
     C 本本本
          THIS ROUTINE IS CALLED BY TESTS TO SET UP THE PARAMETERS
          AFTER THE PERAMETERS HAVE BEEN CHANGED. AND CALCULATE CHI BY
     (本本本本
           BY CALLING ESCPHI
     C.
           COMMON A(50.50).C(50).DPR(50).EMT(5).PR(50).PSAVR(50).Z(50).P(50)
           COMMON PHI (1024) - RESID(1024) - X (1024) - Y (1024) - W (1024)
           COMMON IFCOR.IFPLOT.INDEZ.NB.NBADPT.NCHAN.NE.NP.NPSTEP.NRP.
                  WREAD . NWRITE . NPUNCH
           COMMON AA, CHI, CHIB, CHIA, DEGF
     C*** FIND THE CHANGED RESTRICTED PARAMETERS.
           00 10 I=NB.NE
                  PR(I)=PSAVR(I)+AA*DPR(I)
           PII)=PR(I)
     10
                 CONTINUE
     C
          FIND THE CHI.
     [ 辛辛辛
     C
     60
           CALL LSQPHI
           RETURN
           FND
```

```
COMPILER OPTIONS - NAME - MAIN. OPT=02. LINECHT=57. SIZE=0000K.
                    SOURCE, EBCDIC, NOLIST, DECK, LOAD, MAP, NOEDIT, ID, NOXREF
            SUBROUTINE SCRIBE
     C
     C***
           THIS ROUTINE DOES THE WRITING OUT.
           COMMON A(50,50),C(50),DPR(50),FMT(5),PR(50),PSAVR(50),Z(50),P(50)
           COMMON PHI(1024), RESID(1024), X(1024), Y(1024), W(1024)
           COMMON IFCOR, IFPLOT, INDEZ, NB, NBADPT, NCHAN, NE, NP, NPSTEP, NRP,
                   WREAD, NWRITE, NPUNCH
           COMMON AA, CHI, CHIB, CHIA, DEGF
           DIMENSION PHIN(512), YN(512), XVEL(512)
            WRITE(NWRITE, 1)
           WRITE(NWRITE, 2)(I, PR(I), I=1, NRP)
           WRITE(NWRITE, 3)
     20
           WRITE(NWRITE, 4) (PHI(I), I=1, NCHAN)
     1
           FORMAT ( '-PARAMETER NUMBER', 10X, 'PARAMETER')
           FORMAT (14X,13,5X,E14.7)
     3
           FORMAT (1H1,/,1H-,30X, THEORETICAL FIT',//)
           FORMAT(8(2X, E13, 7))
     30
           RETURN
           END
```

```
COMPILER OPTIONS - NAME = MAIN.OPT=02.LINECNT=57.SIZE=0000K.
                   SOURCE, EBCDIC, NOLIST, DECK, LOAD, MAP, NOEDIT, ID, NOXREF
           SUBROUTINE FRRMAT
    C
     C***
           THIS ROUTINE DOES THE ERROR ANALYSIS.
           COMMON A(50.50; C(50), DPR(50), FMT(5), PR(50), PSAVR(50), Z(50), P(50)
           COMMON PHI(1024), RESID(1024), X(1024), Y(1024), W(1024)
           COMMON IFCOR. IFPLOT. INDEZ. NB. NBADPT. NCHAN. NE. NP. NPSTEP. NRP.
                  NREAD . NWRITE . NPUNCH
           COMMON AA, CHI, CHIB, CHIA, DEGF
           DIMENSION ACSRS(50).AMS(50).FINT(50).SIGI(50)
     C*** TEST FOR NO SUCCESSFUL ITERATIONS.
           IF (CHIA. EQ. 100.0) GO TO 160
           PIE=3.1415927
     C*** INITIALIZE FOR FINDING THE DATA POINT WITH MAXIMUM WEIGHTED
     C*** RESIDUAL.
           TAX=0.0
     C*** FIND CHANNEL # OF MAXIMUM WEIGHTED RESIDUAL AND WRITE IT OUT.
           DD 10 I=1.NCHAN
                 Q=ABS(RESID(I))
                 IF (Q.LT.TAX) GO TO 10
                 ITRACK=I
                 TAX=0
                 CONTINUE
     10
           WRITE(NWRITE, 1) ITRACK
    C*** FIND THE PROBABILITY FOR THE FITTED CHI.
           AR = SORT(2.0*CH[A*DEGF]-SORT(2.0*DEGF-1.0)
           PROB=ERFC(AR)/2.0
           NC H=NC HAN-NB ADPT
    C*** WRITE OUT STATISTICAL INFORMATION.
```

C\*\*\* TEST CHI FOR LESS THAN 1.0 AND SET TO ONE IS SATISFIED.

WRITE(NWRITE, 2)NCH, NRP, CHIA, PROB

IFICHIA.LT.1.3) CHIA=1.3

```
APPENDIX E
```

```
C*** CALL GRIND TO FIND THE ERROR MATRIX.
20
      CALL GRIND(1)
C*** FIND THE STANDARD DEVIATIONS ON THE FITTED PARAMETERS.
      DJ 30 I=1.NRP
             ACSRS(I)=SQRT(A(I,I)*CHIA)
30
            CONTINUE
      FIND THE CORRELATION COEFFICIENTS.
[ 本本本
      II = NRP - I
      DO 50 I=1.II
             JJ=[+]
             SIG=ACSRS(I)
             DO 40 J=JJ, NRP
                   A(I,J)=A(I,J)*CH(A/(SIG*ACSRS(J))
40
                   CONTINUE
50
             CONTINUE
      WRITE OUT THE PARAMETERS AND THEIR STANDARD DEVIATIONS.
      WRITE(NWRITE.3)
      WRITE(NWRITE, 4) (I, PR(I), ACSRS(I), I=1, NRP)
C*** TEST FOR CORRELATION COFFFICIENTS DESIRED.
100
      IF (IFCOR) 110, 150, 120
C*** IF IFCOR < ZERO PUNCH OUT PARAMETERS.
110
      WRITE(NPUNCH.11)[PR(I).ACSRS(I).I=1.NRP)
120
      WRITE(NWRITE, 12)
C*** WRITE OUT CORRELATION COEFFICIENTS.
      II = NRP - 1
      DO 140 [=1.1]
              WRITE (NWRITE, 13) (I, J, A(I, J), J=JJ, NRP)
C***
              TEST FOR PUNCHING.
              IF(IFCOR)130,140,140
130
              WRITE (NPUNCH, 14) (A(I, J), J=JJ, NRP)
```

```
140
              CONTINUE
150
      RETURN
      WRITE(NWRITE, 15)
160
      FORMAT ( '-CHANNEL NUMBER OF THE WORST DATA POINT IS ', 15)
      FORMAT ('-CHI-SQUARED FOR', 15. CHANNELS AND', 15. PARAMETERS IS',
              F11.7./. THE PROBABILITY FOR WHICH IS !. E16.7)
3
      FORMAT (1H1,/, "-PARAMETER NUMBER", 10X, "PARAMETER", 10X, "STANDARD ",
               'DEVIATION')
      FORMAT (14X, [3, 5X, E14, 7, 14X, E14, 7)
      FORMAT (1X, E14.7, 15X, E14.7)
      FORMAT (5E16.7)
11
      FORMAT (1H1, /, 1H-, 30X, 'CORRELATION COEFFICENTS')
12
13
      FORMAT (1H-,6(3H A(,12,1H,,12,2H)=,F10.7))
14
      FORMAT (8F10.5)
15
      FORMAT (' PROGRAM COULD MAKE NO SUCCESSFUL ITERATIONS')
      RETURN
      END
```

```
SOURCE. EBCDIC, NOLIST, DECK, LOAD, MAP, NOEDIT, ID, NOXREF
      SUBROUTINE PLOTS
C
C***
      THIS ROUTINE DOES THE PLOTTING OF THE DATA AND THEORETICAL FIT.
C
      COMMON A(50.50),C(50),DPR(50),FMT(5),PR(50),PSAVR(50),Z(50),P(50)
      COMMON PHI (1024) . RESID(1024) . X (1024) . Y (1024) . W (1024)
      COMMON IFCOR IFPLOT INDEZ NB NBADPT NCHAN NE NP NPSTEP NRP .
             NREAD . NWRITE . NPUNCH
      COMMON AA.CHI.CHIB.CHIA.DEGF
      DIMENSION ARR(126)
      WRITE(NWRITE.1)
      DATA BLANK, STAR, ZERO, PLUS, BAD/1H , 1H*, 1HO, 1H+, 1HB/
C*** INITIALIZE FOR FINDING MAXIMUM AND MINIMUM.
      AMAX=PHI(1)
      AMIN=AMAX .
C*** INITIALIZE THE PLOTTING ARRAY TO BLANKS.
      DO 10 I=1.126
            ARR(I)=BLANK
10
            CONTINUE
C*** FIND MAXIMUM AND MINIMUM OF EITHER THE DATA OR THE FIT.
      DO 30 I=1.NCHAN.NPSTEP
            IF (Y(I). EQ. J. J) GO TO 23
            IF(Y(I),GT,AMAX) AMAX=Y(I)
            IF(Y(I) LT AMIN) AMIN=Y(I)
20
            IF(PHI(I).GT.AMAX) AMAX=PHI(I)
            IF(PHI(I).LT.AMIN) AMIN=PHI(I)
30
            CONTINUE
C*** FIND THE BIN SIZE.
      DBIN=(AMAX-AMIN)/124.0
C*** FIND WHERE TO PUT THE CURVES.
      DO 80 I=1.NCHAN.NPSTEP
            II=IFIX((PHI(I)-AM(N)/DBIN)+2
```

COMPILER OPTIONS - NAME - MAIN-OPT=02.LINECNT=57.SIZE=0000K.

```
IF (Y(1) . NE . 0 . 0) GO TO 43
             JJ=1
            ARR(1)=BAD
             GO TO 60
40
            JJ=IFIX((Y(I)-AMIN)/DBIN)+2
            IF(II.NE.JJ) GO TO 50
            ARR(III) = PLUS
            GO FO 70
50
            ARR(JJ) = ZERO
50
            ARR(III)=STAR
70
            WRITE(NWRITE, 2) I, (ARR(K), K=1, 126)
            ARR(II)=BLANK
            ARR(JJ) = BLANK
80
            CONTINUE
      WRITE(NWRITE, 3)
1
      FORMAT (1H1)
2
      FJRMAT (1X, [4, 2X, 126A1)
      FORMAT ('-0=EXPERIMENTAL POINT, *=FITTED POINT, +=COINCIDENT ',
3
              'EXPERIMENTAL AND FITTED POINTS, B=BAD EXPERIMENTAL ..
              'POINT')
      RETURN
      END
```

SOURCE . EBCDIC . NOLIST . DECK . LOAD . MAP . NOEDIT . ID . NO XREF

COMMON A(50.50).C(50).DPR(50).FMT(5).PR(50).PSAVR(50).Z(50).P(50)

COMMON IFCOR, IFPLOT, INDEZ, NB, NBADPT, NCHAN, NE, NP, NPSTEP, NRP,

C THIS SUBROUTINE CALCULATES THE THEORETICAL FUNCTION FOR THE FITTING
C PROCESS - SPECIFICALLY PHI =SUN(P(K) + EXP(-P(K+1) + X(I)) + COS(P(K+2) + X(I))

COMMON PHI (1024) - RESID(1024) - X(1024) - Y(1024) - W(1024)

COMPILER OPTIONS - NAME = MAIN. OPT=02. LINECNT=57. SIZE=0000K.

C WHERE THE SUM IS OVER THE TOTAL NUMBER OF FUNCTIONS

PHIENC=PHIENC+P(J)\*EXP(-PSX(J+1))\*COS(PSX(J+2))

WREAD . NWRITE . NPUNCH

COMMON AA.CHI.CHIB.CHIA.DEGF

FUNCTION PHIENCILL

DIMENSION PSX(50)

(I) X = I X

CONTINUE

RETURN

PHIFNC=0.0 DO 20 J=1.NP.3

DO 10 K=1.NP

PSX(K) = P(K) \* XI

C

10

```
151
```

```
APPENDIX B
```

152

```
( JAN 73 )
```

RF TURN

# OS/360 FORTRAN H

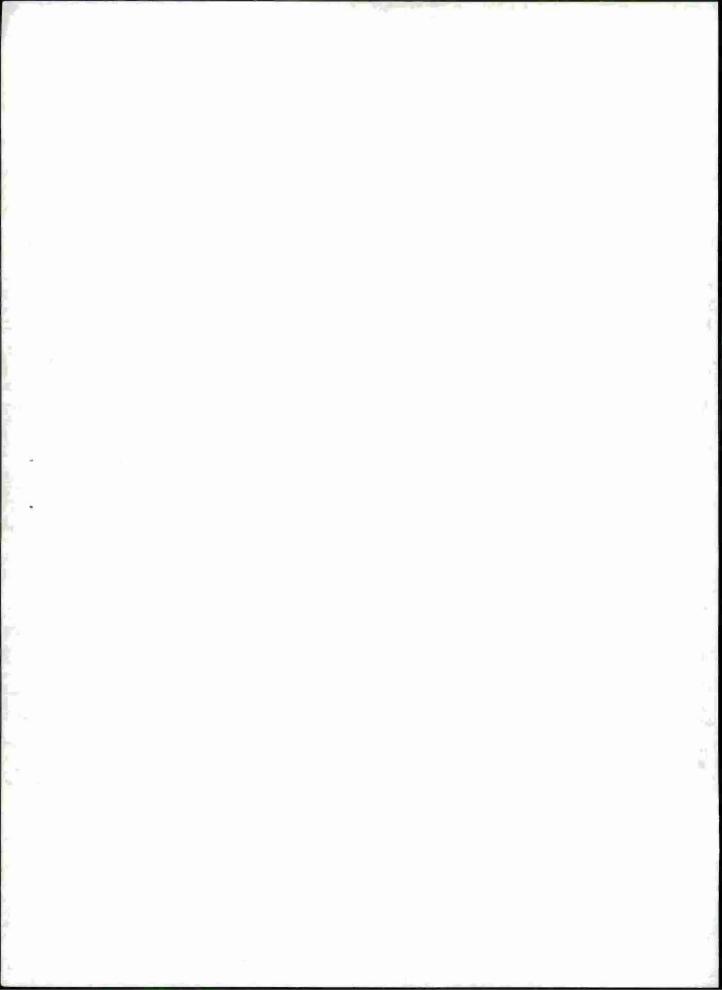
```
COMPILER OPTIONS - NAME MAIN, OPT=02. LINECHT=57, SIZE=0000K,
                     SOURCE, EBCDIC, NOLIST, DECK, LOAD, MAP, NOEDIT, ID, NO XREF
            FUNCTION DPHI(I.J)
           COMMON A(50,50),C(50),DPR(50),FMT(5),PR(50),PSAVR(50),Z(50),P(50)
            CJMMON PHI(1024), RESID(1024), X(1024), Y(1024), W(1024)
           COMMON IFCOR, IFPLOT, INDEZ, NB, NBADPT, NCHAN, NE, NP, NPSTEP, NRP,
                   NREAD, NWRITE, NPUNCH
            COMMON AA, CHI, CHIB, CHIA, DEGF
           DIMENSION PSX(50)
            DIMENSION SAVS(50) SAVC(50)
     C
     C THIS ROUTINE CALCULATES THE DERIVATIVES OF THE THEORETICAL FUNCTION
     C WHERE THE FUNCTION IS GIVEN BY
     C PHI = SUM(P(K)*EXP(-P(K+1)*X(I))*COS(P(K+2)*X(I))) WHERE THE
     C SUM IS OVER ALL FUNCTIONS
     C
            [F(J.NE.1) GO TO 20
           XI = X(I)
           DO 10 K=1, NP
           PSX(K)=P(K)*XI
     10
           CONTINUE
           K = 0
           DO 15 L=1, NP, 3
           K = K + 1
           CON=P(L)*EXP(-PSX(L+1))
            SAVC(K)=CON*COS(PSX(L+2))
           SAVS(K)=CON*SIN(PSX(L+2))
     15
           CONTINUE
     20
           ITE=J-1
           K = MOD(ITE, 3) + 1
           L = (J - 1)/3 + 1
           GJ TO (1,2,3),K
           DPHI=SAVC(L)/P(J)
```

- 2 DPHI = XI \* SAVC(L)
- RETURN

  DPHI=-XI\*SAVS(L)

  RETURN

  END



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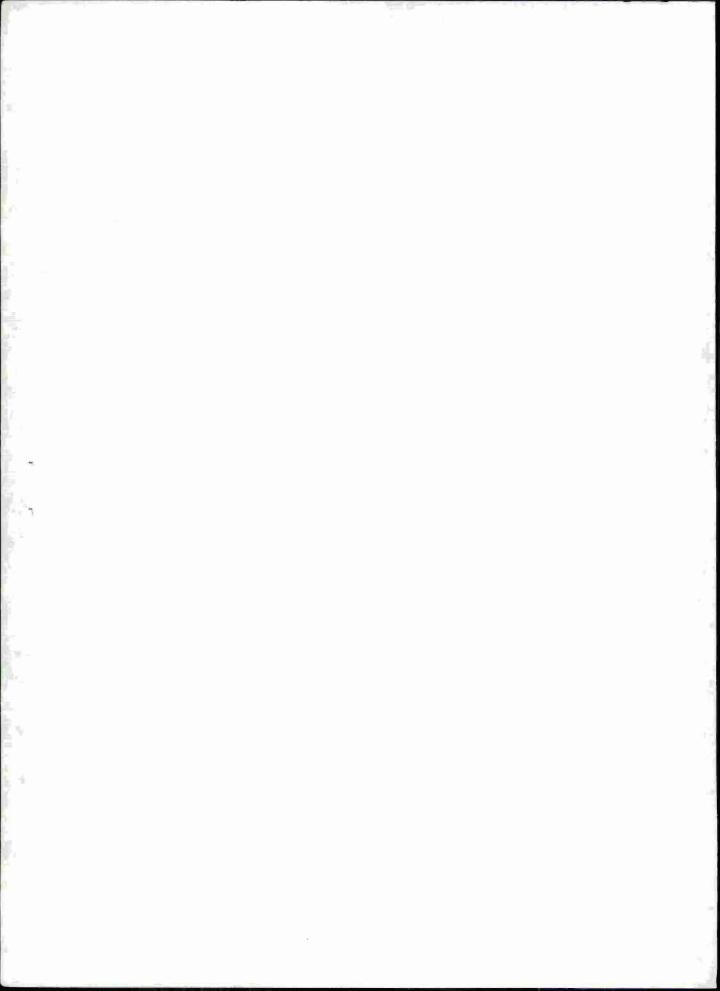
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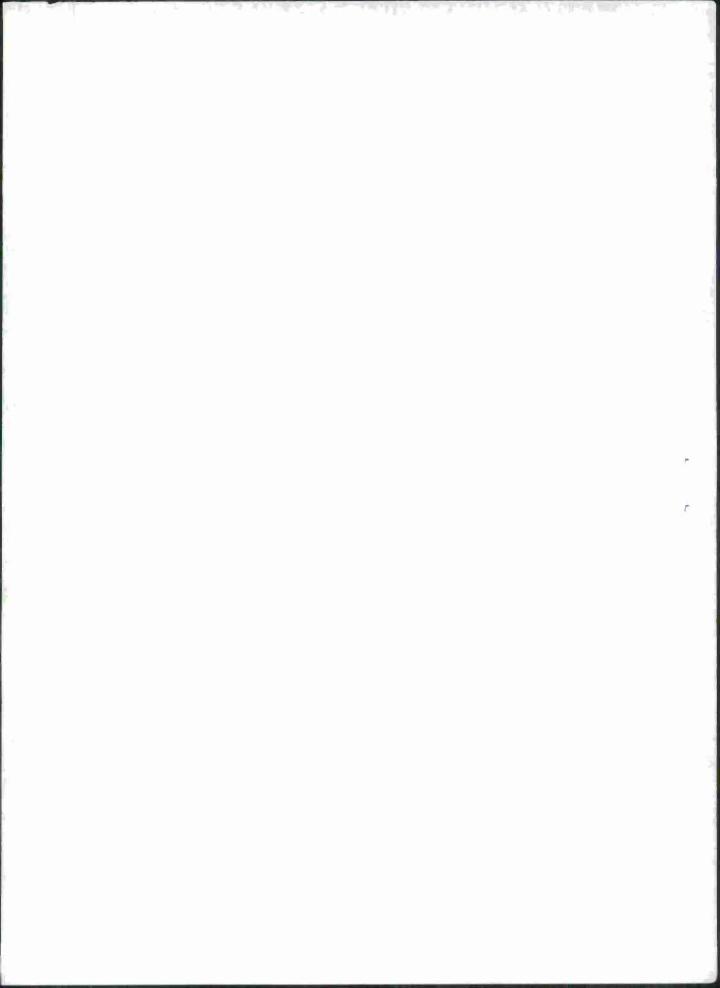
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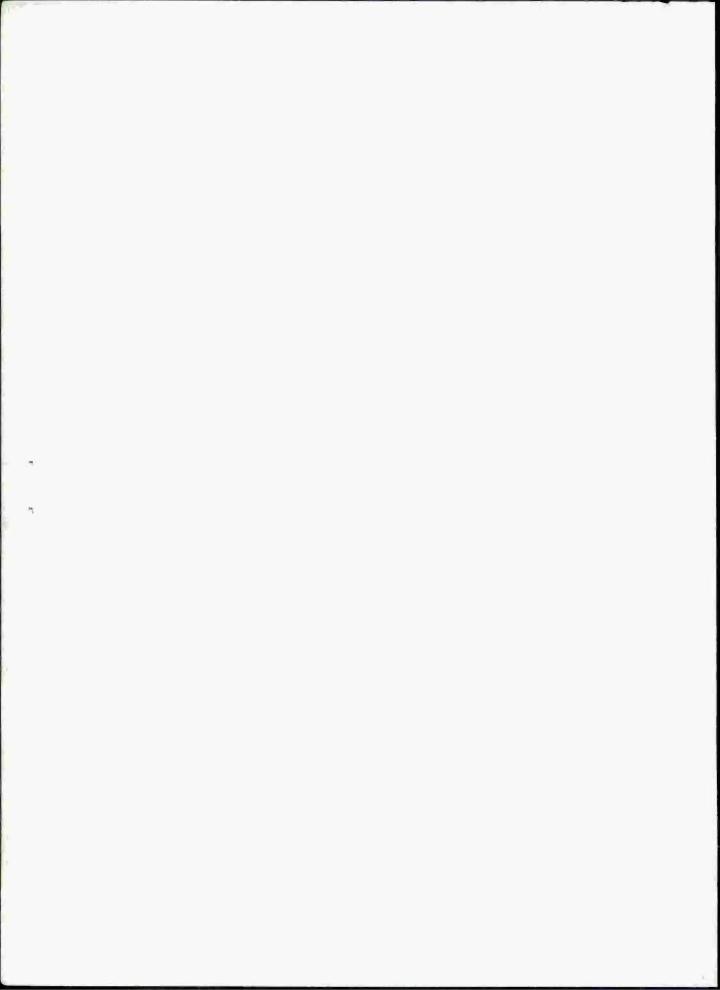
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